

Tackling light higgsinos at the ILC

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In supersymmetric extensions of the Standard Model, higgsino-like charginos and neutralinos are preferred to have masses of the order of the electroweak scale by naturalness arguments. Such light $\tilde{\chi}_1^0$, $\tilde{\chi}_2^0$ and $\tilde{\chi}_1^\pm$ states can be almost mass degenerate, and their decays are then difficult to observe at colliders. In addition to the generic naturalness argument, light higgsinos are well motivated from a top-down perspective. For instance, they arise naturally in certain models of hybrid gauge-gravity mediation. In the present analysis, we study two benchmark points which have been derived in the framework of such a model, which exhibit mass differences of $\mathcal{O}(\text{GeV})$ in the higgsino sector. For chargino pair and neutralino associated production with initial-state photon radiation, we simulate the detector response and determine how accurately the small mass differences, the absolute masses and the cross sections can be measured at the International Linear Collider. Assuming that 500 fb^{-1} has been collected at each of two beam-polarisations $P(e^+, e^-) = (\pm 30\%, \mp 80\%)$, we find that the mass-differences can be measured to 40–300 MeV, the cross sections to 2–5%, and the absolute masses to 1.5–3.3 GeV, where the range of values correspond to the different scenarios and channels. Based on these observables we perform a parameter fit in the MSSM, from which we infer that the higgsino mass parameter μ can be measured to a precision of about $\Delta\mu = 2\text{--}7 \text{ GeV}$. For the electroweak gaugino mass parameters M_1, M_2 , which are chosen in the multi-TeV range, a narrow region is compatible with the measurements. For both parameters independently, we can determine a lower bound.

1 Introduction

The LHC experiments have discovered a Standard Model (SM)-like Higgs boson with $m_h \approx 126 \text{ GeV}$ [1, 2]. However, it is not clear whether this newly discovered particle is a pure SM Higgs boson, or whether it is part of an extension of the SM, which might also include multiple or composite Higgs-like states. Precise measurements of the Higgs couplings and branching ratios will be required to resolve this question. Another, more direct way to gain insights about physics beyond the SM (BSM) would obviously be to discover more new particles directly.

A well-motivated extension of the SM is the Minimal Supersymmetric SM (MSSM), a fully renormalisable theory of high predictive power. In this model, the U(1), SU(2) and SU(3) gauge couplings can unify, and a dark matter candidate can be naturally accommodated. Furthermore, supersymmetry stabilizes the Higgs potential with respect to quadratically divergent radiative corrections. In the MSSM the Higgs mass is no longer a free parameter but becomes strongly constrained.

If nature is supersymmetric at the electroweak scale, a large number of new superpartner mass parameters is introduced whose values depend on the unknown mechanism of supersymmetry breaking at high energies. This poses a challenge for any phenomenological analysis. Due to the large number of unknown parameters it is very difficult to make reliable predictions for SUSY background processes, and in many cases the worst background contributions for SUSY signals originate in fact from SUSY processes.

The experiments at the LHC have not yet found any hints for supersymmetry in squark and gluino searches [3, 4], which points towards a rather heavy coloured superpartner spectrum with soft masses above

a TeV. However, there are large regions in the MSSM parameter space which lead to a SM-like Higgs with $m_h = 126$ GeV, a light electroweak SUSY sector, and a rather heavy coloured SUSY sector beyond the current LHC discovery range. In general, if the electroweak sector is light enough, it should be possible to discover neutralinos and charginos via Drell-Yan production at the LHC [5]. However, in cases where these light states are very close in mass and their decay products are therefore very soft, it will be almost impossible to observe such processes at the LHC. An example is given by scenarios in which the higgsino mass parameter μ is much smaller than the electroweak gaugino masses M_1 and M_2 (current bounds require only $|\mu| \gtrsim 100$ GeV in order to avoid the limits from LEP chargino searches [6]). Then, the new light states are a chargino and two neutralinos which are higgsino-like and almost mass-degenerate. The chargino and the heavier neutralino therefore decay only into extremely soft SM particles and the invisible lightest neutralino.

The most “natural” versions of the phenomenological MSSM which are still not ruled out are indeed characterized by higgsino masses of a few hundred GeV, together with sub-TeV third-generation squarks and a moderately heavy gluino, while the rest of the spectrum can be in the multi-TeV range [7]. Models with light higgsinos are also attractive for cosmology, either in combination with gravitino dark matter or with a higgsino-like neutralino as a non-thermally produced lightest supersymmetric particle (LSP); see e.g. [8].

While the light higgsino scenario is extremely challenging for the LHC, it may well be observable in the clean environment at an e^+e^- Linear Collider, for instance, at the International Linear Collider (ILC) [9, 10]. At the ILC, the striking feature of a precisely known initial state allows to successfully apply the initial state radiation (ISR) method, i.e. to select only those signal events that are accompanied by a sufficiently hard photon. In this context, polarised beams help to significantly enhance the signal cross section, at least in the chargino case.

The subject of the present work is the phenomenology of light quasi-degenerate higgsinos at the ILC. We study their production and decay, and investigate in how far one can derive the fundamental MSSM parameters without assuming any specific SUSY breaking scheme if only the light higgsinos are kinematically accessible. We study the processes $e^+e^- \rightarrow \tilde{\chi}_1^+\tilde{\chi}_1^-$ and $e^+e^- \rightarrow \tilde{\chi}_1^0\tilde{\chi}_2^0$ at $\sqrt{s} = 500$ GeV with polarised beams and initial-state radiation at the ILC. We simulate the detector response for two MSSM benchmark points with different $\mathcal{O}(1$ GeV) mass splittings between the chargino and neutralinos, leading to different final state SM particles. The uncertainties on masses, mass differences and cross sections as determined by our simulation study are then used as inputs for a fit in order to reconstruct the MSSM parameters M_1 , M_2 , μ and $\tan\beta$. Our overall goal is to determine both how well the chargino and neutralino masses can be measured at the ILC, and what could be inferred from such a measurement about the underlying SUSY model.

In the following chapters we first give an overview of the theoretical arguments in favour of light higgsinos and the phenomenological aspects of such scenarios in the MSSM. The expected SM background will also be discussed, including beam-conditions and detector capabilities at the ILC. In Section 3 we define two benchmark scenarios on which our study will be based and provide the details for the detector simulation. In Section 4 we discuss the expected experimental results at the ILC. In Section 5 we present our results on reconstructing the fundamental parameters and the precision we expect to achieve with our analysis strategy.

2 Theoretical context and analysis strategy

There are two main reasons why higgsinos are distinguished among the MSSM superpartners of the SM particles. First, the size of the higgsino mass parameter μ is not necessarily related to the scale of SUSY breaking, as it is allowed by unbroken supersymmetry (it is indeed the only dimensionful MSSM parameter with this property). It can therefore be generated independently of the soft supersymmetry breaking terms, and at leading order does not receive any soft term-dependent radiative corrections in the evolution from the high scale. Second, μ enters the tree-level Higgs potential and is therefore directly connected to the electroweak scale. For instance, the tree-level Z boson mass is given in terms of the Higgs soft masses m_{H_u, H_d}^2 , the ratio $\tan\beta$ of Higgs vacuum expectation values, and μ as

$$m_Z^2 = 2 \frac{m_{H_u}^2 \tan^2 \beta - m_{H_d}^2}{1 - \tan^2 \beta} - 2 |\mu|^2 \approx -2 (m_{H_u}^2 + |\mu|^2). \quad (1)$$

To avoid finely tuned large cancellations among these parameters, it follows that μ should not be too far above the electroweak scale, even if some or all of the squark, slepton and gaugino masses are much larger. These two arguments are the main motivation for investigating models where the higgsinos are light, while the other MSSM superpartners may be much heavier.

2.1 Light higgsinos from UV-scale models

From a top-down perspective, light higgsinos are equally well motivated. For instance, they are generic in certain models of hybrid gauge-gravity mediation [11, 12], whose main features we will now briefly recapitulate. In gauge-mediated SUSY breaking, some of the hidden sector states carry SM gauge charges. These messenger fields acquire large supersymmetric masses M , and messenger loops generate soft terms of the order

$$m_{\text{soft}} \sim \frac{n}{16\pi^2} \frac{F}{M}, \quad (2)$$

where \sqrt{F} is the scale of SUSY breaking, and n roughly counts the number of contributing messenger states. In addition to these, generically there are also gravity-mediated contributions to soft terms as well as a gravity-mediated effective μ parameter, generated by gravitationally suppressed interactions between the hidden sector and the MSSM. They are of the order of the gravitino mass,

$$m_{3/2} \sim \mu \sim \frac{F}{M_{\text{Planck}}}. \quad (3)$$

In models where the MSSM is embedded in a more fundamental theory, for instance in superstring compactifications, the details of supersymmetry breaking are often unknown, while the spectrum of sub-Planckian states can still be calculated. In phenomenologically promising models these typically consist of the MSSM along with a multitude of additional exotic states, some of which are charged under the SM gauge groups. These exotic fields can be given large masses and act as gauge mediation messengers. A natural scale for them to decouple is the scale of the Grand Unification,

$$M \sim M_{\text{GUT}} = 2 \cdot 10^{16} \text{ GeV} \approx \frac{1}{16\pi^2} M_{\text{Planck}}. \quad (4)$$

With some simple assumptions about the couplings of the messengers to the SUSY-breaking fields, and choosing parameters such that $m_{3/2}$ is of the order of the electroweak scale, one can obtain a realistic soft term spectrum. For large messenger multiplicities $n \gtrsim 10$, the soft terms are predominantly generated by gauge mediation and can be in the multi-TeV range, while μ is induced by gravity mediation and of the order of 100–200 GeV.

The benchmark points we are using for the present analysis are derived from such hybrid gauge-gravity mediation models with large messenger numbers. However, we stress that these are by no means the only possibility to obtain light higgsinos from high-scale models, and that our results on the light higgsino phenomenology are independent of a specific SUSY breaking model.

2.2 Phenomenology

In a more general context, scenarios with a rather heavy coloured sector and a mass-degenerate light chargino/neutralino sector can easily be constructed in the unconstrained MSSM, where the soft SUSY-breaking parameters are taken as independent. One simply chooses sufficiently large values for the squark and gluino masses (and possibly also for the slepton masses).

At the tree level the chargino/higgsino sector [13] is determined by the U(1) bino mass parameter M_1 , the SU(2) wino mass parameter M_2 , the higgsino mass parameter μ and $\tan\beta$. The neutralino mass matrix in the $(\tilde{B}, \tilde{W}^0, \tilde{H}_d^0, \tilde{H}_u^0)$ basis is given by

$$Y = \begin{pmatrix} M_1 & 0 & -m_Z \cos\beta \sin\theta_W & m_Z \sin\beta \sin\theta_W \\ 0 & M_2 & m_Z \cos\beta \cos\theta_W & -m_Z \sin\beta \cos\theta_W \\ -m_Z \cos\beta \sin\theta_W & m_Z \cos\beta \cos\theta_W & 0 & -\mu \\ m_Z \sin\beta \sin\theta_W & -m_Z \sin\beta \cos\theta_W & -\mu & 0 \end{pmatrix}. \quad (5)$$

Since Y is complex symmetric, it can be diagonalized by one unitary matrix N via $\mathbf{M}_{\tilde{\chi}^0} = N^* Y N^\dagger$ (where the matrix $\mathbf{M}_{\tilde{\chi}^0}$ contains the neutralino masses $M_{\tilde{\chi}_i^0}$ on the diagonal). The mixing matrix N relates the neutralino mass eigenstates to the neutral gauge eigenstates. For full analytical solutions of the system see [14, 15].

The chargino mass matrix in the $(\tilde{W}^+, \tilde{H}^+)$ basis,

$$X = \begin{pmatrix} M_2 & \sqrt{2} m_Z \cos \theta_W \sin \beta \\ \sqrt{2} m_Z \cos \theta_W \cos \beta & \mu \end{pmatrix}, \quad (6)$$

is not symmetric and therefore has to be diagonalised via the bi-unitary transformation $\mathbf{M}_{\tilde{\chi}^\pm} = U^* X V^\dagger$. U and V relate the chargino mass eigenstates to the charged gauge eigenstates.

One can easily classify three limiting cases: in the case $M_1 < M_2 \ll |\mu|$, one obtains $M_1 \approx M_{\tilde{\chi}_1^0}$ with a bino-like $\tilde{\chi}_1^0$, $M_2 \approx M_{\tilde{\chi}_2^0}, M_{\tilde{\chi}_1^\pm}$ with wino-like $\tilde{\chi}_2^0$ and $\tilde{\chi}_1^\pm$, and higgsino-like heaviest states with masses $|\mu| \approx M_{\tilde{\chi}_{3,4}^0}, M_{\tilde{\chi}_2^\pm}$, cf. Eqs. (5) and (6). By contrast, the hierarchy $M_2 < M_1 \ll |\mu|$ leads to the lightest two states being mass-degenerate and wino-like, $M_2 \approx M_{\tilde{\chi}_1^0}, M_{\tilde{\chi}_1^\pm}$ but a bino-like $\tilde{\chi}_2^0$ with mass $M_1 \approx M_{\tilde{\chi}_2^0}$, while $\tilde{\chi}_{3,4}^0$ and $\tilde{\chi}_2^\pm$ remain higgsino-like and heavy as before.

Choosing $|\mu| \ll M_1 < M_2$ (or $|\mu| \ll M_2 < M_1$) leads to mass-degenerate lightest states with $|\mu| \approx M_{\tilde{\chi}_{1,2}^0}, M_{\tilde{\chi}_1^\pm}$ but $M_1 \approx M_{\tilde{\chi}_3^0}$ (or $M_1 \approx M_{\tilde{\chi}_4^0}$) and $M_2 \approx M_{\tilde{\chi}_2^\pm}, M_{\tilde{\chi}_4^0}$ (or $M_2 \approx M_{\tilde{\chi}_2^\pm}, M_{\tilde{\chi}_3^0}$). The lighter neutralinos and the lightest chargino are then mostly higgsino-like and have unsuppressed couplings to the SM gauge bosons. This is the scenario which we are investigating.

It is instructive to explicitly write down the tree-level masses of the three light states in the case that M_1 and M_2 are large. They are found by diagonalizing the mass matrices of Eqs. (5) and (6), and given by

$$\begin{aligned} M_{\tilde{\chi}_{1,2}^0} &= |\mu| \mp \frac{m_Z^2}{2} (1 \pm \sin 2\beta \operatorname{sign}(\mu)) \left(\frac{\sin^2 \theta_W}{M_1} + \frac{\cos^2 \theta_W}{M_2} \right), \\ M_{\tilde{\chi}_1^\pm} &= |\mu| - \sin 2\beta \operatorname{sign}(\mu) \cos^2 \theta_W \frac{m_Z^2}{M_2}, \end{aligned} \quad (7)$$

up to terms suppressed by higher powers of M_i , see e.g. [16]. The wino and bino admixtures are small, of the order $m_Z/M_{1,2}$. For instance, for $\mu > 0$, the (two-component) mass eigenstates are decomposed as

$$\begin{aligned} \tilde{\chi}_1^0 &= \frac{1}{\sqrt{2}} (\tilde{H}_d^0 - \tilde{H}_u^0) + \frac{\sin \beta + \cos \beta}{\sqrt{2}} \frac{m_Z}{M_1} \sin \theta_W \tilde{B} - \frac{\sin \beta + \cos \beta}{\sqrt{2}} \frac{m_Z}{M_2} \cos \theta_W \tilde{W}^0, \\ \tilde{\chi}_2^0 &= \frac{1}{\sqrt{2}} (\tilde{H}_d^0 + \tilde{H}_u^0) - \frac{|\sin \beta - \cos \beta|}{\sqrt{2}} \frac{m_Z}{M_1} \sin \theta_W \tilde{B} + \frac{|\sin \beta - \cos \beta|}{\sqrt{2}} \frac{m_Z}{M_2} \cos \theta_W \tilde{W}^0, \\ \tilde{\chi}_1^+ &= \tilde{H}_u^+ - \sqrt{2} \sin \beta \frac{m_W}{M_2} \tilde{W}^+, \\ \tilde{\chi}_1^- &= \tilde{H}_d^- - \sqrt{2} \cos \beta \frac{m_W}{M_2} \tilde{W}^-, \end{aligned} \quad (8)$$

where we have only written the respective leading-order terms of the gaugino and higgsino components. For multi-TeV gaugino masses, the gaugino admixtures are of the order of a few percent, and the mass splittings are of the order of 1–10 GeV. Explicitly, the tree-level mass difference between the lightest neutralino and the lightest chargino reads, again for $\mu > 0$,

$$M_{\tilde{\chi}_1^\pm} - M_{\tilde{\chi}_1^0} = \frac{m_Z^2}{2} \left[\sin 2\beta \left(\frac{\sin^2 \theta_W}{M_1} - \frac{\cos^2 \theta_W}{M_2} \right) + \left(\frac{\sin^2 \theta_W}{M_1} + \frac{\cos^2 \theta_W}{M_2} \right) + \mathcal{O} \left(\frac{\mu}{M_i^2} \right) \right], \quad (9)$$

which for $\tan \beta \gg 1$ further simplifies to

$$M_{\tilde{\chi}_1^\pm} - M_{\tilde{\chi}_1^0} = \frac{m_Z^2}{2} \left(\frac{\sin^2 \theta_W}{M_1} + \frac{\cos^2 \theta_W}{M_2} \right) + \mathcal{O} \left(\frac{\mu}{M_i^2}, \frac{1}{\tan \beta} \right). \quad (10)$$

For the neutralinos, the mass difference is

$$M_{\tilde{\chi}_2^0} - M_{\tilde{\chi}_1^0} = m_Z^2 \left(\frac{\sin^2 \theta_W}{M_1} + \frac{\cos^2 \theta_W}{M_2} \right) + \mathcal{O} \left(\frac{\mu}{M_i^2} \right). \quad (11)$$

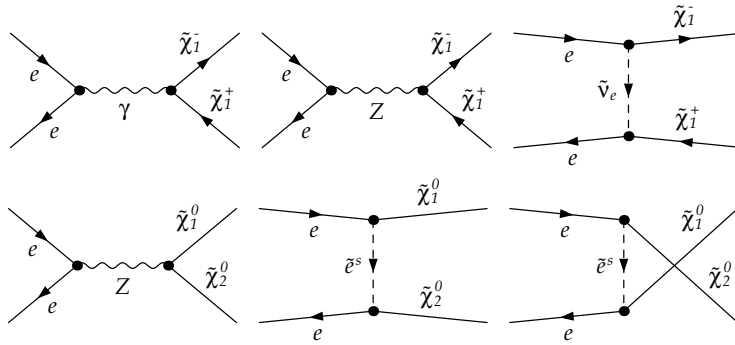


Figure 1: Chargino (upper row) and neutralino (lower row) production graphs at tree level. Note that t - and u -channel selectron/sneutrino exchange graphs are negligible in our scenario due to a small Yukawa coupling between sleptons and higgsinos.

Remarkably, it does not depend on $\tan\beta$ in the leading approximation. The corresponding expressions for $\mu < 0$ are similar and the phenomenology of that parameter range has been discussed in [17].

Comparing with the analytical formulae of the neutralino (chargino) masses and couplings [15, 18], it is obvious that one can express the neutralino and chargino sector as functions of $\sin 2\beta$ and $\cos 2\beta$. In the large $\tan\beta$ limit, $|\cos 2\beta| \rightarrow 1$ and $\sin 2\beta \rightarrow 2/\tan\beta$, therefore all neutralino and chargino masses as well as the cross sections are only weakly dependent on $\tan\beta$. This weak dependence is also reflected in our final results, cf. Section 5.

Radiative corrections to the chargino and neutralino observables, see e.g. [19] (and references therein for a more complete list of studies), can be sizeable and increase the mass splittings by several hundreds of MeV. These corrections will lift the degeneracy between higgsino states. If the tree-level mass difference is substantially less than a GeV, a correction of 200 MeV will have a profound impact on the lifetime and decay branching ratios. This may also be exploited to better determine the underlying MSSM parameters, see [20]. A detailed analysis of NLO corrections in this regard, is, however, beyond the scope of this study and in the following we will assume tree-level relations.

At the ILC higgsino production proceeds via Z or γ exchange in the s -channel, see Fig. 1. In general, t - and u -channel exchange of selectrons or sneutrinos is also possible, but in the higgsino case they are highly suppressed due to the small Yukawa couplings. Hence, even if the selectrons and sneutrinos were relatively light, they would not affect any observables. Higgsinos will be produced in $\tilde{\chi}_1^0\tilde{\chi}_2^0$ or $\tilde{\chi}_1^+\tilde{\chi}_1^-$ pairs (the $Z\tilde{\chi}_1^0\tilde{\chi}_1^0$ and $Z\tilde{\chi}_2^0\tilde{\chi}_2^0$ couplings vanish in the pure higgsino limit).¹ Both $\tilde{\chi}_2^0$ and $\tilde{\chi}_1^\pm$ will undergo three-body decays into $\tilde{\chi}_1^0$ via off-shell Z and W bosons [21]. The decay $\tilde{\chi}_2^0 \rightarrow W^{\pm*}\tilde{\chi}_1^\mp$ is also possible but kinematically highly suppressed. The detector signature comprises exclusively a few, low-energetic leptons or hadrons, whose softness poses a major challenge to experiment. For small mass splittings, the radiative decay $\tilde{\chi}_2^0 \rightarrow \tilde{\chi}_1^0\gamma$ via a W -chargino loop can also become very important [22]. More details on higgsino production and decay will be given below when we discuss the details of our simulation.

2.3 Standard model background

For the discussion of the SM processes at the ILC it is convenient to classify them by the initial state and the number of final state fermions. Due to the very strong fields between the colliding bunches, there is a sizable photon component from back-scattered synchrotron radiation in the ILC beams [23, 24]. Therefore, the initial states can be e^+e^- , but also $\gamma\gamma$, or γe^\pm . For e^+e^- and $\gamma\gamma$, the final state has an even number of fermions: 2f, 4f, 6f, ... , while the γe^\pm initial state gives rise to final states with an odd number of fermions: 1f, 3f, 5f, ... In this notation, γ can represent both real beam-photons or virtual ones. For initial virtual photons, the final state is understood to also contain the electron and/or positron that emitted the photon

¹This is a consequence of the fact that \tilde{H}_u^0 and \tilde{H}_d^0 have opposite weak isospin, and that both $\tilde{\chi}_1^0$ and $\tilde{\chi}_2^0$ contain equally large \tilde{H}_u^0 and \tilde{H}_d^0 admixtures as seen from Eq. (8).

(the “beam-remnant”), in addition to the fermion(s).

The signal comprises events with a few low-energy charged particles and no other activity. Possible SM processes giving this topology are of two types. Either (i) events with large amounts of energy in invisible neutrinos, or (ii) events where a large fraction of the energy escapes with particles going in un-instrumented regions of the detector; in an ILC detector, the only such regions are the holes in the low-angle calorimeter through which the beam pipes enter the detector.

The first background type mainly contains events with two τ s, possibly together with additional neutrinos, i.e. the subset of the classes $e^+e^- \rightarrow 2f$ and $e^+e^- \rightarrow 4f$ where one of the fermion pairs are τ leptons, and the other one (if present) is a neutrino pair. These channels do not have cross sections much larger than the signal. In addition, only a small fraction of the τ s will decay to the very soft visible system which is typical for the signal. Hence, while it is not negligible, this background-type is expected to be manageable without too strong cuts on the event-topology.

The potentially most severe background processes are of type (ii), and are expected to be the high cross section multi-peripheral process $e^+e^- \rightarrow e^+e^-\gamma^*\gamma^* \rightarrow e^+e^-f\bar{f}$, i.e. the $\gamma\gamma \rightarrow 2f$ class. In such events, the final beam-remnant electrons and positrons carry nearly the full beam-energy in most cases, have low transverse momentum, and therefore leave the detector unseen through the out-going beam-pipe. The $f\bar{f}$ -pair, on the other hand, can be emitted at any angle, and tends to have low momentum. In other words, these events have a large resemblance to the signal. However, if one of the incoming electrons emits an ISR photon at an angle and energy high enough to be detected, this electron will recoil against the photon, changing its direction sufficiently to deflect it into the acceptance of the detector. In order to reduce this background to acceptable levels, we therefore require that the accepted signal candidate events should contain a detected ISR photon. Also the other component of the $\gamma\gamma \rightarrow 2f$ class, where the initial photons are real photons from the photon component of the beams, is suppressed by this requirement. This is simply due to the fact that, in this case, there are no initial charged particles that could emit an ISR photon.

Instead, the most severe remaining background is expected to arise from the $e\gamma \rightarrow 3f$ class, in particular from the t -channel diagram with incoming real photons. This process is similar to Compton scattering, with the difference that the scattered photon is virtual, and gives rise to an $f\bar{f}$ -pair. In this case, the hard ISR photon does not recoil against the beam electron, but against the $f\bar{f}$ -system, like in the signal events. However, in the event as a whole, a large amount of energy is carried away by the *exchanged* electron once it has been brought back on-shell by the beam-photon. This electron will receive most of the photon’s energy, and closely follow its direction in the lab frame, i.e. along the beam pipe. As the second incoming particle is neutral in this case, no ISR will be emitted from it, and thus it cannot be deflected into the detector acceptance either.

3 Benchmark points and simulation

For our simulation we chose two benchmark points with the lightest Higgs mass which is compatible with the LHC discovery (within the theory uncertainty), namely 124 and 127 GeV. We will refer to them as dM1600 and dM770, according to their chargino-LSP mass difference. In the following sections we will discuss their low-scale spectrum and the higgsino production and decay modes at the ILC as well as their implementation into our simulation. We especially comment on the decay modes and branching ratios in the nearly mass-degenerate $\tilde{\chi}_1^\pm, \tilde{\chi}_{1,2}^0$ system.

3.1 Spectrum generation, higgsino production and decay

Spectrum generation

The general tree-level spectrum of our model was calculated using the spectrum generator **SOFTSUSY** [25]. The model we are considering is characterized by the high scale parameters specified in Table 1. The parameters in the gaugino, squark, and slepton sector at the high scale are obtained from Eqs. (3) and (4) in [12], while the Higgs sector parameters are computed such that both the model predictions at the high scale are matched and electroweak symmetry breaking at the proper scale is obtained. With the input parameters in Table 1, we obtain a spectrum which is mostly at the multi-TeV scale, with only the two lightest neutralinos

	n_1	n_2	n_3	m_{GM}	m_A	$\tan\beta$	$\mu^{\overline{\text{DR}}}$	$M_1^{\overline{\text{DR}}}$	$M_2^{\overline{\text{DR}}}$
dM1600	20	28	11	190 GeV	2160 GeV	43.81	160 GeV	1.72 TeV	4.33 TeV
dM770	46	46	20	250 GeV	4050 GeV	47.66	160 GeV	5.37 TeV	9.47 TeV

Table 1: Input parameters and running gaugino masses for the benchmark points. Here $n_{1,2,3}$ refers to the messenger indices for $U(1)_Y$, $SU(2)_L$, and $SU(3)_C$ respectively, m_{GM} is the characteristic gauge-mediated soft mass per messenger pair. $\mu^{\overline{\text{DR}}}$ is the running higgsino mass parameter in the $\overline{\text{DR}}$ scheme evaluated at the soft mass scale (as defined e.g. in [25]); it generally differs from the on-shell μ -parameter [26] by a few percent. Moreover, m_A is the pole mass of the pseudoscalar Higgs, and $\tan\beta$ is the ratio of Higgs vacuum expectation values. For details about the precise definitions of n_i and m_{GM} see [12]. We also give the $\overline{\text{DR}}$ values of M_1 and M_2 at the soft mass scale, which are determined at the high scale from the input parameters by the usual minimal gauge mediation relations.

as well as the light chargino masses within current collider ranges, as well as the SM-like Higgs boson with a mass is within a ± 2 GeV theory uncertainty interval from the LHC central value of 125.5 GeV.

The physical chargino and neutralino masses are calculated by `SOFTSUSY` using the running mass parameters $\mu^{\overline{\text{DR}}}$, $M_1^{\overline{\text{DR}}}$ and $M_2^{\overline{\text{DR}}}$ as given in Table 1, which includes one-loop corrections [27]. The light chargino and neutralino spectrum, and the underlying on-shell μ and $M_{1,2}$ parameters are given by

$$\text{dM1600} \quad M_1 = 1.70 \text{ TeV}, M_2 = 4.36 \text{ TeV}, \mu = 165.66 \text{ GeV}, \tan\beta|_{m_Z} = 44, \quad (12)$$

$$M_{\tilde{\chi}_1^\pm} = 165.77 \text{ GeV}, M_{\tilde{\chi}_1^0} = 164.17 \text{ GeV}, M_{\tilde{\chi}_2^0} = 166.87 \text{ GeV}, m_h = 124 \text{ GeV};$$

$$\text{dM770} \quad M_1 = 5.30 \text{ TeV}, M_2 = 9.51 \text{ TeV}, \mu = 167.22 \text{ GeV}, \tan\beta|_{m_Z} = 48, \quad (13)$$

$$M_{\tilde{\chi}_1^\pm} = 167.36 \text{ GeV}, M_{\tilde{\chi}_1^0} = 166.59 \text{ GeV}, M_{\tilde{\chi}_2^0} = 167.63 \text{ GeV}, m_h = 127 \text{ GeV}.$$

Note that the above masses include higher-order corrections, and therefore differ from the tree-level predictions in Eq. (7) at the sub-permil level.² As stated before, the two lightest neutralinos are nearly completely higgsino-like, with negligible gaugino-admixtures. The mass splittings are very small with $M_{\tilde{\chi}_1^\pm} - M_{\tilde{\chi}_1^0} = 770 \text{ MeV}$ (1.6 GeV) and $M_{\tilde{\chi}_2^0} - M_{\tilde{\chi}_1^0} = 1.04 \text{ GeV}$ (2.7 GeV) in the dM770 (dM1600) scenario. This has interesting consequences especially for the decay, which we will discuss in the following.

Production

The production cross sections have been calculated with `Whizard 1.95` [28] using the ILC setup provided by the ILC Generator Group [29]. It includes soft initial state photon emission using an electron structure function [30] as well as the beam energy spectra corresponding to the ILC Technical Design Report (ILC TDR) [31].

Due to the small mass difference, only very few and soft visible particles will result from the higgsino decay. For reasons mentioned above, a realistic analysis will therefore need to require an additional photon from initial state radiation to be seen in the detector. The ISR photon, in addition to its usefulness to suppress background, provides an elegant way to determine the $\tilde{\chi}_1^\pm$ and $\tilde{\chi}_2^0$ masses from their recoil against the photon. This has already been applied successfully for near-degenerate winos in AMSB scenarios [32].

We therefore determine the observable cross sections with one final state photon included in the hard matrix element. This photon is required to have a minimal invariant mass of 4 GeV with the corresponding beam electron and an energy of more than 10 GeV, corresponding to the cut used later at the analysis stage. We have verified that the effect of double-counting with the soft initial state photon emission is negligible.

Figure 2 shows the observable chargino and neutralino cross sections in the dM770 scenario as a function of the centre-of-mass energy, and for two different configurations of the polarisation of the beams, with

²These small differences in absolute masses, however, translate into sizeable corrections of the mass differences between the lightest chargino/second-lightest neutralino and the LSP. We accordingly correct for these higher-order contributions in the fit used for the determination of the fundamental parameters M_1 , M_2 , μ , and $\tan\beta$; see Section 5 for further details.

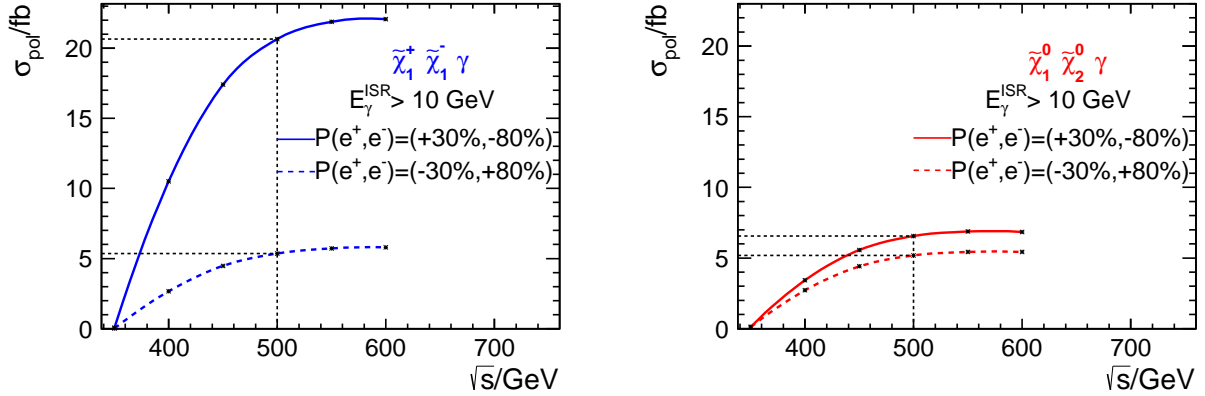


Figure 2: Chargino (left) and neutralino (right) production cross section in scenario dM770 for $P(e^+, e^-) = (-30\%, +80\%)$ and $P(e^-, e^+) = (+30\%, -80\%)$ as a function of the centre-of-mass energy.

the positron beam polarised to $-(+)$ 30%, the electron beam to $+(-)$ 80%.³ Figure 3 shows the same cross sections for a fixed centre-of-mass energy of 500 GeV as a function of the positron polarisation. The chargino production can be strongly enhanced by choosing the appropriate beam helicities, whereas the differences are significantly smaller in the neutralino case. Beyond the signal enhancement, beam polarisation can be used to prove that the observed process is mediated purely by s -channel exchange of a vector particle, i.e. by showing that the cross section vanishes for like-sign helicity configurations. Effects from beam polarisation at an $e^+ e^-$ collider on chargino/neutralino pair-production have already been studied in [15, 18, 21]⁴ and have successfully been exploited for full parameter determination in [37].

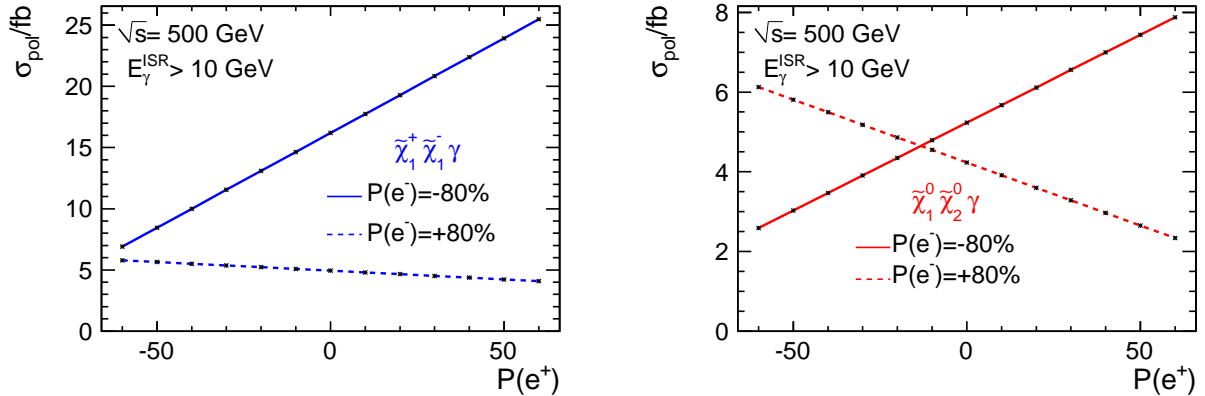


Figure 3: Chargino (left) and neutralino (right) production cross section in scenario dM770 at $\sqrt{s} = 500$ GeV as a function of the positron polarisation for $P(e^-) = +80\%$ and $P(e^-) = -80\%$.

As mentioned, higgsinos are produced by Z (γ, Z) exchange in the s -channel for neutralinos (charginos) respectively, cf. the diagrams shown in Fig. 1. It is then straightforward to explain the differences in production cross sections in Figs. 2 and 3 by considering only these diagrams. For neutralino production, the ratio of $(e_R^+ e_L^- \rightarrow \tilde{\chi}_1^0 \tilde{\chi}_2^0)/(e_L^+ e_R^- \rightarrow \tilde{\chi}_1^0 \tilde{\chi}_2^0) \sim 1.3$ for fully polarised beams can directly be obtained from the ratio of the respective fermion couplings to the Z boson, given by $L^2/R^2 = [1 - 1/(2 \sin^2 \theta_W)]^2 \sim 1.3$, with $L = -\frac{1}{2} + \sin^2 \theta_W$, $R = \sin^2 \theta_W$. For arbitrarily polarised beams the scaling factor between polarised

³In the following, the notation $P(e^+, e^-) = (-30\%, +80\%)$ etc. will be used to denote the assumed beam polarisation.

⁴For a (fully differential) chargino production at NLO, cf. [20, 33, 34, 35, 36].

$\tilde{\chi}_1^+$ decay mode	BR(dM1600)	BR(dM770)
$e\nu\tilde{\chi}_1^0$	17.3%	15.0%
$\mu\nu\tilde{\chi}_1^0$	16.6%	13.7%
$\pi^+\tilde{\chi}_1^0$	16.5%	60.4%
$\pi^+\pi^0\tilde{\chi}_1^0$	28.5%	7.3%
$\pi^+\pi^0\pi^0\tilde{\chi}_1^0$	7.5%	0.03%
$\pi^+\pi^+\pi^-\tilde{\chi}_1^0$	7.1%	0.03%
$\pi^+\pi^+\pi^-\pi^0\tilde{\chi}_1^0$	2.4%	–
$\pi^+\pi^0\pi^0\pi^0\tilde{\chi}_1^0$	0.5%	–
$K^+\tilde{\chi}_1^0$	1.2%	3.5%
$K^0\pi^+\tilde{\chi}_1^0$	1.0%	0.03%
$K^+\pi^0\tilde{\chi}_1^0$	0.5%	0.02%

Table 2: Chargino $\tilde{\chi}_1^+$ decay modes according to **Herwig++ 2.6.0**.

and unpolarised cross section is given by

$$\sigma_{\text{pol}}/\sigma_{\text{unpol}} = \left(1 + \frac{R^2 - L^2}{R^2 + L^2} P_{\text{eff}}\right) \cdot \mathcal{L}_{\text{eff}}/\mathcal{L} = (1 - 0.151 P_{\text{eff}}) \cdot \mathcal{L}_{\text{eff}}/\mathcal{L}, \quad (14)$$

where $\mathcal{L}_{\text{eff}} = (1 - P(e^-)P(e^+))\mathcal{L}$ denotes the normalized effective number of interactions and $P_{\text{eff}} = \frac{P(e^-) - P(e^+)}{1 - P(e^-)P(e^+)}$ the effective polarisation (for details see [38]). It is expected to achieve high $|P_{\text{eff}}| \sim 95\%$ and $\mathcal{L}_{\text{eff}}/\mathcal{L} \sim 1.5$ for center-of-mass energies of $\sqrt{s} \geq 350$ GeV at the ILC [39] which leads to scaling factor of about 1.3 (1.7) with $P(e^-, e^+) = (+80\%, -60\%)$ ($P(e^-, e^+) = (-80\%, +60\%)$) configuration of polarised (e^-, e^+)-beams.

In case of chargino pair production, the large difference in the cross sections stems from positive (in LR configuration) or negative (in RL configuration) interference of the γ and Z s -channel contributions, which is again due to the structure of the fermion- Z couplings. These characteristics (which are independent of additional photon radiation) are well reflected in the total cross sections displayed in Figures 2 and 3.

Decay widths and branching ratios

Due to the quite pronounced mass-degeneracy in the $\tilde{\chi}_1^\pm, \tilde{\chi}_1^0$ system, the $\tilde{\chi}_1^\pm$ decays require a specific treatment, which we will now briefly describe. In case of small mass splittings, a perturbative treatment on parton level, followed by standard hadronization, leads to large uncertainties for slight variations of the masses of the partonic decay products. In the scenarios considered here, this especially plays a role in the decays of the chargino, where the mass difference can be in the sub-GeV range. In fact, the same situation is encountered in the SM for the decay of τs into hadrons. In this case, a more appropriate description then follows an effective theory approach using hadronic currents. For the case of supersymmetry, such a description has been implemented in the event generator **Herwig++** for $\tilde{\chi}_1^\pm$ decays [40, 41]. We therefore used **Herwig++** to generate the according branching ratios and decays widths; they are given in Table 2. On the other hand, the branching ratios of the $\tilde{\chi}_2^0$ are calculated at parton-level followed by hadronization, with the results listed in Table 3.⁵ The decay widths of loop-induced processes have been cross checked with available analytic results [22]. They agree within 5 % with the **Herwig++** values.

⁵In principle, the mass difference for the $\tilde{\chi}_1^\pm$ -LSP in dM1600 scenario and the $\tilde{\chi}_2^0$ -LSP in dM770 scenario are of similar order. However, since for neutral current decays there is no analogue to the charged current τ -decay in the the SM, no dedicated code is yet available which treats the decays of $\tilde{\chi}_2^0$ using hadronic currents. The effects of including such an improved description are in the line of future work.

$\tilde{\chi}_2^0$ decay mode	BR(dM1600)	BR(dM770)
$\gamma\tilde{\chi}_1^0$	23.6%	74.0%
$\nu\bar{\nu}\tilde{\chi}_1^0$	21.9%	9.7%
$e^+e^-\tilde{\chi}_1^0$	3.7%	1.6%
$\mu^+\mu^-\tilde{\chi}_1^0$	3.7%	1.5%
hadrons + $\tilde{\chi}_1^0$	44.9%	12.7%
$\tilde{\chi}_1^\pm + X$	1.9%	0.4%

Table 3: Neutralino $\tilde{\chi}_2^0$ decay modes according to Herwig++ 2.6.0.

3.2 Event generation and detector simulation

Event generation

The simulation part of this study has been performed for a centre-of-mass energy of $\sqrt{s} = 500$ GeV. Signal events for the two benchmark scenarios have been generated with *Whizard*, using the same set-up as for the cross section calculation, described in Section 3.1. The charginos and neutralinos were generated by *Whizard*, with the subsequent decays simulated by *Pythia* [42]. The decay-branching fractions used by *Pythia* were those given in Table 2 and 3. As for the cross section calculation, one photon was included in the hard matrix element, and it was required to have a minimal invariant mass of 4 GeV with the corresponding beam electron. The generator-level cut on the photon energy was however not applied in the event generation step, in order to be able to estimate the experimental acceptance after reconstruction.

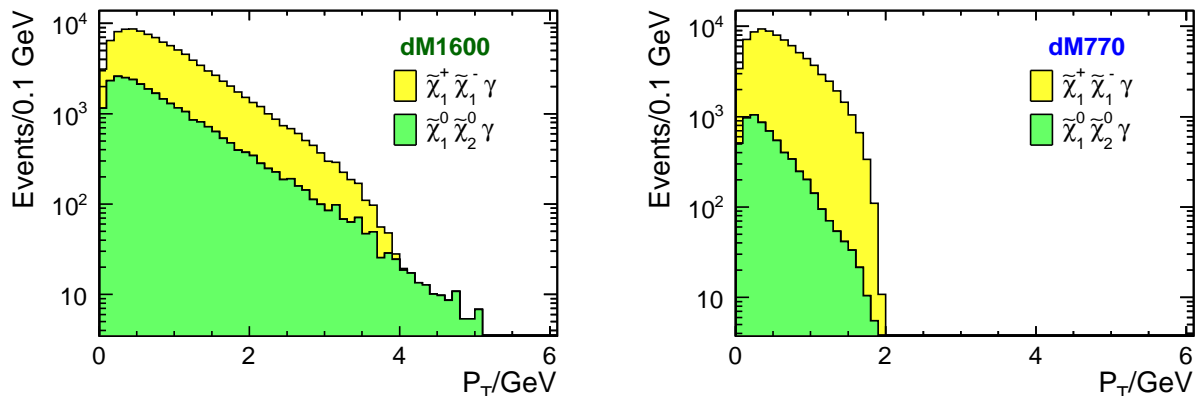


Figure 4: p_t spectrum of visible chargino and neutralino decay products at generator level for the two scenarios. Left: dM1600, right: dM770.

Figure 4 shows the transverse momentum spectrum of the visible decay products of chargino and neutralino events at generator level for both benchmark scenarios. Especially in the dM770 scenario, the visible decay particles of the higgsinos are extremely soft, with $p_t < 2$ GeV. While for larger momenta the tracking at the ILC is almost 100% efficient [43], the situation is more challenging for sub-GeV particles, especially due to the presence of random hits and additional tracks caused by e^+e^- -pair background from beamstrahlung. Therefore we will discuss the tracking efficiency in more detail in the following subsection.

The event samples centrally produced for the detector benchmarking in context of the ILC TDR [29] were used as SM background. Hence, the same *Whizard* version was used both for signal and background generation. Apart from some processes with very high cross sections, the available statistics corresponds at least to an integrated luminosity of 500 fb^{-1} . A notable exception was the $e\gamma \rightarrow 3f$ class, for which only

an integrated luminosity of about 100 fb^{-1} was available. This was taken care of by assigning appropriate weights to these events. Furthermore, radiative Bhabha events and pure photon final states deserve special comments in the context of this analysis. Radiative Bhabha events, i.e. $e^+e^- \rightarrow e^+e^-\gamma$, have been generated in a restricted phase space requiring an invariant mass of at least 4 GeV between any two of the final state particles, a momentum transfer of $Q > 4 \text{ GeV}$ between the incoming and outgoing electron (positron) and between the incoming lepton and the photon, resulting in a minimum acolinearity of the final state leptons. Pure photon final states have not been included in the production. However it has been shown in a previous full simulation study of a pure photon plus missing energy signature that the contribution after a very similar selection is negligible [44]. Also not included are final states with only neutrinos and photons as well as QED Compton events. They could in principle contribute if overlaid with one of the low- p_t $\gamma\gamma \rightarrow$ hadrons events of which on average 1.2 occur per bunch-crossing for TDR machine parameters. However, it is rather improbable that such an overlay will mimic the exclusive decay modes of the higgsinos studied here. A quantitative estimate of the impact of overlay events is only meaningful in full, **Geant4**-based [45] simulation and thus beyond the scope of this study.

Detector simulation

A detailed description of the ILD detector can be found in [43]. The most important detector capabilities for this analysis comprise the tracking efficiency at very low momenta, the photon reconstruction and the hermeticity of the detector, including the ability to veto high energetic electrons in the forward calorimeters (LumiCal, LHCAL and BeamCal). The tracking system is based on a Time Projection Chamber (TPC), augmented with silicon trackers on the inside, the outside and in the forward direction, as well as for vertexing. The transverse momentum resolution ($\Delta(1/p_T) = \Delta(p_T)/p_T^2$) is expected to be $2.0 \times 10^{-5} \text{ GeV}^{-1}$ asymptotically, worsening to $9.0 \times 10^{-5} \text{ GeV}^{-1}$ at 10 GeV , and to $9.0 \times 10^{-4} \text{ GeV}^{-1}$ at 1 GeV . In the low angle region, charged particles will be efficiently detected down to $\theta = 7^\circ$ ($|\cos\theta| = 0.993$), while the only region not in the acceptance of the calorimetric system are the holes in the BeamCal for the beam-pipes. Around the outgoing beam-pipe, the radius of the hole is 20 mm at $z = 3550 \text{ mm}$, corresponding to 5.6 mrad . Since the crossing angle of the beams is 14 mrad and the hole for the incoming beam-pipe has $R = 16 \text{ mm}$, the lower edge of the acceptance increases to 18.5 mrad at $\phi \approx 180^\circ$. The electromagnetic calorimeter (ECAL) is a highly granular SiW sampling calorimeter with a transverse cell size of $5 \text{ mm} \times 5 \text{ mm}$ and 20 layers. In test-beam measurements with a prototype detector a resolution of $(16.6 \pm 0.1)/\sqrt{E(\text{GeV})} \oplus (1.1 \pm 0.1)\%$ has been achieved [46].

We simulate the response of the ILD detector using the fast detector simulation **SGV** [47]. **SGV** derives the full covariance matrix of the track parameter reconstruction from the ILD detector geometry, taking into account point resolutions, positions of measurement planes and material effects like multiple scattering, bremsstrahlung and pair conversion. The calorimeter response has been parametrised to model the performance of the ILD particle flow reconstruction.

The response of the BeamCal is modelled taking into account the large energy depositions from beam backgrounds. The efficiency to tag an electron or photon on top of the beam background has been studied in full simulation and parametrised as function of the energy of the particle and the local background energy density at its impact point on the BeamCal surface.

A special challenge for this analysis is the tiny mass gap, which leads to extremely low momenta of the visible higgsino decay products. Beyond what was done for the TDR studies, the tracking efficiency in the presence of e^+e^- pair background from beamstrahlung was determined in full simulation and reconstruction of the ILD detector. In order to have a conservative estimate, e^+e^- pairs generated with **GuineaPig** [48] were overlaid on events with much higher particle-multiplicity ($t\bar{t}$ events) than what is expected from the signal studied here. These events were treated by **Mokka** [49], the full **Geant4**-based simulation of the ILD detector. For momenta larger than 1 GeV , track finding efficiency was found to be 99.5% . For low- p_t tracks, the track finding efficiency is shown as a function of p_t and as function of $\cos\theta$ in Figure 5. It can be seen that for beam background conditions as expected at $\sqrt{s} = 500 \text{ GeV}$, the tracking system is nearly fully efficient down to p_t values around 400 MeV , dropping to about 60% for p_t between 200 and 350 MeV .⁶ For all angles, the average tracking efficiency for all tracks with $p_t < 5 \text{ GeV}$ was found to be larger than 99% in

⁶For $p_t < 200 \text{ MeV}$, the tracking efficiency has not been studied and therefore we set it to zero here.

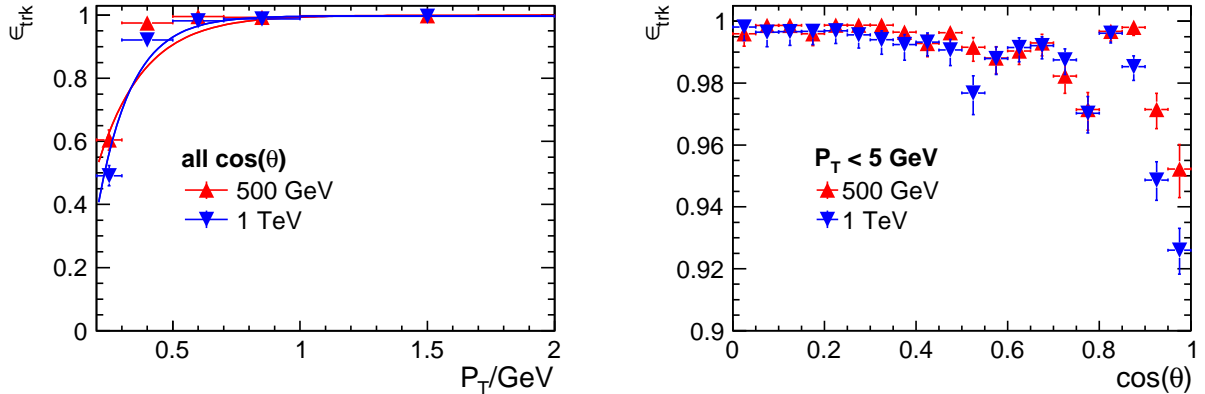


Figure 5: Tracking efficiency in $t\bar{t}$ events in the presence of pair background from beamstrahlung for $\sqrt{s} = 500 \text{ GeV}$ and $\sqrt{s} = 1 \text{ TeV}$, from full simulation of the ILC detector, assuming TDR beam parameters. Left: as a function of the transverse momentum p_t , Right: as a function of the absolute value of the cosine of the polar angle $|\cos \theta|$ for tracks with $p_t < 5 \text{ GeV}$.

the barrel region, dropping slightly in the transition region between the barrel and the end-cap, but staying larger than 95% everywhere in the geometric acceptance of the tracking system.

In SGV, we implemented four parametrisations of the efficiency as function of p_t for different polar angle ranges and applied it to the signal and the dominating background class, namely the $e\gamma \rightarrow 3f$ processes, which is basically the only type of events remaining at the selection stage where the tracking efficiency for low-energetic particles enters.

4 Experimental precision at the ILC

In this section, we describe the event selection and the extraction of masses and cross sections in the two benchmark scenarios. All numbers refer to an integrated luminosity of $\int \mathcal{L} dt = 500 \text{ fb}^{-1}$ each for two configurations of the beam polarisation, namely $P(e^+, e^-) = (+30\%, -80\%)$ and $P(e^+, e^-) = (-30\%, +80\%)$. With the current machine design, collecting 1 ab^{-1} of data at $\sqrt{s} = 500 \text{ GeV}$ takes 4 years of design ILC operation, based on typical machine availability times and efficiencies.

As a first step, a common pre-selection discriminating higgsino events against the majority of SM background processes is performed. Based on this, two different sets of cuts are applied for selecting either chargino or neutralino events and measuring the respective masses and cross sections.

4.1 Preselection

The aim of the preselection is to suppress the majority of SM background, while not yet discriminating charginos from neutralinos. The key features of the signal are that the events only contain a few low energetic but centrally produced decay products of the higgsinos, and the hard ISR photon. Thus, the following selection criteria are applied:

- There should be no significant activity in the BeamCal beyond the beam background in order to suppress events with high energetic electrons scattered under a small angle, e.g. from $\gamma\gamma$ and $e\gamma$ type of processes.
- The total number of reconstructed particles N_{RP} should be less than 15.
- There should be exactly one reconstructed photon with $E_{\text{ISR}} > 10 \text{ GeV}$ and a polar angle θ_{ISR} fulfilling $|\cos \theta_{\text{ISR}}| < 0.993$. The latter condition ensures that the photon is within the acceptance of the tracking system in order to be able to distinguish photons and electrons.

- Any other reconstructed particle in the event is required to be at least 20° away from the beam axis, and thus within the acceptance of the TPC, and to have an energy E_{soft} of less than 5 GeV.⁷
- Finally, the missing energy E_{miss} should be larger than 300 GeV. This rejects the class of events where either the beam electron or positron is scattered under very small angles such that it escapes through the beam-pipe with an energy of roughly 250 GeV.
- In order to ensure that the missing energy is not due to particles escaping along the beam pipe, the missing momentum vector is required to point well into the acceptance region of the detector by fulfilling $|\cos \theta_{\text{miss}}| < 0.992$.

The resulting cut-flow is displayed for $P(e^+, e^-) = (+30\%, -80\%)$ in Table 4. At this stage of the selection, the SM background is reduced by four orders of magnitude and is dominated by $e\gamma \rightarrow 3f$ and $\gamma\gamma \rightarrow 2f$ processes. The biggest loss in the signals is due to the ISR photon requirement, since at event generation only a minimal invariant mass of the photon with respect to the beam electron/positron of 4 GeV was required, while the preselection demands a photon energy of at least 10 GeV, the latter being more stringent.

	dM1600		dM770		Standard Model		
	$\tilde{\chi}_1^+ \tilde{\chi}_1^- \gamma$	$\tilde{\chi}_1^0 \tilde{\chi}_2^0 \gamma$	$\tilde{\chi}_1^+ \tilde{\chi}_1^- \gamma$	$\tilde{\chi}_1^0 \tilde{\chi}_2^0 \gamma$	$ee \rightarrow 2, 4, 6f$	$e\gamma \rightarrow 3, 5f$	$\gamma\gamma \rightarrow 2, 4f$
no cut	38672	24250	38130	23940	2.6434×10^7	8.8820×10^7	9.7554×10^8
BeamCal veto	38591	24187	38054	23874	2.6284×10^7	8.8178×10^7	9.6757×10^8
$N_{\text{RP}} < 15$	38591	24185	38054	23874	6.4968×10^6	6.5811×10^7	6.6308×10^8
$N_{\text{ISR}} = 1$	30058	9551	29675	9317	3.1640×10^6	1.5074×10^7	1.7752×10^7
$ \cos \theta_{\text{soft}} < 0.9397$	21501	7318	23117	7458	7.1453×10^5	4.5646×10^6	4.7083×10^6
$E_{\text{soft}} < 5 \text{ GeV}$	20611	6615	22156	7110	9092	5.9732×10^5	1.2390×10^6
$E_{\text{miss}} > 300 \text{ GeV}$	20611	6615	22156	7110	6462	1.5822×10^5	4.6306×10^5
$ \cos \theta_{\text{miss}} < 0.992$	19872	6365	21558	6872	5731	1.1837×10^5	3.3051×10^5

Table 4: Cut-flow table of the preselection for an integrated luminosity of 500 fb^{-1} and $P(e^+, e^-) = (+30\%, -80\%)$.

4.2 Chargino selection

In order to discriminate chargino pair events from the associated neutralino production, we select semi-leptonic events, i.e. events with one chargino decaying leptonically and the other hadronically. This signature does not appear in the associated neutralino production, since the $\tilde{\chi}_2^0$ decays via a virtual Z boson and thus either fully leptonically or fully hadronically. The main background contributions to the semi-leptonic signature are τ -pairs, either from $e^+e^- \rightarrow \tau^+\tau^-$, from $e^\pm\gamma \rightarrow e^\pm\tau^+\tau^-$ or from $\gamma\gamma \rightarrow \tau^+\tau^-$.

The cleanest semi-leptonic signature arises when the hadronic chargino decays to a single charged pion. In the dM770 scenario, this signature occurs in as many 35% of the chargino pair events, cf. Table 2. In the dM1600 scenario, the branching ratio to single pions is significantly lower due to the larger mass difference, so that only 11% of the chargino pairs give the semi-leptonic single pion signature. Therefore the $\pi^\pm\pi^0$ channel is included for the dM1600 scenario by allowing the charged pion to be accompanied by two photons, raising the fraction of events with the target signature to 30%. For the dM770 scenario, the gain in signal is too small to balance the penalty of larger background, and thus only the single pion mode is allowed in that scenario.

⁷For charged particles, the energy is reconstructed from the momentum measured in the tracking system assuming the π^\pm mass.

A very discriminative variable against τ -events can be constructed from hadronic chargino decay:

$$E_{\pi}^* = \frac{(\sqrt{s} - E_{\gamma})E_{\pi} + \vec{p}_{\pi} \cdot \vec{p}_{\gamma}}{\sqrt{s'}}. \quad (15)$$

The reduced center-of-mass energy $\sqrt{s'}$ of the system recoiling against the ISR photon is defined as $s' = s - 2\sqrt{s}E_{\gamma}$, where $\sqrt{s} = 500 \text{ GeV}$ is the nominal center-of-mass energy. E_{π}^* is related to the energy of the pion(s) in the rest frame of the chargino-pair, which is only a few GeV for pions from chargino decays, but large for τ decays.⁸ Therefore, we require $E_{\pi}^* < 3 \text{ GeV}$. As can be seen from Table 5, this cut is fully efficient in the dM770 scenario, but not in the dM1600 scenario. The reason is that the values of E_{π}^* are closely related to the mass difference between the chargino and the LSP, which is nearly twice as large in the dM1600 scenario. However the losses occur in the less relevant regime far from the endpoint. At this stage, the background is dominated by contributions from $\gamma\gamma$ events, in particular $\gamma\gamma \rightarrow \tau^+\tau^-$, which are typically back-to-back in the transverse plane and have large $\sqrt{s'}$. Therefore the combined requirement of $\Phi_{\text{acop}} < 2$ or $\sqrt{s'} < 480 \text{ GeV}$, where Φ_{acop} is the acoplanarity angle between the leptonic and the hadronic system, reduced this background by a factor of four.

	dM1600		Standard Model		
	$\tilde{\chi}_1^+ \tilde{\chi}_1^- \gamma$	$\tilde{\chi}_1^0 \tilde{\chi}_2^0 \gamma$	$ee \rightarrow 2, 4, 6f$	$e\gamma \rightarrow 3, 5f$	$\gamma\gamma \rightarrow 2, 4f$
after preselection	19872	6365	5731	1.1837×10^5	3.3051×10^5
$l^{\pm} \pi^{\pm} (\pi^0)$	5509	134	38	6197	13991
$E_{\pi}^* < 3 \text{ GeV}$	4435	103	0	2635	6162
$\Phi_{\text{acop}} < 2$ or $\sqrt{s'} < 480 \text{ GeV}$	3813	97	0	2564	1452
$E_{\text{miss}} > 350 \text{ GeV}$	3812	97	0	1016	511

	dM770		Standard Model		
	$\tilde{\chi}_1^+ \tilde{\chi}_1^- \gamma$	$\tilde{\chi}_1^0 \tilde{\chi}_2^0 \gamma$	$ee \rightarrow 2, 4, 6f$	$e\gamma \rightarrow 3, 5f$	$\gamma\gamma \rightarrow 2, 4f$
after preselection	21558	6872	5731	1.1837×10^5	3.3051×10^5
$l^{\pm} \pi^{\pm}$	5489	38	19	2478	6754
$E_{\pi}^* < 3 \text{ GeV}$	5489	38	0	1465	4755
$\Phi_{\text{acop}} < 2$ or $\sqrt{s'} < 480 \text{ GeV}$	4600	36	0	1417	782
$E_{\text{miss}} > 350 \text{ GeV}$	4599	36	0	536	218

Table 5: Cut flow tables of the chargino selection in the two scenarios for an integrated luminosity of 500 fb^{-1} and $P(e^+, e^-) = (+30\%, -80\%)$ following the preselection (cf. Tab. 4). For the dM1600 scenario, also the two-pion decay mode is included, therefore also the SM expectation changes. The chargino mass measurement is performed after the cut on the acoplanarity angle and the reduced center of mass energy in both scenarios. The respective last line is applied for the cross section measurement and the determination of the mass difference.

Table 5 shows that at this level only a very small contribution from $\tilde{\chi}_1^0 \tilde{\chi}_2^0$ events remains in the sample. This is important since we can assume the ability to model electroweak SM processes to any necessary precision by the time an ILC is actually running, whereas background contributions from other new physics processes, in our case the neutralino production, are a priori unknown.

⁸In case of the additional π^0 , the reconstructed four-momenta of the two photons are added to the charged pion four-momentum.

Chargino mass reconstruction

The chargino mass is reconstructed using the reduced center-of-mass energy, $\sqrt{s'}$ introduced above. At the threshold, the chargino pair is produced nearly at rest, and $\sqrt{s'}$ is twice the chargino mass, thus:

$$M_{\tilde{\chi}_1^\pm} = \frac{1}{2}\sqrt{s'} = \frac{1}{2}\sqrt{s - 2\sqrt{s}E_\gamma}. \quad (16)$$

Figure 6 shows the resulting $\sqrt{s'}$ distributions for both scenarios for an integrated luminosity of 500fb^{-1} with $P(e^+, e^-) = (+30\%, -80\%)$. The onset of the signal is clearly visible on top of the SM background, which stretches out to lower $\sqrt{s'}$ values. The cut-off in the SM background near $\sqrt{s'} = 230\text{ GeV}$ is due to the cut on $E_{\text{miss}} > 300\text{ GeV}$. This is chosen on purpose so that a signal-free region is available to fix the background level, here by fitting an exponential function with two free parameters $f(x) = p_1 \cdot e^{-p_2 \cdot x}$ to the SM prediction only (blue line). In a second step, a straight line is added on top of the background to model the signal contribution and fitted to the simulated data in the endpoint region (red line). The two parameters of the SM background function are fixed to the values obtained from the SM-only fit in the wider $\sqrt{s'}$ window. It has been verified that the results are stable against reasonable variations of the fit ranges or the bounds on the background parameters. The chargino mass is fitted to $168.0 \pm 1.4\text{ GeV}$ in the dM1600 scenario and to $168.6 \pm 1.0\text{ GeV}$ in the dM770 case.

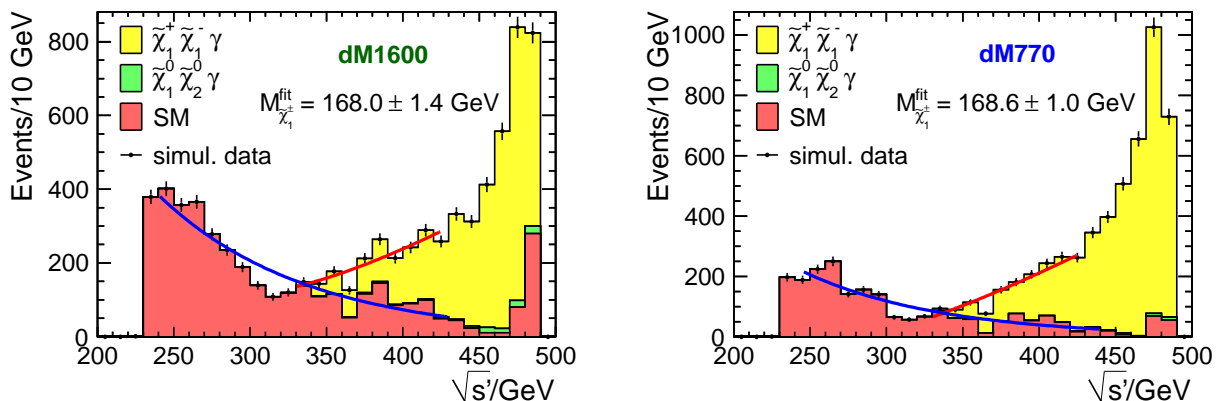


Figure 6: Distribution of the reduced center-of-mass energy ($\sqrt{s'}$) of the system recoiling against the hard ISR photon for all events passing the chargino selection for an integrated luminosity of 500fb^{-1} with $P(e^+, e^-) = (+30\%, -80\%)$. $M_{\tilde{\chi}_1^\pm}$ is determined from a linear fit to the distribution near the endpoint. Left: dM1600 scenario; Right: dM770 scenario.

The fitted central values agree within 1.6 (1.2) standard deviations with the respective input masses of $M_{\tilde{\chi}_1^\pm} = 165.77\text{ GeV}$ ($M_{\tilde{\chi}_1^\pm} = 167.36\text{ GeV}$) in the dM1600 (dM770) scenario. Since the relation between $\sqrt{s'}$ and the chargino mass is only approximate, e.g. due to the approximation of equal chargino energies, but even more so due to the beam energy spectrum, an exact agreement is not necessarily expected. Therefore we investigated the dependence of the fitted mass on the input mass by simulating signal samples with different chargino masses. In order to minimize a possible bias due to changes in the acceptance, all higgsino masses were varied simultaneously, so that e.g. the momentum distribution of the decay products does not change significantly.

Figure 7 shows the fitted mass as function of the true mass for both scenarios. They clearly display a linear behaviour, which can easily be used to calibrate the reconstruction method. The calibrated mass (and its uncertainty) can be found on the x -axis as a projection of the fitted values on the y -axis as indicated by the lines.⁹

⁹The uncertainty on the calibration curve itself is not propagated to the final result, since its origin, namely the limited amount of available MC statistics especially for the SM background, will not be an issue in a real ILC measurement.

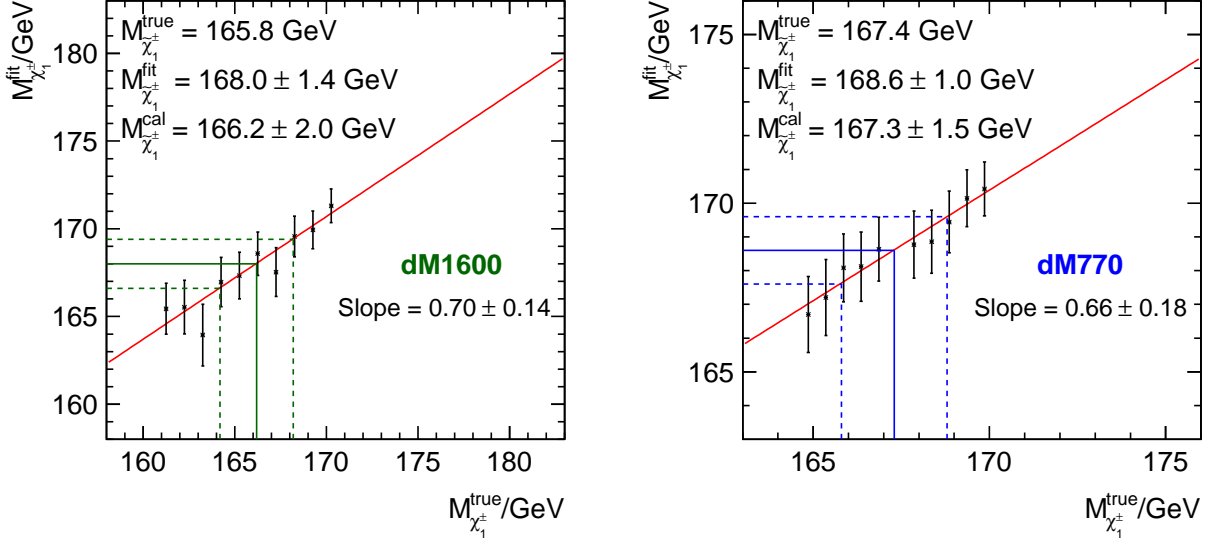


Figure 7: Calibration of the $\sqrt{s'}$ method for the chargino mass determination. Left: dM1600 scenario; Right: dM770 scenario.

At a first sight it might appear worrisome that both scenarios have different correlations between fitted and true masses. However one needs to remember that in the dM1600 case, an additional decay channel is included and thus different selections are applied in the two scenarios, which in this case changes the behaviour of the reconstruction. A substantial component of the uncertainty is correlated between all masses. This is due to the remaining background, and is most visible in the dM770 scenario, where the signal rate is higher and thus statistical fluctuations between the different signal samples at the different masses are smaller than in the dM1600 case. In the absence of any background, the statistical error on an individual mass determination would shrink to about 0.5 GeV.

After applying the calibration, we find an excellent agreement with the input mass values, at the price of the uncertainties increasing by $\sim 50\%$:

$$\text{dM1600} : M_{\tilde{\chi}_1^\pm}^{\text{cal}} = 166.2 \pm 2.0 \text{ GeV} \quad (M_{\tilde{\chi}_1^\pm}^{\text{true}} = 165.8 \text{ GeV}), \quad (17)$$

$$\text{dM770} : M_{\tilde{\chi}_1^\pm}^{\text{cal}} = 167.3 \pm 1.5 \text{ GeV} \quad (M_{\tilde{\chi}_1^\pm}^{\text{true}} = 167.4 \text{ GeV}). \quad (18)$$

Reconstruction of the chargino-LSP mass difference

After a clear observation of the signal in the $\sqrt{s'}$ spectrum and the determination of the chargino mass from the endpoint, the background can be further reduced by tightening the missing energy cut to $E_{\text{miss}} > 350 \text{ GeV}$. The resulting event count is displayed for both scenarios in the respective last line of Table 5.

The mass difference between the chargino and the LSP is very small in both scenarios, namely 1.6 GeV (770 MeV) in the dM1600 (dM770) scenario. It is nevertheless accessible via the energy of the visible chargino decay products in the chargino-pair rest frame, calculated according to the definition in Eq. (15). If the charginos are actually produced at rest, E_π^* can only take one value, *viz*:

$$E_\pi^* = \frac{(M_{\tilde{\chi}_1^\pm} - M_{\tilde{\chi}_1^0})(M_{\tilde{\chi}_1^\pm} + M_{\tilde{\chi}_1^0}) + m_\pi^2}{2M_{\tilde{\chi}_1^\pm}} = \frac{\Delta(M)\Sigma(M) + m_\pi^2}{2M_{\tilde{\chi}_1^\pm}} = \frac{1}{1/\Delta(M) + 1/\Sigma(M)} + \frac{m_\pi^2}{2M_{\tilde{\chi}_1^\pm}}. \quad (19)$$

As both the mass difference and m_π are several orders of magnitude smaller than the higgsino masses, E_π^* being equal to the mass difference itself to a very good approximation. This is illustrated by the upper row of plots in Figure 8, which show the selected events in the $E_\pi^* - \sqrt{s'}$ plane for both scenarios. The signal events

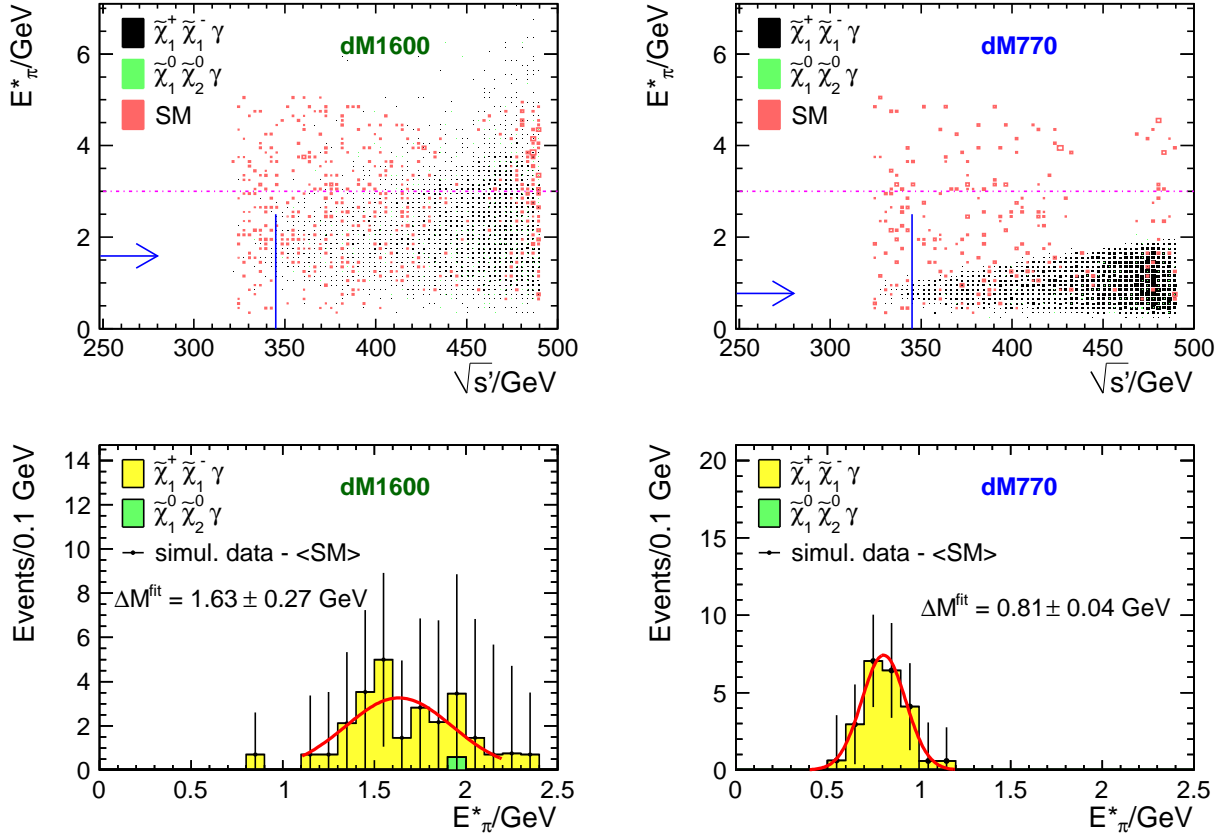


Figure 8: Measurement of the chargino-LSP mass difference for an integrated luminosity of 500fb^{-1} with $P(e^+, e^-) = (+30\%, -80\%)$. The upper row shows E_π^* vs. $\sqrt{s'}$. The horizontal dashed line indicates the cut on E_π^* , the vertical line the cut on $\sqrt{s'}$. The arrow indicates the input value of the chargino-LSP mass difference. The lower row the E_π^* after cutting on $\sqrt{s'}$. Left: dM1600 scenario; Right: dM770 scenario.

form a distinctive triangular region, with a tip just at the mass difference indicated by the blue arrow. We select the tip of the triangle by requiring $\sqrt{s'} < 345$ GeV as indicated by the blue line, and then determine the tip position from the projection of the remaining events onto the E_π^* axis. In a more complex analysis the whole two-dimensional $\sqrt{s'}$ distribution could be exploited to determine the mass difference, potentially even together with the cross section.

The resulting E_π^* distributions are displayed in the lower row of Figure 8, again for both scenarios. The mass difference is then reconstructed as the mean of a Gaussian fitted to the distribution. Due to the limited amount of available SM events in the simulation, we subtract the SM background, assuming that precise theoretical predictions and sufficient MC statistics would be available for a real ILC measurement. However the error bars include the full expected statistical fluctuations from the background. Especially in the case of the dM1600 scenario, the remaining signal is rather small, but still has a significance of more than 5σ . The mass difference is reconstructed as:

$$\text{dM1600} : \Delta M_{\tilde{\chi}_1^\pm - \tilde{\chi}_1^0}^{\text{rec}} = 1630 \pm 270 \text{ MeV} \quad (\Delta M_{\tilde{\chi}_1^\pm - \tilde{\chi}_1^0}^{\text{true}} = 1600 \text{ MeV}), \quad (20)$$

$$\text{dM770} : \Delta M_{\tilde{\chi}_1^\pm - \tilde{\chi}_1^0}^{\text{rec}} = 810 \pm 40 \text{ MeV} \quad (\Delta M_{\tilde{\chi}_1^\pm - \tilde{\chi}_1^0}^{\text{true}} = 770 \text{ MeV}). \quad (21)$$

These values agree with the input mass differences within the uncertainty.

Measurement of the polarised chargino cross sections

The achievable precision on the polarised cross sections is estimated based on the event counts in the final row of Table 5, i.e. after requiring $E_{\text{miss}} > 350 \text{ GeV}$, according to

$$\frac{\delta\sigma}{\sigma} = \frac{1}{\sqrt{\epsilon \cdot \pi \cdot \sigma \cdot \int \mathcal{L} dt}}. \quad (22)$$

Table 6 displays the corresponding efficiencies (ϵ) and purities (π) as well as the resulting precision on the cross sections assuming an integrated luminosity of 500 fb^{-1} for each polarisation configuration. Since the efficiencies in this case refer to the total number of produced chargino events regardless of their decay channel, we display for information the branching ratios of the decay modes to which the selections are tailored, as well as the assumed production cross sections.

	$P(e^+, e^-) = (+30\%, -80\%)$		$P(e^+, e^-) = (-30\%, +80\%)$	
	dM1600	dM770	dM1600	dM770
$\int \mathcal{L} dt$	500 fb^{-1}	500 fb^{-1}	500 fb^{-1}	500 fb^{-1}
σ	78.7 fb	77.0 fb	20.4 fb	19.9 fb
BR of selected mode(s)	30.5%	34.7%	30.5%	34.7%
efficiency ϵ	9.9%	12.1%	9.5%	12.2%
purity π	70.1%	85.3%	36.4%	56.1%
$\delta\sigma/\sigma$	1.9%	1.6%	5.3%	3.8%

Table 6: Efficiency, purity, and relative statistical precision on the visible cross section for charginos. The cross section values given here are higher than those displayed in Fig. 2 and 3 since in the event generation step no generator level cut on the energy of the ISR photon was applied, cf. Sec. 3.2.

For the signal-enriching choice of $P(e^+, e^-) = (+30\%, -80\%)$, the cross section can be determined to 1.9% (1.6%) in the dM1600 (dM770) scenario. For $P(e^+, e^-) = (-30\%, +80\%)$, the signal cross sections are lower, while the dominating SM background processes do not depend on the polarisation. Therefore the resolution worsens to 5.3% and 3.8% in the two scenarios, respectively. These statistical uncertainties refer to the cross section for the selected semi-leptonic final state. The branching ratio is determined completely by the mass difference. Therefore the total production cross section can be extracted, where the additional uncertainty on the branching ratio from the limited knowledge of the mass difference needs to be included in addition. The corresponding numbers will be given in Section 5.

4.3 Neutralino selection

The neutralino events are selected by the decay to photon and the LSP. For the dM770 scenario, this is the dominant decay by far, occurring in 74% of the events. In the dM1600 case, hadronic decay modes dominate, but the branching ratio for the very clean radiative decay is still sizable with 24%. Therefore, after the common preselection described in Section 4.1, events with only photons in the final state are selected by applying a veto against reconstructed leptons and hadrons. All “soft” photons, i.e. all photons apart from the hard ISR photon candidate, are required to fulfill $|\cos \theta_{\gamma_{\text{soft}}}| < 0.85$. If more than one “soft” photon is reconstructed in addition to the hard ISR photon, then the candidate with the highest transverse momentum is considered as the neutralino decay photon. In analogy to E_π^* in the chargino case, the variable $E_{\gamma_{\text{soft}}}^*$ is defined as

$$E_{\gamma_{\text{soft}}}^* = \frac{(\sqrt{s} - E_\gamma) E_{\gamma_{\text{soft}}} + \vec{p}_{\gamma_{\text{soft}}} \cdot \vec{p}_\gamma}{\sqrt{s'}}, \quad (23)$$

and $E_{\gamma_{\text{soft}}}^* > 0.5$ is required at the final selection step. The resulting event counts are summarized in Table 7.

	dM1600		dM770		Standard Model		
	$\tilde{\chi}_1^+ \tilde{\chi}_1^- \gamma$	$\tilde{\chi}_1^0 \tilde{\chi}_2^0 \gamma$	$\tilde{\chi}_1^+ \tilde{\chi}_1^- \gamma$	$\tilde{\chi}_1^0 \tilde{\chi}_2^0 \gamma$	$ee \rightarrow 2, 4, 6f$	$e\gamma \rightarrow 3, 5f$	$\gamma\gamma \rightarrow 2, 4f$
after preselection	19872	6365	21558	6872	5731	1.1837×10^5	3.3051×10^5
Photon final state	53	1733	155	5224	399	1217	2254
$ \cos \theta_{\gamma_{\text{soft}}} < 0.85$	38	1467	120	4538	233	800	1145
$E_{\gamma_{\text{soft}}}^* > 0.5$	19	1395	22	4095	109	242	413
$E_{\text{miss}} > 350 \text{ GeV}$	19	1395	22	4095	90	180	384

Table 7: Number of events passing the final neutralino selection, following the preselection in Tab. 4, for an integrated luminosity of 500 fb^{-1} and $P(e^+, e^-) = (+30\%, -80\%)$.

Neutralino mass reconstruction

As in the chargino case, the mass of the $\tilde{\chi}_2^0$ is reconstructed from $\sqrt{s'}$ determined using the ISR photon. As opposed to the chargino case, the masses of the two produced neutralinos are not equal, but their difference is smaller than the resolution of the $\sqrt{s'}$ method. Therefore we use the approximation of two equal mass particles being produced and calibrate the method in the end.

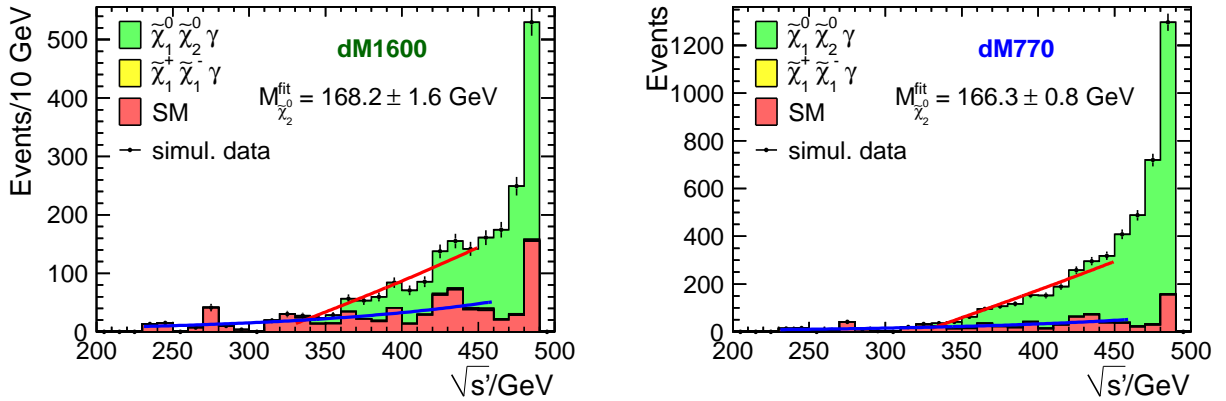


Figure 9: Distribution of the reduced center-of-mass energy ($\sqrt{s'}$) of the system recoiling against the hard ISR photon for all events passing the neutralino selection for an integrated luminosity of 500 fb^{-1} with $P(e^+, e^-) = (+30\%, -80\%)$. $M_{\tilde{\chi}_2^0}^{\text{fit}}$ is determined from fitting the sum (red curve) of a straight line for the signal and the background parametrisation (blue curve) to the distribution near the endpoint. Left: dM1600 scenario; Right: dM770 scenario.

Figure 9 shows the $\sqrt{s'}$ distribution obtained in both scenarios. The signal is clearly visible above the background, which has only a negligible contribution from chargino production. Like in the chargino case, the SM background is fitted first with an exponential (blue line). In a second step, a straight line is added on top of the background to model the signal contribution and fitted to the simulated data in the endpoint region (red line). Again the parameters of the SM background function are fixed to the values obtained from the SM only fit in the wider $\sqrt{s'}$ window. The neutralino mass is fitted to $168.2 \pm 1.6 \text{ GeV}$ in the dM1600 scenario and to $166.3 \pm 0.8 \text{ GeV}$ in the dM770 case. Both numbers agree with the input values within 1–1.5 σ .

Again the correlation between fitted mass and input mass is investigated, resulting in the calibration curve displayed in Figure 10. Since in both scenarios the same decay modes are selected, there is no reason to assume a different calibration behaviour in the other scenario. Therefore the same calibration curve is used. Since its slope is only about 0.5, the final statistical uncertainty on the calibrated neutralino masses

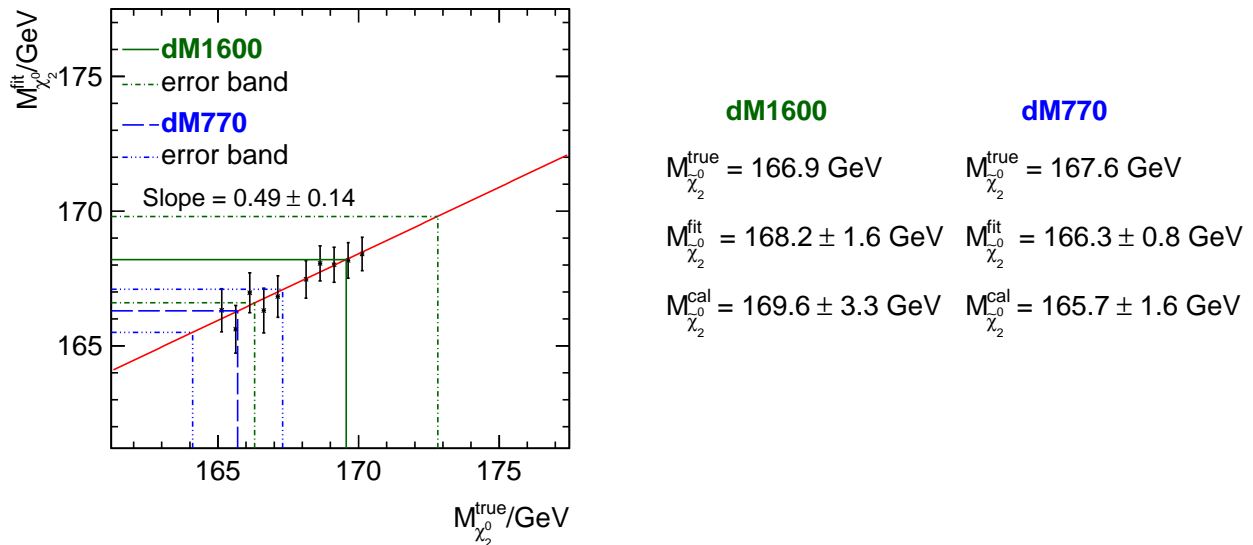


Figure 10: Calibration of the $\sqrt{s'}$ method for the neutralino mass determination. Since here the same decay mode is selected for both scenarios, the same calibration curve can be used.

grows by a factor of 2. After applying the calibration, the final results are:

$$\text{dM1600} : M_{\chi_2^0}^{\text{cal}} = 169.6 \pm 3.3 \text{ GeV} \quad (M_{\chi_2^0}^{\text{true}} = 166.9 \text{ GeV}), \quad (24)$$

$$\text{dM770} : M_{\chi_2^0}^{\text{cal}} = 165.7 \pm 1.6 \text{ GeV} \quad (M_{\chi_2^0}^{\text{true}} = 167.6 \text{ GeV}). \quad (25)$$

Measurement of the polarised neutralino cross sections

The achievable precision on the polarised cross sections is estimated like in the chargino case, based on the event counts given in Table 7. Table 8 displays the corresponding efficiencies and purities as well as the resulting precision on the cross sections assuming an integrated luminosity of 500 fb^{-1} for each polarisation configuration. Since the efficiencies in this case refer to the total number of produced neutralino events regardless of their decay channel, the branching ratios of the radiative decay are displayed in addition for information, as well as the assumed production cross sections.

	$P(e^+, e^-) = (+30\%, -80\%)$		$P(e^+, e^-) = (-30\%, +80\%)$	
	dM1600	dM770	dM1600	dM770
$\int \mathcal{L} dt$	500 fb^{-1}	500 fb^{-1}	500 fb^{-1}	500 fb^{-1}
σ	49.0 fb	48.4 fb	38.9 fb	38.4 fb
BR of selected mode	23.6%	74.0%	23.6%	74.0%
efficiency ϵ	5.8%	17.1%	6.0%	17.2%
purity π	67.4%	85.8%	62.3%	82.5%
$\delta\sigma/\sigma$	3.2%	1.7%	3.7%	1.9%

Table 8: Efficiency, purity, and relative statistical precision on the visible cross section for associated neutralino production. The cross section values given here are higher than those displayed in Fig. 2 and 3 since in the event generation step no generator level cut on the energy of the ISR photon was applied, cf. Sec. 3.2.

Observable	dM1600	dM770
$M_{\tilde{\chi}_1^\pm}$ [GeV]	166.2 ± 2.0	167.3 ± 1.5
$M_{\tilde{\chi}_2^0}$ [GeV]	169.6 ± 3.3	165.7 ± 1.6
$M_{\tilde{\chi}_1^\pm} - M_{\tilde{\chi}_1^0}$ [GeV]	1.63 ± 0.27	0.81 ± 0.04
$\delta\sigma/\sigma(\tilde{\chi}_1^+ \tilde{\chi}_1^- \gamma)_{(-0.3,0.8)}^{500}$	16%	4.3%
$\delta\sigma/\sigma(\tilde{\chi}_1^+ \tilde{\chi}_1^- \gamma)_{(0.3,-0.8)}^{500}$	15%	2.6%
$\delta\sigma/\sigma(\tilde{\chi}_1^+ \tilde{\chi}_1^- \gamma)_{(-0.3,0.8)}^{350}$	33%	20%
$\delta\sigma/\sigma(\tilde{\chi}_1^+ \tilde{\chi}_1^- \gamma)_{(0.3,-0.8)}^{350}$	18%	9.0%
$\delta\sigma/\sigma(\tilde{\chi}_1^0 \tilde{\chi}_2^0 \gamma)_{(-0.3,0.8)}^{500}$	4.0%	2.4%
$\delta\sigma/\sigma(\tilde{\chi}_1^0 \tilde{\chi}_2^0 \gamma)_{(0.3,-0.8)}^{500}$	3.5%	2.3%
$\delta\sigma/\sigma(\tilde{\chi}_1^0 \tilde{\chi}_2^0 \gamma)_{(-0.3,0.8)}^{350}$	20%	10%
$\delta\sigma/\sigma(\tilde{\chi}_1^0 \tilde{\chi}_2^0 \gamma)_{(0.3,-0.8)}^{350}$	18%	9.4%

Table 9: Observables used in the fit: masses with uncertainties taken from Eqs. (17), (18), (20), (21), (24), (25); relative cross sections uncertainties from Tabs. 8 and 6 including systematic uncertainty due to branching ratios as discussed in text. The subscript on $\delta\sigma/\sigma$ denotes the beam polarisation $P(e^+, e^-)$ and the superscript the center-of-mass energy in GeV.

For neutralino production, the total cross section is much less sensitive to beam polarisation than for chargino production, cf. Figure 2. Therefore, the precision depends only marginally on the beam polarisation. Due to the different branching ratios for the radiative decay, the cross sections can be determined more precisely, to about 1.7% in the dM770 scenario, while the measurements are about a factor two less precise in the dM1600 case.

5 Parameter determination

In the final step of the analysis, we estimate the accuracy to which underlying MSSM parameters can be determined. As discussed in Section 2.2, there are four real parameters that define the chargino and neutralino sector at tree-level:

$$M_1, \quad M_2, \quad \mu, \quad \tan\beta. \quad (26)$$

They can be extracted using the measurements discussed in the previous section: the masses of the light chargino $\tilde{\chi}_1^\pm$ and the neutralino $\tilde{\chi}_1^0$, the mass difference $M_{\tilde{\chi}_1^\pm} - M_{\tilde{\chi}_1^0}$, the chargino pair production cross section and the $\tilde{\chi}_1^0 \tilde{\chi}_2^0$ production cross section. The respective values and uncertainties are summarised in Tab. 9. Since we have not performed a detector simulation for $\sqrt{s} = 350$ GeV, the uncertainties in this case are obtained by rescaling the $\sqrt{s} = 500$ GeV result. The ratio of the cross sections is ~ 30 and we conservatively scale the errors by $\sqrt{30}$. One should note that at $\sqrt{s} = 350$ GeV the SM background contributions are expected to be smaller, therefore the errors are likely to be lower than our current estimate.

The cross section errors obtained in Sec. 4 did not take into account the uncertainty on the branching ratios. The branching ratios only depend on the available phase space for the decay of the virtual W boson for charginos, given by the chargino-LSP mass difference. With the previously obtained precision on the mass difference, the uncertainty on the BR of the selected decay modes has to be taken into account. In case of the dM1600 scenario, due to the significant uncertainty on the chargino-LSP mass difference, it amounts to 15% for the chargino cross section measurement. For the dM770 scenario it is 2%, comparable to the experimental uncertainty. These additional uncertainties are included in Table 9.

		$\sqrt{s} = 500 \text{ GeV}$			$\sqrt{s} = 350 \text{ \& } 500 \text{ GeV}$	
		input	lower	upper	lower	upper
500/fb	M_1 [TeV]	1.7	~ 0.8	no	~ 0.8	no
	M_2 [TeV]	4.36	~ 1.5	no	~ 1.5	no
	μ [GeV]	165.66	165.2	172.5	165.4	170.2
2/ab	M_1 [TeV]	1.7	~ 1	~ 6	~ 1	~ 6
	M_2 [TeV]	4.36	~ 2.5	~ 8.5	~ 2.5	~ 8.5
	μ [TeV]	165.66	166.2	170.1	166.4	170.0

Table 10: The $1\text{-}\sigma$ allowed ranges for the parameter fit in scenario dM1600. $\tan\beta$ is allowed to vary in the range $[1, 60]$. Determination of the parameters for integrated luminosities of $\int \mathcal{L} dt = 500 \text{ fb}^{-1}$ and 2 ab^{-1} per polarisation configuration is shown. The high-luminosity fit also includes the mass difference between neutralinos, $M_{\tilde{\chi}_2^0} - M_{\tilde{\chi}_1^0}$. For an exact shape of the allowed M_1 - M_2 region, see Fig. 11. The input values, see Eq. (12), are also shown.

Using Minuit [50] we minimize a χ^2 function defined as

$$\chi^2 = \sum_i \left| \frac{\mathcal{O}_i - \bar{\mathcal{O}}_i}{\delta\mathcal{O}_i} \right|^2, \quad (27)$$

where the sum runs over the input observables \mathcal{O}_i , with their corresponding experimental uncertainties $\delta\mathcal{O}_i$. The theoretical values calculated using the fitted MSSM parameters, Eq. (26), are denoted by $\bar{\mathcal{O}}_i$. Note that throughout the simulation the one-loop corrected masses, given by Eqs. (12) and (13), were used. The difference between the tree-level values and the ones obtained from SOFTSUSY is accordingly taken into account in the fit in order to ensure consistency. The variation of the observables, with respect to the central values, on the SUSY parameters, Eq. (26), is in the following calculated at the tree-level. Therefore, the results of the fit are to be understood as the tree-level fit.

It turns out that using the experimental information one cannot constrain $\tan\beta$. This is not very surprising, as with the values of parameters used here, there is very little mixing between higgsinos and gauginos which could have provided some sensitivity to $\tan\beta$. As can be seen from Eq. (7), the chargino and neutralino masses depend only on $\sin 2\beta$ and the sensitivity is rather weak, especially in the large $\tan\beta$ limit. In order to have a converging fit we therefore fix $\tan\beta$ to values in the range $\tan\beta \in [1, 60]$ and fit the remaining three parameters of Eq. (26). In principle, the cross sections depend also on the sneutrino and selectron masses but since the couplings to higgsinos are negligible, no dependence on these parameters enters our observables.

The results of the fit are shown in Tables 10 and 11. The crucial observable for the determination of M_1 and M_2 is the mass difference between the LSP and the light chargino. On the other hand, the cross section measurements improve the determination of the μ parameter. In general, only lower limits can be obtained for M_1 and M_2 . However, it turns out that M_1 and M_2 are strongly correlated and the allowed regions are only narrow strips in the M_1 - M_2 parameter space, see Figs. 11 and 12. The μ parameter can be determined with an accuracy $\sim 2.5 \text{ GeV}$ in dM770 scenario and 6.8 GeV in dM1600 scenario.

The impact of including the cross section measurement at $\sqrt{s} = 350 \text{ GeV}$ only effects the determination of the μ parameter. The low cross section results in a large error, larger by a factor of 5.5 than for $\sqrt{s} = 500 \text{ GeV}$. However, being close to the production threshold and on the quickly-rising slope of the cross section, see Fig. 2, helps to improve the determination of the μ parameter in both scenarios. Taking the cross section at a higher center of mass energy, e.g. $\sqrt{s} = 370 \text{ GeV}$, does not yield a significant improvement. The increased statistics does not compensate the advantage of being close to the production threshold. The masses and mass differences require larger statistics, and therefore, measurements at higher center-of-mass energies (eg. at 500 GeV) are indispensable. Finally, we note that the main driver for the rather high μ error is the

		$\sqrt{s} = 500 \text{ GeV}$			$\sqrt{s} = 350 \text{ \& } 500 \text{ GeV}$	
		input	lower	upper	lower	upper
500/fb	M_1 [TeV]	5.3	~ 2	no	~ 2	no
	M_2 [TeV]	9.51	~ 3	no	~ 3	no
	μ [GeV]	167.22	164.8	167.8	165.2	167.7
2/ab	M_1 [TeV]	5.3	~ 3	no	~ 3	no
	M_2 [TeV]	9.51	~ 7	~ 15	~ 7	~ 15
	μ [GeV]	167.22	165.2	167.4	165.3	167.4

Table 11: The $1\text{-}\sigma$ allowed ranges for the parameter fit in scenario dM770. $\tan\beta$ is allowed to vary in the range $[1, 60]$. Determination of the parameters for integrated luminosities of $\int \mathcal{L} dt = 500 \text{ fb}^{-1}$ and 2 ab^{-1} per polarisation configuration is shown. The high-luminosity fit also includes the mass difference between neutralinos, $M_{\tilde{\chi}_2^0} - M_{\tilde{\chi}_1^0}$. For an exact shape of the allowed M_1 – M_2 region, see Fig. 12. The input values, see Eq. (13), are also shown.

variation of $\tan\beta$, which shifts the central values of the μ parameter by approximately $\sim 1 \text{ GeV}$. This additional dependence will remain regardless of additional measurements and/ or increased accuracy.

In order to better understand the interdependence of the M_1 and M_2 determination, in the left panels of Figs. 11 and 12 we show the $\Delta\chi^2 = 1$ contours in the M_1 – M_2 plane. Different contours correspond to different values of $\tan\beta$. For low values of either M_1 or M_2 the allowed region for the other parameter extends to arbitrary high values and cannot be reliably constrained. Significant variation with $\tan\beta$ is also visible, especially in case of scenario dM770.

Finally, we assess how much of an improvement can be expected after a high-luminosity run with a total of $\int \mathcal{L} dt = 2 \text{ ab}^{-1}$ integrated luminosity at $\sqrt{s} = 500 \text{ GeV}$. We assume that with the increased luminosity the experimental errors would be reduced by a factor of 2. We furthermore include the measurement of the mass difference $M_{\tilde{\chi}_2^0} - M_{\tilde{\chi}_1^0}$, which was not studied here. For the μ parameter the increased luminosity narrows the allowed region by 2–3.5 GeV, as can be seen in the last rows of Tabs. 10 and 11. In case of the M_1 and M_2 parameters the results are shown in the right panels of Figs. 11 and 12 for scenarios dM1600 and dM770, respectively. The neutralino mass difference, $M_{\tilde{\chi}_2^0} - M_{\tilde{\chi}_1^0}$, helps to break the correlation between M_1 and M_2 from the chargino-LSP mass difference, since it exhibits a different dependence on the fundamental parameters, in particular by introducing $\tan\beta$, see Eqs. (9) and (11). Hence, the inclusion of the neutralino mass difference helps to constrain the parameters in the low $\tan\beta$ regime, where the $\tan\beta$ dependence is numerically most significant. This effect would not be achieved by solely increasing the luminosity. The absolute mass measurements, which are obviously improved with higher statistics, have only a very limited impact on determination of the multi-TeV parameters, M_1 and M_2 . The neutralino mass difference, due to its different dependence on the fundamental parameters, provides important additional information to the fit. This underlines the importance of including this measurements in future analyses.

For the scenario dM1600, M_2 is now constrained in the 2.5–8.5 TeV range, while M_1 in the 1–6 TeV range. Also for dM770 one obtains a visible improvement, but still with rather broad ranges of allowed values for M_1 and M_2 . In both cases, the contours for other values of $\tan\beta$ are mostly covered by the high $\tan\beta$ contour and are therefore not shown separately. This result is due to the weak dependence on $\tan\beta$ of the complete neutralino/chargino sector, cf. Section 2.2. Additional measurements, involving e.g. third generation sfermions, would be needed to obtain meaningful constraints on $\tan\beta$.

6 Conclusions

The current LHC results provide strong bounds on constrained supersymmetric models. In particular, the measured Higgs mass of $m_h \approx 126 \text{ GeV}$ results in a tension with the naturalness argument. However, in a

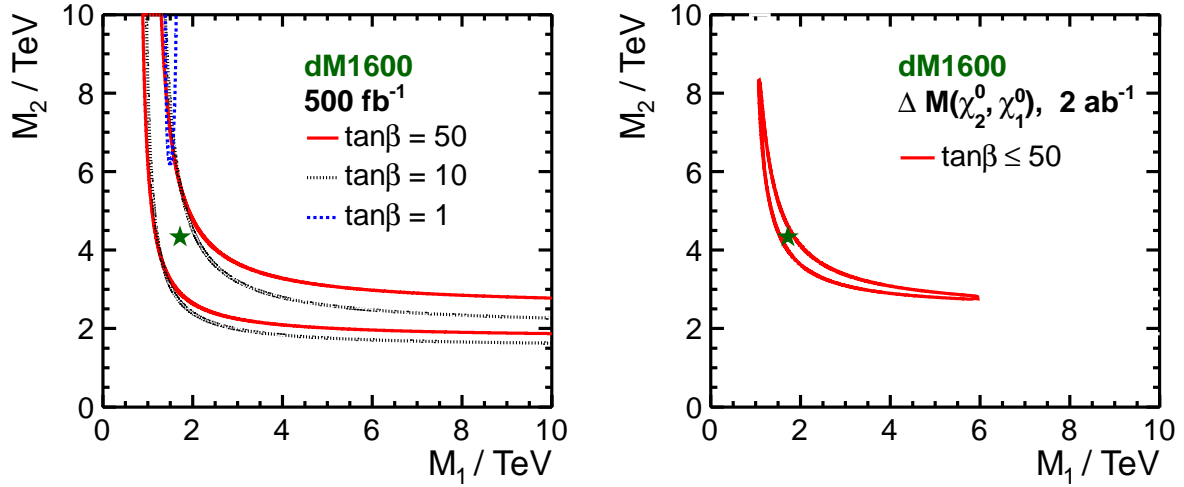


Figure 11: The $\Delta\chi^2 = 1$ contours for determination of M_1 and M_2 in scenario dM1600. The input values of M_1 and M_2 are indicated by the star, and $(\tan\beta)_{\text{true}} = 44$. **Left:** The fit using input values listed in Tab. 9. Black (solid), red (dashed), and blue (dotted) contours are for $\tan\beta = 50, 10$ and 1 , respectively. **Right:** Projected fit results after a high luminosity run, $\int \mathcal{L} dt = 2 \text{ ab}^{-1}$, with experimental errors improved by factor 2 and with the mass difference $M_{\tilde{\chi}_2^0} - M_{\tilde{\chi}_1^0}$ measurement included. Black and red contours are for $\tan\beta = 50$ and 1 , respectively.

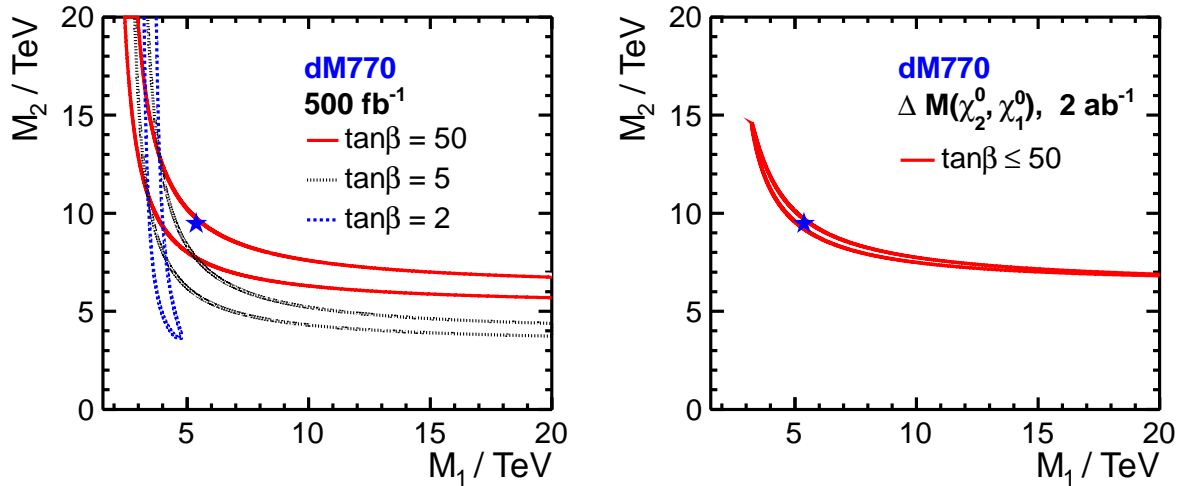


Figure 12: The $\Delta\chi^2 = 1$ contours for determination of M_1 and M_2 in scenario dM770. The input values of M_1 and M_2 are indicated by the star, and $(\tan\beta)_{\text{true}} = 48$. **Left:** The fit using input values listed in Tab. 9. Black (solid), red (dashed) and blue (dotted) contours are for $\tan\beta = 50, 5$ and 2 , respectively. **Right:** Projected fit results after a high luminosity run, $\int \mathcal{L} dt = 2 \text{ ab}^{-1}$, with experimental errors improved by factor 2 and with the mass difference $M_{\tilde{\chi}_2^0} - M_{\tilde{\chi}_1^0}$ measurement included for fixed $\tan\beta = 50$. Other values of $\tan\beta$ do not further extend allowed regions and the respective contours are inside.

specific class of scenarios, a rather light μ -parameter in conjunction with a heavy coloured SUSY and gaugino spectrum can be well-motivated from a model-building perspective. Such scenarios result in a strong mass degeneracy between the higgsino-like light neutralinos and chargino, $\tilde{\chi}_{1,2}^0$ and $\tilde{\chi}_1^\pm$. Due to the close mass degeneracies, three-body decays as well as radiative decays are dominant leading to extremely challenging experimental signatures.

We therefore performed a simulation for the ILD detector in two different benchmark scenarios at a center-of-mass energy of 500 GeV, assuming an integrated luminosity of 500fb^{-1} each for the two relevant polarisation configurations, which corresponds to about 4 years of ILC operation with design accelerator parameters. In both scenarios, the μ -parameter is around 160 GeV, but the SUSY gaugino parameters M_1 and M_2 are in the multi-TeV region. This leads to mass differences between the light chargino and lightest neutralino of about 1.6 GeV and 770 MeV in the two benchmarks, respectively. These small mass differences lead to a very soft transverse momentum spectrum of the visible decay products, which ends at 5 and 2 GeV for the two different scenarios. In this kinematic regime, background processes from collisions involving the photon-component of the electron and positron beams are about 5 orders of magnitude larger than the signal. However, they can be heavily reduced by requiring the signal to be accompanied by a hard ISR photon.

In conjunction with the known beam energy, the $\tilde{\chi}_1^\pm$ and $\tilde{\chi}_2^0$ masses can be determined with a precision between 1–2% from the reduced center-of-mass energy of the system recoiling against the photon. The mass difference with the LSP can be reconstructed by boosting the visible decay products into the rest frame of the produced gaugino pair. This allows one to clearly resolve the lightest and next-to-lightest SUSY particles and to determine the corresponding mass difference in the two benchmarks with a relative accuracy of 17% and 5%, respectively. Finally, the polarised cross sections can be measured with an accuracy of a few percent. The polarisation dependency gives a clear indication of the higgsino nature.

In the last step of this analysis, we investigated whether the relevant SUSY parameters, namely the gaugino mass parameters M_1 and M_2 , the μ -parameter and $\tan\beta$ can be extracted from the experimental observables. We find that the μ -parameter can be determined to $\pm 4\%$. In case of M_1 and M_2 , lower bounds in the multi-TeV range can be set, which depend slightly on the value of $\tan\beta$, which cannot be fixed from these measurements alone. If the measurement uncertainties could be reduced by a factor of 2, e.g. by increasing the luminosity, or by improving the analysis, and if a measurement of the mass difference between $\tilde{\chi}_2^0$ and $\tilde{\chi}_1^0$ could be performed with the same precision as for the chargino mass difference, the constraints on M_1 and M_2 would be significantly more restrictive and become less dependent on $\tan\beta$.

Our results demonstrate the strong physics potential of the ILC. High precision is mandatory in order to resolve even such challenging scenarios that could be easily missed at the LHC. The high experimental precision of an ILC of course has to be balanced by suitable higher order predictions implemented in a Monte Carlo generator on a fully differential level. First steps in this direction have already been presented in [35, 36], but a full study including decays would be needed and is in the line of future work.

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References

- [1] G. Aad *et al.* [ATLAS Collaboration], Phys. Lett. B **716** (2012) 1 [arXiv:1207.7214 [hep-ex]].
- [2] S. Chatrchyan *et al.* [CMS Collaboration], Phys. Lett. B **716** (2012) 30 [arXiv:1207.7235 [hep-ex]].
- [3] G. Aad *et al.* [ATLAS Collaboration], Eur. Phys. J. C **73** (2013) 2362 [arXiv:1212.6149 [hep-ex]]; ATLAS-CONF-2013-054; ATLAS-CONF-2013-047; ATLAS-CONF-2013-037; ATLAS-CONF-2013-026; ATLAS-CONF-2013-024; ATLAS-CONF-2013-007; ATLAS-CONF-2012-144; ATLAS-CONF-2012-140.

- [4] S. Chatrchyan *et al.* [CMS Collaboration], Phys. Rev. D **87** (2013) 072001 [arXiv:1301.0916 [hep-ex]]; Eur. Phys. J. C **73** (2013) 2404 [arXiv:1212.6428 [hep-ex]]; JHEP **1303** (2013) 037 [arXiv:1212.6194 [hep-ex]], JHEP **1303** (2013) 111 [arXiv:1211.4784 [hep-ex]]; arXiv:1305.2390 [hep-ex]; arXiv:1303.2985 [hep-ex]; arXiv:1301.3792 [hep-ex]; arXiv:1301.2175 [hep-ex]; arXiv:1212.6961 [hep-ex]; CMS-PAS-SUS-13-011; CMS-PAS-SUS-13-008; CMS-PAS-SUS-13-007.
- [5] [ATLAS Collaboration], ATLAS-CONF-2013-035 and ATLAS-CONF-2013-049.
- [6] J. Beringer *et al.* [Particle Data Group Collaboration], Phys. Rev. D **86** (2012) 010001; A. Heister *et al.* [ALEPH Collaboration], Phys. Lett. B **533** (2002) 223 [hep-ex/0203020]; J. Abdallah *et al.* [DELPHI Collaboration], Eur. Phys. J. C **31** (2003) 421 [hep-ex/0311019]; M. Acciarri *et al.* [L3 Collaboration], Phys. Lett. B **482** (2000) 31 [hep-ex/0002043]; G. Abbiendi *et al.* [OPAL Collaboration], Eur. Phys. J. C **35** (2004) 1 [hep-ex/0401026].
- [7] M. Papucci, J. T. Ruderman and A. Weiler, JHEP **1209** (2012) 035 [arXiv:1110.6926 [hep-ph]].
- [8] R. Allahverdi, B. Dutta and K. Sinha, Phys. Rev. D **86** (2012) 095016 [arXiv:1208.0115 [hep-ph]].
- [9] H. Baer, V. Barger and P. Huang, JHEP **1111** (2011) 031 [arXiv:1107.5581 [hep-ph]].
- [10] H. Baer, V. Barger, P. Huang and X. Tata, JHEP **1205** (2012) 109 [arXiv:1203.5539 [hep-ph]].
- [11] F. Brümmer and W. Buchmüller, JHEP **1107** (2011) 010 [arXiv:1105.0802 [hep-ph]].
- [12] F. Brümmer and W. Buchmüller, JHEP **1205** (2012) 006 [arXiv:1201.4338 [hep-ph]].
- [13] H. E. Haber and G. L. Kane, Phys. Rept. **117** (1985) 75.
- [14] A. Bartl, H. Fraas, W. Majerotto and N. Oshimo, Phys. Rev. D **40** (1989) 1594.; M. M. El Kheishen, A. A. Aboshousha and A. A. Shafik, Phys. Rev. D **45** (1992) 4345.; M. Guchait, Z. Phys. C **57** (1993) 157 [Erratum-ibid. C **61** (1994) 178]; G. J. Gounaris, C. Le Mouel and P. I. Porfyriadis, Phys. Rev. D **65** (2002) 035002 [hep-ph/0107249].
- [15] S. Y. Choi, J. Kalinowski, G. A. Moortgat-Pick and P. M. Zerwas, Eur. Phys. J. C **22** (2001) 563 [Addendum-ibid. C **23** (2002) 769] [hep-ph/0108117].
- [16] J. F. Gunion and H. E. Haber, Phys. Rev. D **37** (1988) 2515.
- [17] S. Ambrosanio and B. Mele, Phys. Rev. D **55** (1997) 1399 [Erratum-ibid. D **56** (1997) 3157] [hep-ph/9609212].
- [18] S. Y. Choi, M. Guchait, J. Kalinowski and P. M. Zerwas, Phys. Lett. B **479** (2000) 235 [hep-ph/0001175].
- [19] A. Bharucha, A. Fowler, G. Moortgat-Pick and G. Weiglein, arXiv:1211.3134 [hep-ph].
- [20] A. Bharucha, J. Kalinowski, G. Moortgat-Pick, K. Rolbiecki and G. Weiglein, arXiv:1211.3745 [hep-ph].
- [21] G. A. Moortgat-Pick and H. Fraas, Phys. Rev. D **59** (1999) 015016 [hep-ph/9708481]; G. A. Moortgat-Pick, H. Fraas, A. Bartl and W. Majerotto, Eur. Phys. J. C **7** (1999) 113 [hep-ph/9804306]; G. A. Moortgat-Pick, H. Fraas, A. Bartl and W. Majerotto, Eur. Phys. J. C **9** (1999) 521 [Erratum-ibid. C **9** (1999) 549] [hep-ph/9903220].
- [22] H. E. Haber and D. Wyler, Nucl. Phys. B **323** (1989) 267.
- [23] K. Yokoya and P. Chen, Lect. Notes Phys. **400** (1992) 415.
- [24] P. Chen, Phys. Rev. D **46** (1992) 1186.
- [25] B. C. Allanach, Comput. Phys. Commun. **143** (2002) 305 [hep-ph/0104145].
- [26] T. Fritzsche and W. Hollik, Eur. Phys. J. C **24** (2002) 619 [hep-ph/0203159]; T. Fritzsche, "Berechnung von Observablen zur supersymmetrischen Teilchenerzeugung an Hochenergie-Collidern unter Einschluss hoeherer Ordnungen", PhD Thesis, Cuvillier Verlag, Göttingen 2005, ISBN 3-86537-577-4.
- [27] D. M. Pierce, J. A. Bagger, K. T. Matchev and R. -j. Zhang, Nucl. Phys. B **491** (1997) 3 [hep-ph/9606211].
- [28] W. Kilian, T. Ohl and J. Reuter, Eur. Phys. J. C **71** (2011) 1742 [arXiv:0708.4233 [hep-ph]].
- [29] T. Behnke, J. E. Brau, P. N. Burrows, J. Fuster, M. Peskin, M. Stanitzki, Y. Sugimoto and S. Yamada *et al.*, "The International Linear Collider Technical Design Report - Volume 4: Detectors," Volume 4, (2013), 32-36, arXiv:1306.6329 [physics.ins-det].
- [30] M. Skrzypek and S. Jadach, Z. Phys. C **49** (1991) 577.
- [31] C. Adolphsen, M. Barone, B. Barish, K. Büsser, P. Burrows, J. Carwardine, J. Clark and Hln. M. Durand *et al.*, "The International Linear Collider Technical Design Report - Volume 3.II: Accelerator Baseline Design,"(2013), 8-10, arXiv:1306.6328 [physics.acc-ph].
- [32] C. Hensel, "Search for nearly mass degenerate charginos and neutralinos in e^+e^- collisions," DESY-THESIS-2002-047. <http://www-library.desy.de/cgi-bin/showprep.pl?desy-thesis-02-047>
- [33] T. Fritzsche and W. Hollik, Nucl. Phys. Proc. Suppl. **135** (2004) 102 [hep-ph/0407095].
- [34] W. Oller, H. Eberl and W. Majerotto, Phys. Rev. D **71** (2005) 115002 [arXiv:hep-ph/0504109].
- [35] W. Kilian, J. Reuter and T. Robens, Eur. Phys. J. C **48** (2006) 389 [hep-ph/0607127].
- [36] T. Robens, "Event Generation for Next to Leading Order Chargino Production at the International Linear Collider," hep-ph/0610401.
- [37] K. Desch, J. Kalinowski, G. A. Moortgat-Pick, M. M. Nojiri and G. Polesello, JHEP **0402** (2004) 035 [hep-ph/0312069].

- [38] G. Moortgat-Pick, T. Abe, G. Alexander, B. Ananthanarayan, A. A. Babich, V. Bharadwaj, D. Barber and A. Bartl *et al.*, *Phys. Rept.* **460** (2008) 131 [hep-ph/0507011].
- [39] G. Moortgat-Pick, I. Fleck, S. Riemann, F. Simon, O. S. Adeyemi, G. Alexander, M. S. Amjad and V. V. Andreev *et al.*, “Helmholtz Alliance Linear Collider Forum : Proceedings of the Workshops Hamburg, Munich, Hamburg 2010-2012, Germany,” DESY 12-123H.
- [40] D. Grellscheid and P. Richardson, arXiv:0710.1951 [hep-ph].
- [41] M. Bahr, S. Gieseke, M. A. Gigg, D. Grellscheid, K. Hamilton, O. Latunde-Dada, S. Platzer and P. Richardson *et al.*, *Eur. Phys. J. C* **58** (2008) 639 [arXiv:0803.0883 [hep-ph]].
- [42] T. Sjostrand, S. Mrenna and P. Z. Skands, *JHEP* **0605** (2006) 026 [hep-ph/0603175].
- [43] T. Behnke, J. E. Brau, P. N. Burrows, J. Fuster, M. Peskin, M. Stanitzki, Y. Sugimoto and S. Yamada *et al.*, “The International Linear Collider Technical Design Report - Volume 4: Detectors,” Volume 4, (2013), 181-314, arXiv:1306.6329 [physics.ins-det].
- [44] C. Bartels, M. Berggren and J. List, *Eur. Phys. J. C* **72** (2012) 2213 [arXiv:1206.6639 [hep-ex]].
- [45] S. Agostinelli *et al.* [GEANT4 Collaboration], *Nucl. Instrum. Meth. A* **506** (2003) 250.
- [46] T. Abe *et al.* [ILD Concept Group - Linear Collider Collaboration], “The International Large Detector: Letter of Intent,” arXiv:1006.3396 [hep-ex].
- [47] M. Berggren, “SGV 3.0 - a fast detector simulation,” arXiv:1203.0217 [physics.ins-det].
- [48] D. Schulte, “Study of Electromagnetic and Hadronic Background in the Interaction Region of the TESLA Collider,” TESLA-97-08.
- [49] P. Mora de Freitas and H. Videau, “Detector simulation with MOKKA / GEANT4: Present and future,” LC-TOOL-2003-010.
- [50] F. James and M. Roos, *Comput. Phys. Commun.* **10** (1975) 343.