# DEUTSCHES ELEKTRONEN-SYNCHROTRON DESY

DESY 87-004 January 1987 87-4-229 高工研図書室

CHIRAL ANOMALY FROM THE FOKKER-PLANCK FORMALISM

by.

J.A. Magpantay, M. Reuter

Deutsches Elektronen-Synchrotron DESY, Hamburg

ISSN 0418-9833

NOTKESTRASSE 85 · 2 HAMBURG 52

DESY behält sie	ch alle Rechte für den Verwertung der in die		· ·	und für die wirtschaftliche mationen vor.
DESY reserves a	all rights for commerc	ial use of i	nformation included	in this report, especially in
	case of filing a	application	n for or grant of pate	nts.

To be sure that your preprints are promptly included in the HIGH ENERGY PHYSICS INDEX, send them to the following address ( if possible by air mail ):

DESY Bibliothek Notkestrasse 85 2 Hamburg 52 Germany DESY 87-004 January 1987

ISSN 0418-9833

Chiral Anomaly from the Fokker-Planck Formalism

J.A. Magpantay <sup>†</sup> and M. Reuter

Deutsches Elektronen-Synchrotron DESY, Hamburg

#### Abstract

We show how chiral anomalies arise in the Fokker-Planck formulation of stochastic quantization. Starting from a noise correlation function which is non-local in real space-time, a gauge invariantly regularized Fokker-Planck Hamiltonian is derived and used to compute the anomaly.

## I. Introduction

Recently various authors /1/ treated the problem of anomalous symmetry breaking in the framework of stochastic quantization /2,3,4,5/. All these derivations of the chiral anomaly, for instance, have been done using the Langevin formulation. In this letter we look at the same problem from the Fokker-Planck point of view. At first glance, it might seem that this can be done trivially because the relationship between the Fokker-Planck and the Langevin formulation is wellestablished. However, actually this is not the case. The reason is that, to regulate the quantum field theory, one uses the Breit-Gupta-Zaks /3/ regularized noise which basically makes the process non-Markov. Thus, it is not evident if there is a Fokker-Planck formulation at all because the equivalence between the two formulations is defined by a single-"time" equation /4/. To circumvent this problem we propose to use the following different regularization scheme: instead of smearing out the noise correlation in the ? -direction, we replace the spacetime  $\delta$  -function by a smooth, Lorentz-invariant regulator function. This leads to a Fokker-Planck Hamiltonian which is a non-local, second order functional operator. Calculating the anomaly in this scheme turns out to be basically equivalent to the well-known point-splitting method /6,7/.

### II. The Gauge-Invariant, Regularized Fokker-Planck Hamiltonian

In this section we will derive the gauge-invariant, regularized Fokker-Planck Hamiltonian using the canonical procedure.

For illustrative purposes consider the Euclideanized action of massless  $\mathsf{QED}_\mathtt{A}$ 

<sup>&</sup>lt;sup>†</sup> Alexander von Humboldt fellow. On leave of absence from the University of the Philippines.

$$S = -\int d^4x \ \overline{+} (i\mathcal{B}) \ 4 \ , \tag{1}$$

where  $\mathcal{J}=\mathcal{J}+i\mathcal{A}$  and we follow the convention  $g_{\mu\nu}=-\delta_{\mu\nu}$ ,  $\gamma_{\mu}^{+}=-\gamma_{\mu}$ . The Langevin equations are then

$$\frac{\partial \Psi}{\partial \tau} = -\mathcal{P}^2 \Psi + i \mathcal{P} \eta \qquad (2a)$$

$$\frac{34}{23} = -4\pi^2 + \eta$$
 (2b)

where  $\overline{\mathcal{H}}=\overline{\mathcal{H}}-\lambda\mathcal{H}$ . The Langevin equations (2a,b) are gauge invariant provided the noise terms  $\gamma$  and  $\overline{\gamma}$  transform like  $\Psi$  and  $\overline{\Psi}$ , respectively.

At this point, we specify how the noise is regularized. It is a common practice to regularize the noise in the  $\Upsilon$  direction. As pointed out already in the introduction, this makes the stochastic process non-Markov. As an alternative, we propose to regulate the noise in the  $\chi_{\mu}$  direction. Because of the gauge transformation properties of  $\gamma$  and  $\overline{\gamma}$ , we propose the following correlation

$$\langle \eta_{\alpha}(q_{3}) \overline{\eta}_{\beta}(x',s') \rangle = \lambda \delta_{\alpha\beta} \delta(s-s') \beta_{\beta}(x-x') \phi(x,x'),$$
 (3)

where  $\beta_{\Lambda}(x'x')$  is a Lorentz-invariant smearing function which approaches the Dirac delta as  $\Lambda$  (which has the dimensionality of a mass) approaches  $\infty$ . We also assume that  $\int \beta_{\Lambda}(x-x') \ d^{\frac{n}{4}}x' = 1$ . The phase factor

$$\Phi(x,x') = e^{i\int_{x}^{x'} A \cdot dy}$$
(4)

is necessary because of the gauge transformation properties of the noises. The correlation (3) implies that the noise have the distribution

To derive the Fokker-Planck Hamiltonian, we consider the Wiener integral corresponding to the Langevin process given by (2a,b).

$$\langle f_{j} T = T_{j} | i_{j} T = 0 \rangle = \int [d \overline{g} (y_{j} \xi)] [d \eta (y_{j} y_{j})] e^{-\frac{1}{2} \int d \tau d x_{j} d x_{j} \tau_{j} (y_{j} \tau_{j}) \beta_{j} (x_{j} - x_{j})} \phi(x_{j} x_{j}) \eta(x_{j} x_{j})} + \int d \tau d x_{j} d x$$

Using the Langevin equations, we transform this into integrals over 4 and  $\bar{4}$ .

$$\langle f, s = \overline{y} | i, t = \sigma \rangle = \int \frac{\left[ d\overline{y}_{i,s} \right] \left[ d\psi(x,s) \right]}{\left[ d\psi(x,s) \right] \left[ \left( \frac{\eta_{i}}{\sqrt{T}} \right) \exp \left( \frac{1}{2} - \frac{1}{R} \right) d\tau d^{2}x d^{2}x^{2}} \right]}$$

$$= \exp \left[ \frac{1}{2} + \overline{\partial}^{2} \right] \beta_{\Lambda}(x - x^{2}) \left( \phi(x, x^{2}) \left( \frac{1}{2} + \overline{\partial}^{2} \right) + \overline{\partial}^{2} \right) + \left[ \frac{1}{2} \left( \frac{3}{2} + \overline{\partial}^{2} \right) + \overline{\partial}^{2} \right] \right]$$

$$= \exp \left[ \frac{1}{2} + \overline{\partial}^{2} \right] \beta_{\Lambda}(x - x^{2}) \left( \phi(x, x^{2}) \left( \frac{1}{2} + \overline{\partial}^{2} \right) + \overline{\partial}^{2} \right) + \left[ \frac{1}{2} + \overline{\partial}^{2} \right]$$

$$= \exp \left[ \frac{1}{2} + \overline{\partial}^{2} \right] \beta_{\Lambda}(x - x^{2}) \left( \phi(x, x^{2}) \left( \frac{1}{2} + \overline{\partial}^{2} \right) + \overline{\partial}^{2} \right) + \left[ \frac{1}{2} + \overline{\partial}^{2} \right] \beta_{\Lambda}(x - x^{2}) \left( \phi(x, x^{2}) \left( \frac{1}{2} + \overline{\partial}^{2} \right) + \overline{\partial}^{2} \right) \right]$$

$$= \exp \left[ \frac{1}{2} + \overline{\partial}^{2} \right] \beta_{\Lambda}(x - x^{2}) \left( \frac{1}{2} + \overline{\partial}^{2} \right)$$

$$= \exp \left[ \frac{1}{2} + \overline{\partial}^{2} \right] \beta_{\Lambda}(x - x^{2}) \left( \frac{1}{2} + \overline{\partial}^{2} \right) + \overline{\partial}^{2} \right]$$

$$= \exp \left[ \frac{1}{2} + \overline{\partial}^{2} \right] \beta_{\Lambda}(x - x^{2}) \left( \frac{1}{2} + \overline{\partial}^{2} \right)$$

$$= \exp \left[ \frac{1}{2} + \overline{\partial}^{2} \right] \beta_{\Lambda}(x - x^{2}) \left( \frac{1}{2} + \overline{\partial}^{2} \right)$$

$$= \exp \left[ \frac{1}{2} + \overline{\partial}^{2} \right] \beta_{\Lambda}(x - x^{2}) \left( \frac{1}{2} + \overline{\partial}^{2} \right)$$

$$= \exp \left[ \frac{1}{2} + \overline{\partial}^{2} \right] \beta_{\Lambda}(x - x^{2}) \left( \frac{1}{2} + \overline{\partial}^{2} \right)$$

$$= \exp \left[ \frac{1}{2} + \overline{\partial}^{2} \right] \beta_{\Lambda}(x - x^{2}) \left( \frac{1}{2} + \overline{\partial}^{2} \right)$$

$$= \exp \left[ \frac{1}{2} + \overline{\partial}^{2} \right] \beta_{\Lambda}(x - x^{2}) \left( \frac{1}{2} + \overline{\partial}^{2} \right)$$

$$= \exp \left[ \frac{1}{2} + \overline{\partial}^{2} \right] \beta_{\Lambda}(x - x^{2}) \left( \frac{1}{2} + \overline{\partial}^{2} \right)$$

$$= \exp \left[ \frac{1}{2} + \overline{\partial}^{2} \right] \beta_{\Lambda}(x - x^{2}) \left( \frac{1}{2} + \overline{\partial}^{2} \right)$$

$$= \exp \left[ \frac{1}{2} + \overline{\partial}^{2} \right] \beta_{\Lambda}(x - x^{2}) \left( \frac{1}{2} + \overline{\partial}^{2} \right)$$

$$= \exp \left[ \frac{1}{2} + \overline{\partial}^{2} \right] \beta_{\Lambda}(x - x^{2}) \left( \frac{1}{2} + \overline{\partial}^{2} \right)$$

$$= \exp \left[ \frac{1}{2} + \overline{\partial}^{2} \right] \beta_{\Lambda}(x - x^{2}) \left( \frac{1}{2} + \overline{\partial}^{2} \right)$$

$$= \exp \left[ \frac{1}{2} + \overline{\partial}^{2} \right] \beta_{\Lambda}(x - x^{2}) \left( \frac{1}{2} + \overline{\partial}^{2} \right)$$

$$= \exp \left[ \frac{1}{2} + \overline{\partial}^{2} \right]$$

$$= \exp \left[ \frac{1}{2} + \overline{\partial}^{$$

and the Jacobian of transformation is

$$J = dit \left[ \frac{3t + 2x^{2}}{3t + 2x^{2}} \right) \delta(t-t^{2})$$

$$\int dt \left[ \frac{3t}{3t + 2x^{2}} \right) \delta(t-t^{2})$$

$$\int dt \left[ \frac{3t}{3t + 2x^{2}} \right) \delta(t-t^{2})$$

$$\int dt \left[ \frac{3t}{3t + 2x^{2}} \right) \delta(t-t^{2})$$

In (7), g(x-x') is the Greens function of  $\lambda \mathcal{N}$  . Factoring out

$$J_0 = \det \begin{bmatrix} \frac{3(x-x')\frac{2}{25}}{\frac{2}{5}} \delta(5-5') & 0 \\ 0 & \frac{2}{21} \delta(x-x')\delta(5-5') \end{bmatrix}$$

and using the midpoint rule  $\,\Theta(\mathfrak{d})\,\mathfrak{z}\,\,\mathcal{V}_{\!\!\boldsymbol{\lambda}}$  , we find

From (8) and the exponential term in (6) we read-off the Fokker-Planck Lagrangian /5/

$$L_{FP} = \frac{1}{2} \int d^{3}x \, d^{3}x' \left[ \frac{\partial Y}{\partial Y} + \overline{Y} + \overline{P}^{2} \right]_{X} \beta_{2}(x-x') \phi(y,x') \left[ \frac{1}{2} \overline{P} \left( \frac{\partial Y}{\partial Y} + \overline{P}^{2} Y \right) \right]_{X}$$

$$- \frac{1}{2} \delta^{4}(0) \int d^{3}x \left[ \overline{Q}^{2} + \overline{P}^{2}_{X} \right]. \tag{9}$$

We now use the canonical procedure. The conjugate momenta are

$$\Pi_{\Psi} = \frac{S L_{EP}}{S(2, \Psi)} = \frac{1}{2} \int d^{3}x d^{3}x_{2} \left[ \frac{\partial \Psi}{\partial T} + \overline{\Psi} \hat{F}^{3} \right]_{X_{1}} \beta_{\Lambda}(x_{1}, x_{2}) \phi(x_{1}, x_{2}) g(x_{2} - x), \quad (10a)$$

$$\Pi_{\vec{q}} = \frac{\delta L_{\text{FP}}}{\delta (\partial_{\vec{q}} \vec{q})} = \frac{1}{a} \int d^{4}x_{1} d^{4}x_{2} \beta_{\Lambda}(x-x_{1}) \phi(x_{1}x_{2}) \frac{\partial (x_{1}-x_{2})}{\partial x_{1}} \frac{\partial (x_{1}-x_{2})}{\partial x_{2}} \frac{\partial$$

To solve for the velocities, we use

$$\int d^{1}x_{1} \, \beta_{\Lambda}(x-x_{1}) \, \phi(x_{1}x_{1}) \, \beta_{\Lambda}^{-1}(x_{1}-y_{2}) \, \phi(x_{1},x_{2}) \; = \; S^{4}(x-x_{2}) \; .$$

The "Hamiltonian" is given by the Legendre transformation

$$H_{FP} = \int d^4x \left[ \Pi_{\Psi_n} \frac{\partial \Psi_n}{\partial r} + \overline{\Pi}_{\overline{\Psi}_n} \frac{\partial \overline{\Psi}_n}{\partial r} \right] - L_{FP}$$
(11)

and after operator ordering gives

$$\hat{H}_{FP} = \int d^{4}x \, d^{4}x' \left\{ \frac{S}{S + ix} \right\} \, i \mathcal{P}_{x} \left( \beta_{n}^{-1}(x - x') \phi(x, x') \right) \frac{S}{S + ix'}$$

$$- \frac{S}{S + ix} \left( \phi^{-1}(x, x') \beta_{n}^{-1}(x - x') \right) i \hat{\mathcal{P}}_{x} \, \frac{S}{S + ix'}$$

$$+ \int d^{4}x \left\{ \frac{S}{S + ix} \right\} \mathcal{P}_{x}^{2} \, \mathcal{H}(x) + \frac{S}{S + ix'} \left( \hat{\mathcal{F}}(x) \hat{\mathcal{F}}_{x}^{A} \right) \right\}. \tag{12}$$

Equation (12) is the gauge-invariant, regularized Fokker-Planck Hamiltonian. By doing a similarity transformation

$$\hat{H}_{FP} = \chi^{\hat{A}} \hat{H}_{FP} L^{\hat{A}}$$
(13a)

$$\hat{A} = \int d^{\prime}x \, d^{\prime}x' \frac{\delta}{\delta \psi(x)} \, \vartheta(x-x') \, \beta_{\Lambda}^{-1}(x-x') \, \hat{\varphi}(x,x') \, \frac{\delta}{\delta \psi(x')} \qquad (13b)$$

we find

$$\hat{H}_{FP} = \int d^4x \left\{ \frac{5}{5 \Psi x} \mathcal{B}_x^2 \Psi x \right\} + \frac{5}{5 \overline{\Psi}_{\alpha}} (\overline{\Psi}_{\alpha}, \widehat{\mathcal{B}}_x^2) \right\}. \tag{14}$$

Note that  $\hat{H}_{FP}^{I}$  does not depend on the regulator and thus its spectrum is independent of  $\Lambda$ . Sakita had shown that  $\hat{H}_{FP}^{I}$  is a positive-definite operator and thus it is also true that the point-splitted, gauge-invariant Fokker-Planck Hamiltonian (12) is also positive definite. Another prerequisite is the existence of a mass gap between the zero mode and the non-zero modes. Making the usual assumption that  $\hat{H}_{FP}^{I}$  has a mass gap then  $\hat{H}_{FP}$  also has a mass gap.

# III. Derivation of the anomaly

We determine the anomaly as the Jacobian  $J[\alpha]$  associated with an infinitesimal chiral transformation

$$\psi(x) = e^{i\kappa(x)\delta_5} \psi(x) ,$$

$$\psi'(x) = \overline{\psi}(x) e^{i\alpha(x)Y_5} \tag{15}$$

For this Jacobian we make the ansatz

$$J[\alpha] = exp.(-i\int d^{4}x \; \alpha(x) \; d_{1}(x)). \tag{16}$$

Now one considers the normalization condition

$$\int [d\Psi] [d\overline{\Psi}] P[\Psi, \overline{\Psi}; \tau] = 1 , \qquad (17)$$

and changes the integration variables from  $\psi$  and  $\bar{\psi}$  to  $\psi'$  and  $\bar{\psi}'$  according to (15) for infinitesimal  $\chi(x)$ . This implies

where  $\delta P(\psi,\tau) = P(\psi,\tau';\tau) - P(\psi,\tau';\tau)$ . To make this equation more transparent, we introduce a complete set  $\{F_n(\psi,\psi)\}$  of eigenfunctionals of the Fokker-Planck Hamiltonian:  $\hat{H}_{FP}$   $F_n = E_n$   $F_n$ . In terms of the  $F_n$ 's the distribution function may be expanded as

$$P[4,\bar{4},r] = \sum_{n}^{\infty} C_n e^{-E_n t} F_n[4,4]$$
 (19)

The coefficients  $\left\{ c_{n} \right\}$  are to be determined from the initial conditions. Thus we have

$$i\int d^3x \, \alpha(x) \, d(x) = \sum_{n} c_n e^{-E_n T} \int [d+7[d\bar{\psi}] \, \delta F_n \, [\psi, \bar{\psi}] \, . \tag{20}$$

This equation nicely exhibits the  $\Upsilon$  -evolution of the anomaly. Since in general the eigenfunctionals  $F_n$  for  $E_n \nearrow 0$  are not known explicitly, we can evaluate the RHS of (20) only in the limit  $\Upsilon \rightarrow \infty$ . In this case only the zero-mode  $F_0 \equiv Z^{-1} \exp(-S_{\xi}) \text{ with the partition function } \vec{Z} \equiv \int \!\!\! \int \!\!\! d\psi \, \mathcal{V} \, d\vec{\Psi} \, \mathcal{V} \exp(-S_{\xi})$  contributes. Hence one obtains

$$i\int d^{4}x \, \alpha(x) \, dx \, |x| = - \frac{7}{2} \int [a47] \, d^{2}7 \, \delta S_{\epsilon} \, exp. \left(-S_{\epsilon}\right) = -\left(\delta S_{\epsilon}\right).$$
 (21)

Contrary to Fujikawa's /8/ approach, where A is evaluated directly from the definition of the Jacobian  $\delta (\Psi, \bar{\Psi})/\delta (\Psi, \bar{\Psi})$ , we here determine A by calculating  $\langle \delta S_{\epsilon} \rangle$ .

From the Hamiltonian (12) it is easy to see that  $\mathcal{S}_{\mathbf{c}}$  is given by

Recalling that for  $\Lambda \to \infty$ , the Lorentz invariant function  $\beta_{\Lambda}$  approaches a  $\delta$ -function, we may replace  $\lim_{\Lambda \to \infty} \int_{0}^{d\xi} \beta_{\Lambda}^{(\xi)}(\dots)$  by  $\lim_{\xi \to 0} \dots$  where an average over the directions of  $\xi^{n}$  according to  $\frac{\xi^{n}\xi^{n}}{\xi^{n}} \to \frac{\xi^{n}}{\xi^{n}}$ , etc. is understood. Expanding the bracket of (22) to first order in  $\xi$ , which is equivalent to the first order in  $1/\Lambda$ , one obtains

$$S_{\xi} = -\lim_{\xi \to 0} \int d^{4}x \, \bar{\psi}(x + \underline{\xi}) \left\{ i\partial_{\xi} - A + \bar{\xi}^{M} A_{m} \partial_{\xi} + i A_{\xi}^{M} A_{m} + i A_{\xi}^{M} \partial_{\xi} + i A_{\xi}^{M} \partial_{\xi} \partial_{\xi}$$

The change of (23) under an infinitesimal chiral rotation reads

$$SS_{\varepsilon} = -\int d^{4}x \, \overline{\Psi}(x + \frac{\varepsilon}{2}) \left\{ -(\partial \alpha) + \varepsilon^{4}(\Omega_{\alpha})(\partial x + iA) + \frac{1}{2} \partial (\varepsilon^{4}) \right\}$$

$$+ i \, \varepsilon^{4} A_{n}(\partial \alpha) \, d^{4}x \, d^{4}x + \frac{\varepsilon}{2} \, d^{4}x \, d^{4}x + \frac{\varepsilon}{2} \, d^{4}x \, d^{4}x + \frac{\varepsilon}{2} \, d^{4}x + \frac{$$

Next, one has to compute the vacuum expectation value of (24) in the limit  $\xi \to 0$  . Using the well known matrix element /7/

and the Heisenberg equations for  $\Psi$  it is straightforward to arrive at

$$\langle \delta S_{\xi} \rangle = -i \lim_{\xi \to 0} \int d^{4}x \, F \, \left[ \sum_{n} \langle \overline{\Psi}(x + \underline{\xi}) \rangle^{n} \, V_{\delta} + (x - \underline{\xi}) \rangle \right]$$

$$= -i \int d^{4}x \, \alpha(x) \, \mathcal{A}(x) \, , \qquad (26)$$

with

$$A(x) = \frac{1}{511^2} F_{\mu\nu} F^{\mu\nu}. \qquad (27)$$

Together with (16) this yields the standard result /8/ for the Jacobian of the chiral transformations (15):

$$J[x] = exp\left(-\frac{i}{\delta \Pi^{2}}\int d^{4}x F_{\mu\nu}^{*} F^{\mu\nu}\right) \qquad (28)$$

To obtain the divergence of the axial vector current one notes that using the equations of motion for  $\psi$  the expectation value of (24) can be written as

Inserting this together with (27) into (21) one ends up with the desired relation /7/

This completes the proof that the Fokker-Planck dynamics described by our  $\hat{H}_{FP}$  leads to the correct anomaly for  $\mathcal{T} \rightarrow \infty$  .

#### IV. Conclusions

In this letter we proposed a new way of regulating the noise in stochastic quantization. We find that this scheme gives rise to a (gauge-invariant) Fokker-Planck Hamiltonian whereas in the Breit-Gupta-Zaks scheme it is not evident if there is a simple Fokker-Planck formulation. As a first application, we derive the chiral anomaly of QED<sub>4</sub>. The generalization to more complicated theories is in principle straightforward.

### References

- /1/ J. Alfaro, M.B. Gavela, Phys. Lett. 158B (1985) 473;
  M.B. Gavela, N. Parga, Phys. Lett. 174B (1986) 319;
  R. Kirschner, E.R. Nissimov, S.J. Pacheva, Phys. Lett. 174B (1986) 324;
  E.S. Egorian, E.R. Nissimov, S.J. Pacheva, Lett. Math. Phys. 11 (1986) 209;
  E.R. Nissimov, S.J. Pacheva, Lett. Math. Phys. 11 (1986) 373;
  R. Tzani, Phys. Rev. D33 (1986) 1146;
  J.P. Ader, J.C. Wallet, Z. Phys. C32 (1986) 575;
  Z. Bern, H.S. Chan, M.B. Halpern, Z. Phys. C 33 (1986) 77;
  E.R. Nissimov, S.J. Pacheva, Phys. Lett. 171B (1986) 267 and CERN-TH.4435/86, unpublished;
  M. Reuter, DESY-preprint (1986), unpublished
- /2/ G. Parisi, Y. Wu, Sci. Sin. 24 (1981) 483; P.H. Damgaard, K. Tsokos, Nucl. Phys. B235 FS11 (1984) 75
- /3/ J.D. Breit, S. Gupta, A. Zaks, Nucl. Phys. B233 (1984) 61
- /4/ 8. Sakita, Proceedings of the Johns Hopkins Workshop 1983, World Scientific, Singapore
- /5/ E. Gozzi, Phys. Rev. D28 (1983) 1922
- /6/ J. Schwinger, Phys. Rev. 82 (1951) 664
- 77/ R. Jackiw, K. Johnson, Phys. Rev. 182 (1969) 1459;
   R. Jackiw in Lectures on current algebra and its applications; ed.
   D. Gross, R. Jackiw, S. Treiman, Princeton University Press (1972)
- /8/ K. Fujikawa, Phys. Rev. Lett. 42 (1979) 1195 and Phys. Rev. D21 (1980) 2848