# Quantization conditions and functional equations in ABJ(M) theories 

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#### Abstract

The partition function of $\operatorname{ABJ}(\mathrm{M})$ theories on the three-sphere can be regarded as the canonical partition function of an ideal Fermi gas with a non-trivial Hamiltonian. We propose an exact expression for the spectral determinant of this Hamiltonian, which generalizes recent results obtained in the maximally supersymmetric case. As a consequence, we find an exact WKB quantization condition determining the spectrum which is in agreement with numerical results. In addition, we investigate the factorization properties and functional equations for our conjectured spectral determinants. These functional equations relate the spectral determinants of ABJ theories with consecutive ranks of gauge groups but the same Chern-Simons coupling.


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## 1. Introduction

In the last years, there has been a lot of progress in understanding $\operatorname{ABJ}(\mathrm{M})$ theory $[1,2]$. In [3] the partition function of $\mathrm{ABJ}(\mathrm{M})$ theory on the three-sphere was reduced to a matrix integral which turned out to be closely related to topological strings on local $\mathbb{P}^{1} \times \mathbb{P}^{1}$ [4]. In [5] the connection with topological strings was used to compute recursively the full 't Hooft $1 / N$ expansion, which by the AdS/CFT correspondence corresponds to the genus expansion of a dual type IIA superstring theory. In order to understand the M-theory lifting of this result, one has to study ABJM theory in a different regime, usually called the M-theory regime or M-theory expansion, in which $N$ is large but the coupling constant is fixed. The study of the matrix models computing partition functions of Chern-Simons-matter theories in the M-theory regime was initiated in [6], where the strict large $N$ limit was solved for a large class of theories.

In [7], a different approach was proposed to study the M-theory regime of ABJM theory and related models. In this approach, the partition function of $\operatorname{ABJ}(\mathrm{M})$ is interpreted as the partition function of an ideal Fermi gas. The M-theory limit corresponds to the thermodynamic limit of this gas, and the coupling constant of ABJM theory becomes Planck's constant. The Fermi gas formulation of ABJM theory has been intensively studied in $[7,8,9,10,11,12,13]$, leading to an exact expression for the the partition function of $\mathrm{ABJ}(\mathrm{M})$ theory which resums the 't Hooft expansion and includes as well non-perturbative, large $N$ instanton corrections [14].

An important aspect of the Fermi gas approach is that, since we are dealing with an ideal gas, all the physics of the problem is encoded in the spectrum of the one-particle Hamiltonian. Therefore one should be able to reproduce the results of [13] by studying the spectral problem associated to the Fermi gas of $\mathrm{ABJ}(\mathrm{M})$ theory. Conversely, the exact expression for the partition function should encode all the information about the spectrum of the Fermi gas. In [15, 16], a WKB quantization condition for $\mathrm{ABJ}(\mathrm{M})$ theory has been proposed by studying the relation between the spectrum of the Hamiltonian and the thermodynamics of the Fermi gas. This condition turns out to be exact in the cases in which the theory has maximal supersymmetry [17], but it needs additional corrections in the general case, as it was recently pointed out in [18] (see also [19]) by a detailed numerical analysis.

One of the key results of [17] is that, in the maximally supersymmetric cases, one can write an explicit expression for the grand canonical partition function of the Fermi gas, which is nothing but the spectral determinant of the Hamiltonian. This expression involves a Jacobi theta function, and the spectrum can be read from the vanishing locus of this theta function. In this paper we generalize the results of $[17]$ to $\operatorname{ABJ}(\mathrm{M})$ theories with $\mathcal{N}=6$ supersymmetry. We write a general formula for the spectral determinant of these theories, which involves now a generalization of the theta function. In particular, we derive an exact WKB quantization condition for the spectrum. The quantization condition proposed in $[15,16]$ is only an approximation to the exact quantization condition, and in general it receives corrections that we can compute analytically in this paper. Our general result explains why the quantization conditions of $[15,16]$ are valid in the maximally supersymmetric cases. It reproduces the corrections found numerically in the case of ABJM theory in [18], and we also test these corrections in detail against explicit calculations of the spectrum in both ABJM theory and ABJ theory.

As it was emphasized in the companion paper [20], where we studied the implications of these ideas for topological string theory, the quantization condition is obtained as a corollary of a stronger result, namely a conjectural exact expression for the spectral determinant. This expression was tested in detail in [17] in the maximally supersymmetric case, where it was shown that it reproduces the values for the partition functions calculated in [8, 9, 22, 23]. In the general case with $\mathcal{N}=6$ supersymmetry our conjecture for the spectral determinant is more difficult to verify, since this involves a resummation of the Gopakumar-Vafa expansion of the topological string free energy [21]. However, we are able to perform this resummation in one special case (ABJM theory with $k=4$ ), and we obtain a generating functional for the partition functions of this theory in full agreement with existing calculations [10].

In exactly solvable cases, spectral determinants enjoy very interesting properties. They can be factorized according to the parity of the eigenfunctions, and they satisfy functional equations (see for example [24, 25]). In this paper, we initiate the study of such properties in the spectral problem of $\operatorname{ABJ}(\mathrm{M})$ theory. We find for example an explicit factorization of the spectral determinant in the maximally supersymmetric case $k=1$, as well as conjectural functional equations akin to those found in $[24,25]$ in Quantum Mechanics. These functional equations relate the spectral determinants of ABJ theories with consecutive ranks of gauge groups. In particular, if the Chern-Simons levels are odd, these equations determine all the ABJ spectral determinants from the ABJM ones via the Seiberg-like duality of ABJ theory [2].

This paper is organized as follows. In section 2 we review some general aspects of the Fermi gas formalism. In section 3 we introduce the spectral determinant and the generalized theta function associated to the $\mathrm{ABJ}(\mathrm{M})$ grand potential, and we deduce the exact quantization condition for the energy levels by looking at the zeros of this generalized theta function. We also give strong numerical evidence in support of our computations. In section 4 we study an example
with $\mathcal{N}=6$ supersymmetry and we show that the full genus expansion can be resummed into an explicit function on the moduli space. In section 5 we discuss the factorization of the spectral determinant according to the parity of the energy levels and in section 6 we give evidence for some functional identities. In section 7 we draw some conclusions. There are also three appendices. In appendix A and B we give some details for computations appearing in section 4 and in section 5. In appendix $C$ we explain the numerical technique used to compute the spectrum.

## 2. The Fermi gas approach to ABJ(M) Theory

The $\operatorname{ABJ}(\mathrm{M})$ theory [1, 2] is an $\mathcal{N}=6$ superconformal Chern-Simons-matter theory with gauge group $U\left(N_{1}\right)_{k} \times U\left(N_{2}\right)_{-k}$. It consists in two Chern-Simons nodes, with couplings $k$ and $-k$, respectively, together with four hypermultiplets in the bifundamental representation. By using localization techniques it is possible to reduce the $\operatorname{ABJ}(\mathrm{M})$ partition function on $\mathbb{S}^{3}$ to the following matrix integral [3]:

$$
\begin{align*}
& Z\left(N_{1}, N_{2}, k\right) \\
& =\frac{\mathrm{i}^{-\frac{1}{2}\left(N_{1}^{2}-N_{2}^{2}\right)}}{N_{1}!N_{2}!} \int \prod_{i=1}^{N_{1}} \frac{\mathrm{~d} \mu_{i}}{2 \pi} \prod_{j=1}^{N_{2}} \frac{\mathrm{~d} \nu_{j}}{2 \pi} \frac{\prod_{i<j}\left(2 \sinh \left(\frac{\mu_{i}-\mu_{j}}{2}\right)\right)^{2}\left(2 \sinh \left(\frac{\nu_{i}-\nu_{j}}{2}\right)\right)^{2}}{\prod_{i, j}\left(2 \cosh \left(\frac{\mu_{i}-\nu_{j}}{2}\right)\right)^{2}} \mathrm{e}^{-\frac{\mathrm{i} k}{4 \pi}\left(\sum_{i} \mu_{i}^{2}-\sum_{j} \nu_{j}^{2}\right) .} \tag{2.1}
\end{align*}
$$

When $N_{1}=N_{2}=N$ the above matrix integral can be also written as $[26,7]$

$$
\begin{equation*}
Z(N, k)=\frac{1}{N!} \int \prod_{i=1}^{N} \frac{\mathrm{~d} x_{i}}{4 \pi k} \frac{1}{2 \cosh \frac{x_{i}}{2}} \prod_{i<j}\left(\tanh \left(\frac{x_{i}-x_{j}}{2 k}\right)\right)^{2} \tag{2.2}
\end{equation*}
$$

These matrix integrals can be studied in the conventional 't Hooft expansion [4, 5]. In [7] it was pointed out that, to fully understand the non-perturbative effects, one has to go beyond the 't Hooft $1 / N$ expansion and study the M-theory regime of (2.1). In this regime, the ranks of the gauge groups are large but the coupling $k$ is fixed. To study this regime it is convenient to use the Fermi gas approach [7] in which we rewrite the matrix integral as the canonical partition function of a one-dimensional ideal Fermi gas. In this approach, the Chern-Simons coupling $k$ plays the role of the Planck constant:

$$
\begin{equation*}
\hbar=2 \pi k \tag{2.3}
\end{equation*}
$$

The Fermi gas formulation of the ABJ matrix integral was proposed in [27, 28] where (2.1) was written as

$$
\begin{equation*}
Z(N, N+M, k)=\mathrm{e}^{\mathrm{i} \vartheta(N, M, k)} Z_{\mathrm{CS}}(M, k) \hat{Z}(N, k ; M) \tag{2.4}
\end{equation*}
$$

and we used the notation

$$
\begin{equation*}
N=N_{1}, \quad M=N_{2}-N_{1} \tag{2.5}
\end{equation*}
$$

In the following we will suppose that

$$
\begin{equation*}
k \geq 0, \quad M \geq 0 \tag{2.6}
\end{equation*}
$$

The phase factor appearing in (2.4) is given by

$$
\begin{equation*}
\mathrm{e}^{\mathrm{i} \vartheta(N, M, k)}=\mathrm{i}^{N M} \mathrm{e}^{-\frac{\mathrm{i} \pi}{6 k} M\left(M^{2}-1\right)}, \tag{2.7}
\end{equation*}
$$

and $Z_{\mathrm{CS}}(M, k)$ is the $U(M)$ Chern-Simon partition function on $\mathbb{S}^{3}[29]$ :

$$
\begin{equation*}
Z_{\mathrm{CS}}(M, k)=k^{-\frac{M}{2}} \prod_{s=1}^{M-1}\left(2 \sin \frac{\pi s}{k}\right)^{M-s} . \tag{2.8}
\end{equation*}
$$

The factor $\hat{Z}(N, k ; M)$ has the form,

$$
\begin{equation*}
\hat{Z}(N, k ; M)=\frac{1}{N!} \sum_{\sigma \in S_{N}} \int \mathrm{~d}^{N} x \prod_{i=1}^{N} \rho\left(x_{i}, x_{\sigma(i)}\right) . \tag{2.9}
\end{equation*}
$$

The function $\rho\left(x_{1}, x_{2}\right)$ is given as

$$
\begin{equation*}
\rho\left(x_{1}, x_{2}\right)=\frac{1}{2 \pi k} \frac{V_{M}^{1 / 2}\left(x_{1}\right) V_{M}^{1 / 2}\left(x_{2}\right)}{2 \cosh \left(\frac{x_{1}-x_{2}}{2 k}\right)}, \tag{2.10}
\end{equation*}
$$

where

$$
\begin{equation*}
V_{M}(x)=\frac{1}{\mathrm{e}^{x / 2}+(-1)^{M} \mathrm{e}^{-x / 2}} \prod_{s=-\frac{M-1}{2}}^{\frac{M-1}{2}} \tanh \frac{x+2 \pi \mathrm{i} s}{2 k} . \tag{2.11}
\end{equation*}
$$

This function can be interpreted as a canonical density matrix,

$$
\begin{equation*}
\left\langle x_{1}\right| \hat{\rho}\left|x_{2}\right\rangle=\rho\left(x_{1}, x_{2}\right), \quad \hat{\rho}=\mathrm{e}^{-\hat{H}}, \tag{2.12}
\end{equation*}
$$

where $\hat{H}$ is the one-particle Hamiltonian, and (2.9) is then the canonical partition function of an ideal Fermi gas with Hamiltonian $\hat{H}$. It is useful to write

$$
\begin{equation*}
V_{M}(x)=\frac{1}{\mathrm{e}^{x / 2}+(-1)^{M} \mathrm{e}^{-x / 2}}\left(\tanh \frac{x}{2 k}\right)^{p_{M}} \prod_{s=\frac{1+p_{M}}{2}}^{\frac{M-1}{2}} \tanh \frac{x+2 \pi \mathrm{i} s}{2 k} \tanh \frac{x-2 \pi \mathrm{i} s}{2 k}, \tag{2.13}
\end{equation*}
$$

where

$$
p_{M}= \begin{cases}1 & \text { if } M \text { is odd }  \tag{2.14}\\ 0 & \text { if } M \text { is even. }\end{cases}
$$

It is now manifest that $V_{M}(x)$ is a real and positive functions: indeed, we have

$$
\begin{equation*}
\tanh \frac{x+2 \pi \mathrm{i} s}{2 k} \tanh \frac{x-2 \pi \mathrm{i} s}{2 k}=\frac{\cosh \left(\frac{x}{k}\right)-\cos \left(\frac{2 \pi s}{k}\right)}{\cosh \left(\frac{x}{k}\right)+\cos \left(\frac{2 \pi s}{k}\right)} \geq 0 . \tag{2.15}
\end{equation*}
$$

When $M=0$, the ABJ partition function becomes the partition function of ABJM theory in (2.2):

$$
\begin{equation*}
Z(N, N, k)=\hat{Z}(N, k ; 0)=Z(N, k) . \tag{2.16}
\end{equation*}
$$

The density matrix (2.10) defines a positive-definite, Hilbert-Schmidt integral kernel. The spectrum $E_{n}$ of the associated Hamiltonian is determined by the integral equation

$$
\begin{equation*}
\int_{-\infty}^{\infty} \rho\left(x_{1}, x_{2}\right) \phi_{n}\left(x_{2}\right) \mathrm{d} x_{2}=\mathrm{e}^{-E_{n}} \phi_{n}\left(x_{1}\right), \quad n \geq 0, \tag{2.17}
\end{equation*}
$$



Figure 1: The standard Airy contour $\mathcal{C}$ used to compute the canonical partition function from the modified grand potential.
where we have ordered the eigenvalues as

$$
\begin{equation*}
E_{0}<E_{1}<E_{2}<\cdots \tag{2.18}
\end{equation*}
$$

As it is well-known, ideal quantum gases are better studied in the grand canonical ensemble. The grand canonical partition function is defined by

$$
\begin{equation*}
\Xi(\kappa, k, M)=\operatorname{det}(1+\kappa \hat{\rho})=\prod_{n=0}^{\infty}\left(1+\kappa \mathrm{e}^{-E_{n}}\right) \tag{2.19}
\end{equation*}
$$

where

$$
\begin{equation*}
\kappa=\mathrm{e}^{\mu} \tag{2.20}
\end{equation*}
$$

is the fugacity. We will use the notation $\Xi(\kappa, k, M)$ and $\Xi(\mu, k, M)$ interchangably. When $M=0$ we will write

$$
\begin{equation*}
\Xi(\kappa, k)=\Xi(\kappa, k, 0) \text {. } \tag{2.21}
\end{equation*}
$$

The grand canonical partition function can be also interpreted as the spectral determinant (or Fredholm determinant) of the operator $\hat{\rho}$. Since this operator is positive-definite and HilbertSchmidt, it is of trace class and therefore its spectral determinant is well-defined ${ }^{1}$. It has two important properties that we will use later on. The first one is that $\Xi(\kappa, k, M)$ is an entire function of the the fugacity [31], and the second one is that one can read off the physical energy spectrum by looking at the zeros of (2.19). Indeed it is easy to see from the definition (2.19) that

$$
\begin{equation*}
\Xi(E+\mathrm{i} \pi, k, M) \tag{2.22}
\end{equation*}
$$

has simple zeros for

$$
\begin{equation*}
E=E_{n} . \tag{2.23}
\end{equation*}
$$

[^0]The spectral determinant can be also regarded as a generating function for the partition functions $\hat{Z}(N, k, M)$ :

$$
\begin{equation*}
\Xi(\kappa, k, M)=1+\sum_{N \geq 1} \hat{Z}(N, k, M) \kappa^{N} . \tag{2.24}
\end{equation*}
$$

We define the standard grand potential as

$$
\begin{equation*}
\mathcal{J}(\mu, k, M)=\log \Xi(\kappa, k, M) \tag{2.25}
\end{equation*}
$$

Its power series expansion around $\kappa=0$

$$
\begin{equation*}
\mathcal{J}(\mu, k, M)=\sum_{\ell \geq 1} \frac{(-1)^{\ell-1}}{\ell} Z_{\ell} \kappa^{\ell} \tag{2.26}
\end{equation*}
$$

involves the spectral traces of the canonical density matrix,

$$
\begin{equation*}
Z_{\ell}=\operatorname{Tr} \hat{\rho}^{\ell}=\sum_{n \geq 0} \mathrm{e}^{-\ell E_{n}} \tag{2.27}
\end{equation*}
$$

In the context of $\operatorname{ABJ}(\mathrm{M})$ theory it is convenient to use the modified grand potential $J(\mu, k, M)$, which was introduced in [10]. It is related to the partition function by

$$
\begin{equation*}
\hat{Z}(N, k, M)=\int_{\mathcal{C}} \frac{\mathrm{d} \mu}{2 \pi \mathrm{i}} \mathrm{e}^{J(\mu, k, M)-N \mu} \tag{2.28}
\end{equation*}
$$

where $\mathcal{C}$ is the standard Airy contour shown in Figure 1. The standard and the modified grand potential are related via

$$
\begin{equation*}
\mathrm{e}^{\mathcal{J}(\mu, k, M)}=\sum_{n \in \mathbb{Z}} \mathrm{e}^{J(\mu+2 \pi \mathrm{i} n, k, M)} \tag{2.29}
\end{equation*}
$$

The modified grand potential of ABJM theory was determined in a series of works $[7,8,10,11$, $12,13]$, and it can be written down in terms of the standard and refined topological strings on local $\mathbb{P}^{1} \times \mathbb{P}^{1}$. This result was extended to ABJ theory in $[23,22]$. One has the following result:

$$
\begin{equation*}
J(\mu, k, M)=J^{(\mathrm{p})}\left(\mu_{\mathrm{eff}}, k, M\right)+J^{\mathrm{WS}}\left(\mu_{\mathrm{eff}}, k, M\right)+\mu_{\mathrm{eff}} \widetilde{J}_{b}\left(\mu_{\mathrm{eff}}, k, M\right)+\widetilde{J}_{c}\left(\mu_{\mathrm{eff}}, k, M\right) \tag{2.30}
\end{equation*}
$$

The perturbative piece $J^{(\mathrm{p})}$ is a cubic polynomial in $\mu$ :

$$
\begin{equation*}
J^{(\mathrm{p})}(\mu, k, M)=\frac{C(k)}{3} \mu^{3}+B(k, M) \mu+A(k, M) \tag{2.31}
\end{equation*}
$$

where

$$
\begin{equation*}
C(k)=\frac{2}{\pi^{2} k}, \quad B(k, M)=\frac{1}{3 k}-\frac{k}{12}+\frac{k}{2}\left(\frac{M}{k}-\frac{1}{2}\right)^{2} \tag{2.32}
\end{equation*}
$$

The constant term is given by

$$
\begin{equation*}
A(k, M)=-\log \left|Z_{\mathrm{CS}}(M, k)\right|+\frac{2 \zeta(3)}{\pi^{2} k}\left(1-\frac{k^{3}}{16}\right)+\frac{k^{2}}{\pi^{2}} \int_{0}^{\infty} \mathrm{d} x \frac{x}{\mathrm{e}^{k x}-1} \log \left(1-\mathrm{e}^{-2 x}\right) \tag{2.33}
\end{equation*}
$$

where $Z_{\mathrm{CS}}(M, k)$ is the same as in (2.8). In particular for $M=0$ we have

$$
\begin{equation*}
Z_{\mathrm{CS}}(0, k)=1 \tag{2.34}
\end{equation*}
$$

and we recover the constant map contribution of ABJM [32, 33]. The exact values of this constant map contribution for arbitrary integral $k$ are found in [33]. The effective chemical potential $\mu_{\text {eff }}$ was introduced in [12] to take into account bound states of worldsheet instantons and membrane instantons. It is given by

$$
\begin{equation*}
\mu_{\mathrm{eff}}=\mu-\frac{1}{2} \sum_{\ell=1}^{\infty}(-1)^{M \ell} \hat{a}_{\ell}(k) \mathrm{e}^{-2 \ell \mu} \tag{2.35}
\end{equation*}
$$

In [13], it was shown that the coefficients $\hat{a}_{\ell}(k)$ are the so-called quantum mirror map of local $\mathbb{P}^{1} \times \mathbb{P}^{1}$ introduced in [34]. For the first few terms we have

$$
\begin{align*}
& \hat{a}_{1}(k)=2\left(q^{1 / 2}+q^{-1 / 2}\right) \\
& \hat{a}_{2}(k)=5\left(q+q^{-1}\right)+8  \tag{2.36}\\
& \hat{a}_{3}(k)=2\left(q^{5 / 2}+q^{-5 / 2}\right)+\frac{62}{3}\left(q^{3 / 2}+q^{-3 / 2}\right)+44\left(q^{1 / 2}+q^{-1 / 2}\right)
\end{align*}
$$

and we denoted

$$
\begin{equation*}
q=\mathrm{e}^{\mathrm{i} \pi k} \tag{2.37}
\end{equation*}
$$

When $k$ is an integer the effective potential can be written in closed form [23]

$$
\mu_{\text {eff }}= \begin{cases}\mu-2(-1)^{\frac{k}{2}-M} \mathrm{e}^{-2 \mu}{ }_{4} F_{3}\left(1,1, \frac{3}{2}, \frac{3}{2}, 2,2,2 ;(-1)^{\frac{k}{2}-M} 16 \mathrm{e}^{-2 \mu}\right), & \text { if } k \text { is even }  \tag{2.38}\\ \mu+\mathrm{e}^{-4 \mu}{ }_{4} F_{3}\left(1,1, \frac{3}{2}, \frac{3}{2}, 2,2,2 ;-16 \mathrm{e}^{-4 \mu}\right), & \text { if } k \text { is odd }\end{cases}
$$

The membrane part of the grand potential consists in two function $\widetilde{J}_{b}\left(\mu_{\text {eff }}, k, M\right)$ and $\widetilde{J}_{c}\left(\mu_{\text {eff }}, k, M\right)$ whose large $\mu$ expansion reads:

$$
\begin{equation*}
\widetilde{J}_{b}\left(\mu_{\mathrm{eff}}, k, M\right)=\sum_{\ell=1}^{\infty} \widetilde{b}_{\ell}(k)(-1)^{M \ell} \mathrm{e}^{-2 \ell \mu_{\mathrm{eff}}}, \quad \widetilde{J}_{c}\left(\mu_{\mathrm{eff}}, k, M\right)=\sum_{\ell=1}^{\infty} \widetilde{c}_{\ell}(k)(-1)^{M \ell} \mathrm{e}^{-2 \ell \mu_{\mathrm{eff}}} \tag{2.39}
\end{equation*}
$$

The coefficients $\widetilde{b}_{\ell}(k)$ are related to the quantum $B$ period of local $\mathbb{P}^{1} \times \mathbb{P}^{1}[13]$, and can be expressed in terms of the refined BPS invariants $N_{j_{L}, j_{R}}^{d_{1}, d_{2}}$ of this CY [21, 35, 36], as

$$
\begin{equation*}
\widetilde{b}_{\ell}(k)=-\frac{\ell}{2 \pi} \sum_{j_{L}, j_{R}} \sum_{\ell=d w} \sum_{d_{1}+d_{2}=d} N_{j_{L}, j_{R}}^{d_{1}, d_{2}} q^{\frac{w}{2}\left(d_{1}-d_{2}\right)} \frac{\sin \frac{\pi k w}{2}\left(2 j_{L}+1\right) \sin \frac{\pi k w}{2}\left(2 j_{R}+1\right)}{w^{2} \sin ^{3} \frac{\pi k w}{2}} . \tag{2.40}
\end{equation*}
$$

The particular combination of invariants appearing here involves only the so-called NekrasovShatashvili limit [37] of the topological string. The coefficients $\widetilde{c}_{\ell}(k)$ can be computed from the $\widetilde{b}_{\ell}(k)$ by using the relation conjectured in [12]

$$
\begin{equation*}
\widetilde{c}_{\ell}(k)=-k^{2} \frac{\partial}{\partial k}\left(\frac{\widetilde{b}_{\ell}(k)}{2 \ell k}\right) \tag{2.41}
\end{equation*}
$$

The worldsheet part of the grand potential $J^{\mathrm{WS}}(\mu, k)$ takes the following form

$$
\begin{equation*}
J^{\mathrm{WS}}(\mu, k, M)=\sum_{m \geq 1}(-1)^{m} d_{m}(k, M) \mathrm{e}^{-4 m \mu / k} \tag{2.42}
\end{equation*}
$$

where the coefficient $d_{m}(k, M)$ can be expressed in terms of standard Gopakumar Vara invariants $n_{g}^{d}$ of local $\mathbb{P}^{1} \times \mathbb{P}^{1}[21]$. It reads (see also [16]) $)^{2}$

$$
\begin{equation*}
d_{m}(k, M)=\sum_{g \geq 0} \sum_{d n=m} \sum_{d_{1}+d_{2}=d} n_{g}^{d_{1}, d_{2}} \beta^{\frac{d_{2}-d_{1}}{d} m}\left(2 \sin \frac{2 \pi n}{k}\right)^{2 g-2} \frac{1}{n} \tag{2.43}
\end{equation*}
$$

where

$$
\begin{equation*}
\beta=\mathrm{e}^{-2 \pi \mathrm{i} M / k} \tag{2.44}
\end{equation*}
$$

These coefficients can also be expressed in terms of BPS invariants of local $\mathbb{P}^{1} \times \mathbb{P}^{1}$

$$
\begin{equation*}
d_{m}(k, M)=\sum_{j_{L}, j_{R}} \sum_{d n=m} \sum_{d_{1}+d_{2}=d} \frac{2 j_{R}+1}{\sin ^{2} \frac{2 \pi n}{k}} \frac{\sin \left(\frac{4 \pi n}{k}\left(2 j_{L}+1\right)\right)}{n \sin \frac{4 \pi n}{k}} N_{j_{L}, j_{R}}^{d_{1}, d_{2}} \beta^{\frac{d_{2}-d_{1}}{d} m} . \tag{2.45}
\end{equation*}
$$

Notice that both the worldsheet instanton contributions (2.42) and the membrane instanton contributions (2.39) have poles at any rational value of $k$. However, these poles cancel and the the full grand potential is a well defined and finite quantity for any value of $k$. This is the so-called HMO cancellation mechanism which was first pointed out in [10] and then verified in [13] for ABJM theory. As noticed in [23], the argument of [13] naturally lifts to ABJ theory. Let us review it briefly. The worldsheet contribution (2.42) has poles whenever $k \in 2 n / \mathbb{N}$, while the membrane instanton contribution (2.39) has poles when $k \in 2 \mathbb{N} / w$. The crucial observation is that these poles mutually cancel order by order in $\mathrm{e}^{-\mu_{\text {eff }}}$ for

$$
\begin{equation*}
k=\frac{2 n}{w}=\frac{2 m}{\ell} \tag{2.46}
\end{equation*}
$$

More precisely, the worldsheet instanton contribution to the singularity is

$$
\begin{align*}
& {\left[\frac{n}{w^{4}\left(k-\frac{2 n}{w}\right)^{2}}+\frac{1}{k-\frac{2 n}{w}}\left(\frac{1}{w^{3}}+\frac{m}{n w^{2}} \mu_{\mathrm{eff}}\right)\right]\left(1+2 j_{L}\right)\left(1+2 j_{R}\right) N_{j_{L}, j_{R}}^{d_{1}, d_{2}} \mathrm{e}^{-\frac{2 m w}{n} \mu_{\mathrm{eff}}}}  \tag{2.47}\\
& \times \frac{(-1)^{m} \mathrm{e}^{-2 \pi \mathrm{i} M m\left(d_{2}-d_{1}\right) / k d}}{\pi^{2}}
\end{align*}
$$

where we used that

$$
\begin{equation*}
\sum_{d_{1}, d_{2}} N_{j_{L}, j_{R}}^{d_{1}, d_{2}}\left(d_{1}-d_{2}\right)=0 \tag{2.48}
\end{equation*}
$$

Similarly, the singular part of $\mu_{\text {eff }} \widetilde{J}_{b}\left(\mu_{\text {eff }}, k\right)$ is

$$
\begin{equation*}
-\frac{(-1)^{M \ell} \mathrm{e}^{\pi \mathrm{i} k w\left(d_{1}-d_{2}\right) / 2}}{\pi^{2}} \frac{\ell}{w^{3}\left(k-\frac{2 n}{w}\right)}\left(1+2 j_{L}\right)\left(1+2 j_{R}\right) N_{j_{L}, j_{R}}^{d_{1}, d_{2}} \mu_{\mathrm{eff}} \mathrm{e}^{-2 \ell \mu_{\mathrm{eff}}} \tag{2.49}
\end{equation*}
$$

where we used the argument explained in [13] to set

$$
\begin{equation*}
(-1)^{n\left(2 j_{L}+2 j_{R}-1\right)}=1 . \tag{2.50}
\end{equation*}
$$

A similar computation gives the pole structure for $\widetilde{J}_{c}\left(\mu_{\mathrm{eff}}, k\right)$, which reads

$$
\begin{equation*}
-(-1)^{M \ell} \frac{\mathrm{e}^{\pi \mathrm{i} k w\left(d_{1}-d_{2}\right) / 2}}{\pi^{2}}\left[\frac{n}{w^{4}\left(k-\frac{2 n}{w}\right)^{2}}+\frac{1}{w^{3}\left(k-\frac{2 n}{w}\right)}\right]\left(1+2 j_{L}\right)\left(1+2 j_{R}\right) N_{j_{L}, j_{R}}^{d_{1}, d_{2}} \mathrm{e}^{-2 \ell \mu_{\text {eff }}} . \tag{2.51}
\end{equation*}
$$

[^1]

Figure 2: The region of phase space (3.1) for ABJM theory, at large energy.

A simple computation shows that

$$
\begin{align*}
\mathrm{e}^{\mathrm{i} \pi k w\left(d_{1}-d_{2}\right) / 2} & =(-1)^{m}, \\
\mathrm{e}^{-2 \pi \mathrm{i} M m\left(d_{2}-d_{1}\right) / k d} & =\mathrm{e}^{\mathrm{i} \pi M \ell} \mathrm{e}^{2 \mathrm{i} \pi M d_{1} 2 n / k}=(-1)^{M \ell}, \tag{2.52}
\end{align*}
$$

where we used (2.46) and the relation $m=d n$. It is now straightforward to see that the singularity in (2.47) precisely cancels against these of (2.49) and (2.51).

## 3. Spectral determinant and quantization conditions

The physical information on the $\operatorname{ABJ}(\mathrm{M})$ Fermi gas can be encoded in many different ways: in the spectrum of the Hamiltonian, in the spectral determinant, and in the modified grand potential. It is clear that these three objects are equivalent, but their relationship is not trivial. The purpose of this paper is to use the explicit answer for the modified grand potential of $\operatorname{ABJ}(\mathrm{M})$ theory in order to find a useful expression for the spectral determinant and to solve the spectral problem of the Hamiltonian. A natural starting point to find the spectrum is to use the WKB method. In the Bohr-Sommerfeld approximation, the energy levels of the Hamiltonian can be obtained in the following way. Let $\operatorname{Vol}_{\mathrm{cl}}(E, M)$ be the classical phase space volume, i.e. the volume of the region

$$
\begin{equation*}
\mathcal{R}_{\mathrm{cl}}(E)=\left\{(x, p) \in \mathbb{R}^{2}: \rho_{\mathrm{cl}}(x, p) \leq \mathrm{e}^{-E}\right\} . \tag{3.1}
\end{equation*}
$$

Here we denoted by $\rho_{\mathrm{cl}}(x, p)$ the classical limit ${ }^{3}$ of the quantum operator $\hat{\rho}$, which is given by

$$
\begin{equation*}
\rho_{\mathrm{cl}}(x, p)=\mathrm{e}^{-T(p)-U(x, M)}, \tag{3.2}
\end{equation*}
$$

where ${ }^{4}$

$$
\begin{equation*}
T(p)=\log \left(2 \cosh \frac{p}{2}\right), \quad U(x, M)=-\log \left(V_{M}(x)\right) . \tag{3.3}
\end{equation*}
$$

[^2]The Bohr-Sommerfeld quantization condition reads

$$
\begin{equation*}
\operatorname{Vol}_{\mathrm{cl}}(E, k, M)=2 \pi \hbar\left(n+\frac{1}{2}\right), \quad n=0,1,2, \cdots \tag{3.4}
\end{equation*}
$$

For large values of the energy one finds

$$
\begin{equation*}
\operatorname{Vol}_{\mathrm{cl}}(E, k, M) \approx 8 E^{2} \tag{3.5}
\end{equation*}
$$

as shown for instance in Figure 2. Notice that the region $\mathcal{R}_{\mathrm{cl}}(E)$ has a finite volume, corresponding to the fact that the operator $\hat{\rho}$ has a discrete spectrum. In [7] it was pointed out that the classical volume receives two types of quantum corrections, perturbative and non perturbative in $\hbar$, and there should be a fully "quantum" version of the classical function $\operatorname{Vol}_{\mathrm{cl}}(E, k, M)$ incorporating these corrections, which we will denote by $\operatorname{Vol}(E, k, M)$. It is convenient to write the quantum volume as in [15],

$$
\begin{equation*}
\operatorname{Vol}(E, k, M)=\operatorname{Vol}_{\mathrm{p}}(E, k, M)+\operatorname{Vol}_{\mathrm{np}}(E, k, M) \tag{3.6}
\end{equation*}
$$

where $\operatorname{Vol}_{\mathrm{p}}(E, k, M)$ contains the full series of perturbative corrections in $\hbar$, while $\operatorname{Vol}_{\mathrm{np}}(E, k, M)$ contains the non-perturbative corrections in $\hbar$. The Bohr-Sommerfeld quantization condition should be promoted to an exact WKB quantization condition of the form

$$
\begin{equation*}
\operatorname{Vol}(E, k, M)=2 \pi \hbar\left(n+\frac{1}{2}\right), \quad n=0,1,2, \cdots \tag{3.7}
\end{equation*}
$$

similar to what happens in some problems in ordinary Quantum Mechanics [38]. In [15], the relation between the quantum volume and the grand potential was studied by using the relationship,

$$
\begin{equation*}
\mathcal{J}(\mu, k, M)=\frac{1}{2 \pi \hbar} \int_{E_{0}}^{\infty} \frac{\operatorname{Vol}(E, k, M) \mathrm{d} E}{\mathrm{e}^{E-\mu}+1}-\sum_{r \geq 1} \frac{B_{2 r}}{(2 r)!} f^{(2 r-1)}(0) \tag{3.8}
\end{equation*}
$$

where

$$
\begin{equation*}
f(n)=\log \left(1+\mathrm{e}^{\mu-E_{n}}\right) . \tag{3.9}
\end{equation*}
$$

From a practical point of view, this strategy leads to many technical difficulties in the study of the non-perturbative sector. Here we overcome these difficulties by using the approach of [17], where the quantum volume and the spectrum are computed by studying the zeros of the spectral determinant ${ }^{5}$. Our first goal is then to find a convenient expression for the spectral determinant of these theories.

In the case of maximally supersymmetric theories, it was shown in [17] that the sum appearing in (2.29) can be written in terms of Jacobi theta functions. It is easy to see, by a computation similar to the one presented in [20], that the spectral determinant for $\operatorname{ABJ}(\mathrm{M})$ theory (2.29) is given by

$$
\begin{equation*}
\Xi(\mu, k, M)=\mathrm{e}^{J(\mu, k, M)} \Theta(\mu, k, M) \tag{3.10}
\end{equation*}
$$

where

$$
\begin{align*}
\Theta(\mu, k, M)=\sum_{n \in \mathbb{Z}} \exp \{ & -4 \pi^{2} n^{2} C(k) \mu_{\mathrm{eff}}+2 \pi \mathrm{in}\left(C(k) \mu_{\mathrm{eff}}^{2}+B(k, M)+\widetilde{J}_{b}\left(\mu_{\mathrm{eff}}, k, M\right)\right)  \tag{3.11}\\
& \left.+J_{\mathrm{WS}}\left(\mu_{\mathrm{eff}}+2 \pi \mathrm{i} n, k, M\right)-J_{\mathrm{WS}}\left(\mu_{\mathrm{eff}}, k, M\right)-\frac{8 \pi^{3} \mathrm{i} n^{3}}{3} C(k)\right\} .
\end{align*}
$$

[^3]We will call this function a generalized theta function.
As we noted in (2.22), the spectrum of energies in (2.17) can be obtained by looking at the zeros of the spectral determinant, by setting

$$
\begin{equation*}
\mu=E+\pi \mathrm{i} . \tag{3.12}
\end{equation*}
$$

As it was found in $[17,20]$, this involves looking at the zeros of the (generalized) theta function, and leads to a quantization condition of the form (3.7) which incorporates all the perturbative and non-perturbative corrections to the Bohr-Sommerfeld condition (3.4). It is easy to see that

$$
\begin{align*}
\Theta(E+\pi \mathrm{i}, k, M)=\mathrm{e}^{\zeta} \sum_{n \in \mathbb{Z}} \exp \{ & -4 \pi^{2}(n+1 / 2)^{2} C(k) E_{\text {eff }}-\frac{8 \pi^{3} \mathrm{i}(n+1 / 2)^{3}}{3} C(k) \\
& +2 \pi \mathrm{i}(n+1 / 2)\left(C(k) E_{\text {eff }}^{2}+B(k, M)+\tilde{J}_{b}\left(E_{\text {eff }}+\pi \mathrm{i}, k, M\right)\right) \\
& \left.+f_{\mathrm{WS}}\left(E_{\text {eff }}+\pi \mathrm{i}, n\right)-\frac{1}{2} f_{\mathrm{WS}}\left(E_{\text {eff }}+\pi \mathrm{i},-1\right)\right\} . \tag{3.13}
\end{align*}
$$

where

$$
\begin{equation*}
f_{\mathrm{WS}}(\mu, n)=\sum_{m \geq 1} d_{m}(k, M)\left(\mathrm{e}^{-8 \pi \mathrm{i} m n / k}-1\right)(-1)^{m} \mathrm{e}^{-4 m \mu / k}, \tag{3.14}
\end{equation*}
$$

in particular

$$
\begin{equation*}
f_{\mathrm{WS}}\left(E_{\text {eff }}+\pi \mathrm{i},-1\right)=2 \mathrm{i} \sum_{m \geq 1} d_{m}(k, M) \sin \frac{4 \pi m}{k}(-1)^{m} \mathrm{e}^{-4 m E_{\text {eff }} / k} \tag{3.15}
\end{equation*}
$$

We also have that

$$
\begin{align*}
\zeta & =\frac{2}{k} E_{\text {eff }}-\pi \mathrm{i}\left(\frac{2}{\pi^{2} k} E_{\text {eff }}^{2}+B(k, M)+\tilde{J}_{b}\left(E_{\text {eff }}+\pi \mathrm{i}, k, M\right)\right)  \tag{3.16}\\
& +\frac{1}{2} f_{\mathrm{WS}}\left(E_{\text {eff }}+\pi \mathrm{i},-1\right)+\frac{2 \pi \mathrm{i}}{3 k}
\end{align*}
$$

Note that, if we just take into account the terms with $n=0,-1$ of the generalized theta function (3.13) we obtain the quantization condition

$$
\begin{equation*}
\cos (\pi \Omega(E, k, M))=0 \tag{3.17}
\end{equation*}
$$

where

$$
\begin{align*}
\Omega(E, k, M)= & C(k) E_{\text {eff }}^{2}+B(k, M)-\frac{\pi^{2} C(k)}{3}+\sum_{\ell=1}^{\infty} \widetilde{b}_{\ell}(k)(-1)^{M \ell} \mathrm{e}^{-2 \ell E_{\text {eff }}} \\
& -\frac{1}{\pi} \sum_{m \geq 1}(-1)^{m} d_{m}(k, M) \sin \left(\frac{4 \pi m}{k}\right) \mathrm{e}^{-4 m E_{\text {eff }} / k} . \tag{3.18}
\end{align*}
$$

This is precisely the quantization condition proposed in $[15,16]$. However, there will be in general corrections to this condition, due to the remaining terms in (3.13). In order to take them into account systematically, let us call this correction $\lambda(E)$ and write the exact quantization condition as

$$
\begin{equation*}
\Omega(E, k, M)+\lambda(E)=s+\frac{1}{2}, \quad s=0,1,2, \ldots \tag{3.19}
\end{equation*}
$$

A straightforward calculation shows that $\lambda(E)$, which is non-perturbative in $k$ (i.e. in $\hbar$ ), is determined by the following equation [20]

$$
\begin{equation*}
\sum_{n \geq 0} \mathrm{e}^{-\frac{8 n(n+1)}{k} E_{\text {eff }}}(-1)^{n} \mathrm{e}^{f_{\mathrm{c}}(n)} \sin \left(\frac{8 \pi n(n+1)(2 n+1)}{3 k}+f_{\mathrm{s}}(n)+2 \pi\left(n+\frac{1}{2}\right) \lambda(E)\right)=0 \tag{3.20}
\end{equation*}
$$

where

$$
\begin{align*}
& f_{\mathrm{c}}(n)=\sum_{m \geq 1}(-1)^{m} d_{m}(k, M)\left[\cos \left(\frac{4 \pi m(2 n+1)}{k}\right)-\cos \left(\frac{4 \pi m}{k}\right)\right] \mathrm{e}^{-\frac{4 m}{k} E_{\text {eff }}}, \\
& f_{\mathrm{s}}(n)=\sum_{m \geq 1}(-1)^{m} d_{m}(k, M)\left[\sin \left(\frac{4 \pi m(2 n+1)}{k}\right)-(2 n+1) \sin \left(\frac{4 \pi m}{k}\right)\right] \mathrm{e}^{-\frac{4 m}{k} E_{\text {eff }} .} \tag{3.21}
\end{align*}
$$

We also conclude from this analysis that the exact quantum volume is given by

$$
\begin{equation*}
\operatorname{Vol}(E, k, M)=2 \pi \hbar(\Omega(E, k, M)+\lambda(E)) \tag{3.22}
\end{equation*}
$$

Note that the perturbative part of this quantum volume is given precisely by the WKB perturbative contribution, encoded in the quantum B-period,

$$
\begin{equation*}
\frac{1}{2 \pi \hbar} \operatorname{Vol}_{\mathrm{p}}(E, k, M)=C(k) E_{\text {eff }}^{2}+B(k, M)-\frac{\pi^{2} C(k)}{3}+\sum_{\ell=1}^{\infty} \widetilde{b}_{\ell}(k)(-1)^{M \ell} \mathrm{e}^{-2 \ell E_{\mathrm{eff}}}, \tag{3.23}
\end{equation*}
$$

while the non-perturbative contribution is given by

$$
\begin{equation*}
\frac{1}{2 \pi \hbar} \operatorname{Vol}_{\mathrm{np}}(E, k, M)=-\frac{1}{\pi} \sum_{m \geq 1}(-1)^{m} d_{m}(k, M) \sin \left(\frac{4 \pi m}{k}\right) \mathrm{e}^{-4 m E_{\text {eff }} / k}+\lambda(E) \tag{3.24}
\end{equation*}
$$

The perturbative and the non-perturbative part are separately divergent when $k$ is rational, as noted in [15], but the total quantum volume (3.22) is smooth, since the singularities cancel as a consequence of the HMO mechanism (indeed, the quantum volume is obtained from the modified grand potential, which is singularity-free). The non-perturbative part is then needed to cancel the singularities in the WKB perturbative expansion, and it contains crucial information on the spectrum. For example, as shown in appendix A, when $k$ is an integer, the finite part of $\tilde{J}_{b}\left(E_{\text {eff }}+\pi \mathrm{i}, k, M\right)$ vanishes and the quantum volume is largely determined by the worldsheet instanton contribution.

The correction $\lambda(E)$ can be computed analytically, in a series expansion in $\mathrm{e}^{-4 E_{\text {eff }} / k}$. It is easy to see that $\lambda(E)$ is of the form

$$
\begin{equation*}
\lambda(E)=\sum_{\ell \geq 1} \lambda_{\ell} \mathrm{e}^{-\frac{4 \ell+12}{k} E_{\text {eff }}}, \tag{3.25}
\end{equation*}
$$

where the first few terms are explicitly given by

$$
\begin{align*}
\lambda_{1}= & \frac{1}{\pi} \sin (16 x), \\
\lambda_{2}= & \frac{4}{\pi} \sin ^{2}(4 x) \sin (20 x) d_{1}, \\
\lambda_{3}= & \frac{8}{\pi} \sin ^{2}(4 x) \sin (24 x)\left[\sin ^{2}(4 x) d_{1}^{2}-2 \cos ^{2}(4 x) d_{2}\right],  \tag{3.26}\\
\lambda_{4}= & \frac{4}{3 \pi} \sin ^{2}(4 x) \sin (28 x)\left[3\left(d_{1}^{3}-2 d_{1} d_{2}+3 d_{3}\right)-4 \cos (8 x)\left(d_{1}^{3}-3 d_{3}\right)\right. \\
& \left.+\cos (16 x)\left(d_{1}^{3}+6 d_{1} d_{2}+6 d_{3}\right)\right]
\end{align*}
$$

Energy levels for $k=3, M=0$

| Order | $E_{0}$ | $E_{1}$ |
| :--- | :--- | :--- |
| $\mathrm{e}^{-4 E / 3}$ | $\underline{2.65297702084083921}$ | $\underline{4.68940459079460092512108986442}$ |
| $\mathrm{e}^{-12 E / 3}$ | $\underline{2.65156164019289190}$ | $\underline{4.68940134450544960666103687122}$ |
| $\mathrm{e}^{-24 E / 3}$ | $\underline{2.65156833993530136}$ | $\underline{4.68940134457031678330042507336}$ |
| $\mathrm{e}^{-32 E / 3}$ | $\underline{2.65156833716} 940289$ | $\underline{4.68940134457031677561757482976}$ |
| $\mathrm{e}^{-40 E / 3}$ | $\underline{2.65156833716875544}$ | $\underline{4.68940134457031677561753154101}$ |
| $\mathrm{e}^{-52 E / 3}$ | $\underline{4.68940134457031677561753154681}$ |  |
| Numerical value | 2.6515683371688576155 | 4.68940134457031677561753154681 |

Table 1: The first two energy levels for $k=3$ and $M=0$ calculated analytically from (3.19). In the last line numerical values evaluated by the method in appendix C are given. At each order of the approximation, we underline the digits which agree with the numerical result.

Energy levels for $k=5, M=0$

| Order | $E_{0}$ | $E_{1}$ |
| :--- | :--- | :--- |
| $\mathrm{e}^{-4 E / 5}$ | $\underline{3.0} 475013693$ | $\underline{5.79353763401508120749977}$ |
| $\mathrm{e}^{-16 E / 5}$ | $\underline{3.0724584475}$ | $\underline{5.79369469126135544218070}$ |
| $\mathrm{e}^{-32 E / 5}$ | $\underline{3.0724359155}$ | $\underline{5.79369469107338173784939}$ |
| $\mathrm{e}^{-48 E / 5}$ | $\underline{3.0724358357}$ | $\underline{5.79369469107338158412549}$ |
| Numerical value | 3.0724358360 | 5.79369469107338158412559 |


| Energy levels for $k=6, M=0$ |  |  |
| :--- | :--- | :--- |
| Order | $E_{0}$ | $E_{1}$ |
| $\mathrm{e}^{-4 E / 6}$ | $\underline{3.21322311}$ | $\underline{6.23111654150891713732}$ |
| $\mathrm{e}^{-16 E / 6}$ | $\underline{3.23510192}$ | $\underline{6.23141999560231213049}$ |
| $\mathrm{e}^{-32 E / 6}$ | $\underline{3.23464705}$ | $\underline{6.23141998018954896785}$ |
| $\mathrm{e}^{-48 E / 6}$ | $\underline{3.234644} 06$ | $\underline{6.23141998018953286312}$ |
| Numerical value | 3.23464413 | 6.23141998018953286330 |

Table 2: The first two energy levels for $k=5,6$ and $M=0$ calculated analytically from (3.19).
with

$$
\begin{align*}
& d_{1}=d_{1}(k, M)=-\cos (2 M x) \csc ^{2}(2 x) \\
& d_{2}=d_{2}(k, M)=-\csc ^{2}(2 x)-\frac{1}{2} \cos (4 M x) \csc ^{2}(4 x)  \tag{3.27}\\
& d_{3}=d_{3}(k, M)=-3 \cos (2 M x) \csc ^{2}(2 x)-\frac{1}{3} \cos (6 M x) \csc ^{2}(6 x)
\end{align*}
$$

and we have denoted

$$
\begin{equation*}
x=\frac{\pi}{k} \tag{3.28}
\end{equation*}
$$

Note that, when $k=1,2,4$, we have that $f_{\mathrm{s}}(n)=0$, and the first term in the argument of the sine in (3.20) is always a multiple of $\pi$. Therefore, the solution to (3.20) is $\lambda(E)=0$, i.e. the correction vanishes and the quantization condition of $[15,16]$ is exact.


Figure 3: The difference $\Delta_{0}^{(k)}(m)(3.31)$ as a function of $m$, the number of instantons included in the computation. The line on the bottom (in blue) gives $\Delta_{0}^{(3)}(m)$, while the line on the top (in red) gives $\Delta_{0}^{(5)}(m)$.

Energy levels for $k=3, M=1$

| Order | $E_{0}$ | $E_{1}$ |
| :--- | :--- | :--- |
| $\mathrm{e}^{-8 E / 3}$ | $\underline{3.4866971392036144}$ | $\underline{5.190229102060787166657584}$ |
| $\mathrm{e}^{-12 E / 3}$ | $\underline{3.4866953311076197}$ | $\underline{5.190229100088979204304574}$ |
| $\mathrm{e}^{-20 E / 3}$ | $\underline{3.4866953293248867}$ | $\underline{5.190229100088834784550474}$ |
| $\mathrm{e}^{-28 E / 3}$ | $\underline{3.4866953293348793}$ | $\underline{5.190229100088834796616934}$ |
| Numerical value | 3.4866953293348782 | 5.190229100088834796616924 |

Table 3: The first two energy levels for $k=3, M=1$ calculated analytically from (3.19).

|  | Energy levels for $k=5, M=2$ |  |
| :--- | :--- | :--- |
| Order | $E_{0}$ | $E_{1}$ |
| $\mathrm{e}^{-8 E / 3}$ | $\underline{4.8544694530582}$ | $\underline{6.95012371466050570772}$ |
| $\mathrm{e}^{-12 E / 3}$ | $\underline{4.8544533648860}$ | $\underline{6.95012364154401512072}$ |
| $\mathrm{e}^{-20 E / 3}$ | $\underline{4.8544536519209}$ | $\underline{6.95012364179272664768}$ |
| $\mathrm{e}^{-28 E / 3}$ | $\underline{4.8544536515325}$ | $\underline{6.95012364179271448721}$ |
| Numerical value | 4.8544536515315 | 6.95012364179271448613 |

Table 4: The first two energy levels for $k=5, M=2$ calculated analytically from (3.19).

Let us now give some concrete results for the correction series (3.25). In the case of ABJM theory, with $M=0$, the first few corrections read

$$
\begin{align*}
& \lambda_{1}=\frac{1}{\pi} \sin (16 x), \\
& \lambda_{2}=-\frac{4}{\pi} \sin ^{2}(4 x) \sin (20 x) \csc ^{2}(2 x), \\
& \lambda_{3}=\frac{8}{\pi} \sin (24 x) \csc ^{2}(2 x)\left(3 \sin ^{2}(4 x) \sin ^{2}(6 x)+\sin ^{2}(2 x) \sin ^{2}(8 x)+\sin ^{2}(10 x)\right), \\
& \lambda_{4}=-\frac{8}{\pi} \sin ^{2}(4 x) \sin (28 x) \csc ^{2}(2 x)(23+22 \cos (4 x)+19 \cos (8 x)+4 \cos (12 x)+3 \cos (16 x)) . \tag{3.29}
\end{align*}
$$

The results for $\lambda_{1}, \lambda_{2}$ and $\lambda_{3}$ are in perfect agreement with the results of [18] based on numerical fitting. We can check as well the higher order corrections by performing a detailed numerical analysis. First of all, we compute the first two energy levels in (2.17), with high precision. To do this, we follow a procedure inspired by the analysis of [8] and summarized in Appendix C. On the other hand, we use (3.20) to compute the corrections to the quantum volume up to

$$
\begin{equation*}
\lambda_{10} \mathrm{e}^{-\frac{52}{k} E_{\text {eff }}} \tag{3.30}
\end{equation*}
$$

The results are shown in tables 1 and 2 , for $k=3,5,6$. As expected, the more instanton corrections we include in the analytic computation, the more we approach the numerical value. This can be seen in detail by considering the quantity

$$
\begin{equation*}
\Delta_{0}^{(k)}(m)=\log _{10}\left|E_{0}^{\text {num }}(k)-E_{0}^{(m)}(k)\right| \tag{3.31}
\end{equation*}
$$

where $E_{0}^{\text {num }}$ is the numerical value of the ground state energy, and $E_{0}^{(m)}$ is the value computed from (3.20) by including the first $m$ instanton corrections. As shown in Fig. 3, $E_{0}^{(m)}$ converges to $E_{0}^{\text {num }}$ as $m$ grows. However, it does not converge monotonically, in contrast to what happened in the numerical analysis of $[15]$ for $k=1,2$ and $M=0$. In tables 3,4 we give additional numerical evidence for the validity of the quantization condition in the ABJ case with $M \neq 0$.

## 4. A case study with $\mathcal{N}=6$ supersymmetry

As emphasized in [20], the quantization condition studied in the previous section is a consequence of a stronger result, namely our explicit formula (3.10) for the spectral determinant. In principle, using this formula, one can compute the canonical partition functions by performing an expansion around $\kappa=0$ as in (2.24). In [17] this was checked in detail in the maximally supersymmetric cases, by using the computations of the partition functions in $[9,8,10,23,22]$. In the case with maximal supersymmetry, the generalized theta function becomes a standard Jacobi theta function, and the higher genus contribution to the modified grand potential vanish, so the analysis of the spectral determinant is relatively straightforward.

In this section we analyze in detail a case with $\mathcal{N}=6$ supersymmetry, namely ABJM theory with $k=4$. This case is slightly richer than the maximally supersymmetric cases because the grand potential involves the all genus worldsheet instantons. However, the generalized theta function of this theory becomes a conventional theta function, as in the maximally supersymmetric cases, so this case is not the most generic one, but it is a good starting point to start exploring the spectral determinants of theories with $\mathcal{N}=6$ supersymmetry.

It follows from (2.30) that the grand potential of ABJM theory with $k=4$ is given by

$$
\begin{equation*}
J(\mu, 4)=J^{(\mathrm{p})}\left(\mu_{\mathrm{eff}}, 4\right)+J^{\mathrm{WS}}\left(\mu_{\mathrm{eff}}, 4\right)+\mu_{\mathrm{eff}} \widetilde{J}_{b}\left(\mu_{\mathrm{eff}}, 4\right)+\widetilde{J}_{c}\left(\mu_{\mathrm{eff}}, 4\right) . \tag{4.1}
\end{equation*}
$$

To calculate this quantity we have to take the limit $k \rightarrow 4$ in the general expression and be careful with the poles, as in [17]. These however cancel, as we recalled above, so we can compute (4.1) by considering only its finite part. In particular $\widetilde{J}_{b}\left(\mu_{\text {eff }}, 4\right)$ has no finite part, as shown in appendix A, while the finite part of $\widetilde{J}_{c}\left(\mu_{\text {eff }}, 4\right)$ is

$$
\begin{equation*}
\frac{1}{12}\left(-\log \left(1-16 y^{2}\right)+\tilde{\varpi}_{1}\left(-y^{2}\right)\right), \quad y=\mathrm{e}^{-\mu} \tag{4.2}
\end{equation*}
$$

where

$$
\begin{equation*}
\tilde{\varpi}_{1}(z)=-4 z_{4} F_{3}\left(1,1, \frac{3}{2}, \frac{3}{2} ; 2,2,2 ;-16 z\right) . \tag{4.3}
\end{equation*}
$$

Let us now look at the worldsheet instanton part. For general $k$ and $M=0$ the expression (2.42) reads

$$
\begin{equation*}
J^{\mathrm{WS}}(\mu, k)=\sum_{g \geq 0} \sum_{w, d \geq 1} \frac{1}{w}(-1)^{d w} n_{g}^{d}\left(2 \sin \frac{2 \pi w}{k}\right)^{2 g-2} \mathrm{e}^{-4 d w \mu / k} \tag{4.4}
\end{equation*}
$$

It is convenient to split the sum over $w$ into even and odd part. This leads to

$$
\begin{align*}
J^{\mathrm{WS}}(\mu, k)= & \sum_{g \geq 0} \sum_{w, d \geq 1} \frac{1}{2 w} n_{g}^{d}\left(2 \sin \frac{2 \pi w}{(k / 2)}\right)^{2 g-2} \mathrm{e}^{-4 d w \mu /(k / 2)} \\
& +\sum_{g \geq 0} \sum_{w, d \geq 1} \frac{1}{2 w-1}(-1)^{d} n_{g}^{d}\left(2 \sin \frac{2 \pi(2 w-1)}{k}\right)^{2 g-2} \mathrm{e}^{-4 d(2 w-1) \mu / k} \tag{4.5}
\end{align*}
$$

When $k=4$, the first part is (up to a factor $1 / 2$ ) precisely the worldsheet instanton part of the modified grand potential for the maximally supersymmetric theory $(k, M)=(2,1)$ analyzed in [17]. The second part is a non trivial quantity which contains all genus contributions. We can then write the worldsheet instanton part of (4.1) as

$$
\begin{equation*}
J^{\mathrm{WS}}\left(\mu_{\mathrm{eff}}, 4\right)=\frac{1}{2} J^{\mathrm{WS}}\left(\mu_{\mathrm{eff}}, 2,1\right)+I\left(\mu_{\mathrm{eff}}\right) \tag{4.6}
\end{equation*}
$$

The first term was computed in [17] and reads

$$
\begin{equation*}
J^{\mathrm{WS}}(\mu, 2,1)=\frac{\mu_{\mathrm{eff}}^{2}}{2 \pi^{2}} \partial_{t}^{2} F_{0}^{\mathrm{inst}}(t)-\frac{\mu_{\mathrm{eff}}}{2 \pi^{2}} \partial_{t} F_{0}^{\mathrm{inst}}(t)+\frac{1}{4 \pi^{2}} F_{0}^{\mathrm{inst}}(t)+F_{1}^{\mathrm{inst}}(t) \tag{4.7}
\end{equation*}
$$

Here we used $t=2 \mu_{\text {eff }}$ and

$$
\begin{equation*}
F_{0}^{\mathrm{inst}}(t)=\sum_{d, w} n_{0}^{d} \frac{1}{w^{3}} \mathrm{e}^{-d w t}, \quad F_{1}^{\mathrm{inst}}(t)=\sum_{d, w}\left(\frac{n_{0}^{d}}{12}+n_{1}^{0}\right) \mathrm{e}^{-d w t} \tag{4.8}
\end{equation*}
$$

The second part of $J^{\mathrm{WS}}$ contains contributions at all genera and can be written as

$$
\begin{equation*}
I\left(\mu_{\mathrm{eff}}\right)=\sum_{g \geq 0} \sum_{w, d \geq 1} \frac{(-1)^{d} n_{g}^{d}}{2 w-1} 4^{g-1} \mathrm{e}^{-d(2 w-1) \mu_{\mathrm{eff}}}=\sum_{g \geq 0} \sum_{d \geq 1} 4^{g-1}(-1)^{d} n_{g}^{d} \tanh ^{-1}\left(\mathrm{e}^{-d \mu_{\mathrm{eff}}}\right) \tag{4.9}
\end{equation*}
$$

By looking at the small $y$ expansion of this quantity we conjecture that it can be resummed in closed form in terms of an elliptic integral of the first kind,

$$
\begin{equation*}
I\left(\mu_{\mathrm{eff}}\right)=-\frac{1}{4} \log \left(\frac{2 K\left(16 y^{2}\right)}{\pi}\right)+\frac{1}{4} \tanh ^{-1}(4 y) \tag{4.10}
\end{equation*}
$$

We have checked the above equality order by order in a series expansion $y$, but we do not have a general proof of it. However, we will see that this conjecture reproduces the correct $Z(N, 4)$ appearing in the large $y$ expansion of the spectral determinant (2.28). This strongly suggests that (4.10) is a true identity. The existence of an identity like this is remarkable, since it resums
the all-genus Gopakumar-Vafa expansion of the free energy. This resummation is needed if we want to reproduce the canonical partition functions: this requires an expansion around $y \rightarrow \infty$, while the original Gopakumar-Vafa expansion only holds at large radius, i.e. for $y \rightarrow 0$.

By using (4.10) one finds

$$
\begin{align*}
J(\mu, 4)= & A(4,0)+F_{1}(t)+F_{1}^{\mathrm{NS}}(t)+\frac{\mu}{4}+\frac{1}{4} \tanh ^{-1}\left(4 \mathrm{e}^{-\mu}\right) \\
& +\frac{1}{8 \pi^{2}}\left(F_{0}(t)-t \partial_{t} F_{0}(t)+\frac{1}{2} t^{2} \partial_{t}^{2} F_{0}(t)\right) \tag{4.11}
\end{align*}
$$

where

$$
\begin{align*}
F_{0}(t) & =\frac{t^{3}}{6}+F_{0}^{\mathrm{inst}}(t) \\
F_{1}(t) & =-\frac{1}{12} \log \left[64 y^{2}\left(1-16 y^{2}\right)\right]-\frac{1}{2} \log \left(\frac{K\left(16 y^{2}\right)}{\pi}\right),  \tag{4.12}\\
F_{1}^{\mathrm{NS}}(t) & =-\frac{1}{24} \log \frac{\left(1-16 y^{2}\right)}{y^{4}} .
\end{align*}
$$

The constant $A(4,0)$ is the standard constant map contribution of ABJM, whose exact value is given by

$$
\begin{equation*}
A(4,0)=-\frac{\zeta(3)}{4 \pi^{2}}-\frac{\log (2)}{2} \tag{4.13}
\end{equation*}
$$

The derivative of the grand potential takes the following closed form,

$$
\begin{equation*}
\partial_{\mu} J(\mu, k=4)=\frac{\mu_{\mathrm{eff}}^{2}}{8\left(1-16 y^{2}\right) K\left(16 y^{2}\right)^{2}}-\frac{E\left(16 y^{2}\right)}{2\left(16 y^{2}-1\right) K\left(16 y^{2}\right)}+\frac{1-2 y}{8 y-2}+\frac{1}{4} \tag{4.14}
\end{equation*}
$$

The large $\mu$ expansion of (4.11) reads

$$
\begin{align*}
J(\mu, 4)= & \frac{\mu^{3}}{6 \pi^{2}}+\frac{\mu}{4}-\frac{\zeta(3)}{4 \pi^{2}}-\frac{\log (2)}{2}+y+\frac{\left(-4 \mu^{2}-2 \mu-1\right) y^{2}}{2 \pi^{2}}  \tag{4.15}\\
& +\frac{16 y^{3}}{3}+\frac{\left(-208 \mu^{2}-4 \mu+32 \pi^{2}-9\right) y^{4}}{16 \pi^{2}}+\frac{256 y^{5}}{5}+\mathcal{O}\left(y^{6}\right)
\end{align*}
$$

Once we know the modified grand potential, we can use (2.25) and (2.29) to obtain the grand canonical partition function or spectral determinant:

$$
\begin{equation*}
\Xi(\kappa, 4)=\mathrm{e}^{J(\mu, 4)} \vartheta_{3}\left(\frac{1}{2}\left(\xi-\frac{\tau}{4}-\frac{1}{12}\right), \frac{\tau}{2}\right) \tag{4.16}
\end{equation*}
$$

where

$$
\begin{equation*}
\tau=\frac{2 \mathrm{i}}{\pi} \partial_{t}^{2} F_{0}(t), \quad \xi=\frac{1}{2 \pi}\left(t \partial_{t}^{2} F_{0}(t)-\partial_{t} F_{0}(t)\right) \tag{4.17}
\end{equation*}
$$

Note that, although this theory is not maximally supersymmetric, the generalized theta function becomes in this case a Jacobi theta function. As explained in $[17,20]$ the spectrum is determined by the zeros of the theta function, and we find the exact quantization condition

$$
\begin{equation*}
\xi-\frac{\tau}{4}-\frac{1}{12}=2 n+1, \quad n=0,1,2, \cdots \tag{4.18}
\end{equation*}
$$

Interestingly, the quantum volume for $k=4$ is exactly related to that for $k=2$ and $M=1$

$$
\begin{equation*}
\operatorname{Vol}(E, k=4)=\frac{1}{2}\left(\xi-\frac{\tau}{4}-\frac{1}{12}\right) 2 \pi \hbar=\operatorname{Vol}(E, k=2, M=1)+4 \pi^{2}, \quad(\hbar=8 \pi) \tag{4.19}
\end{equation*}
$$

where $\operatorname{Vol}(E, k=2, M=1)$ is the quantum volume for the maximally supersymmetric theory with $k=2, M=1$ obtained in [17]. Notice that, as we mentioned before, in this case the approximate quantization condition of [15] is exact and does not need additional corrections.

By following the arguments of [17], we can now use modular properties and analytic continuation to write the special determinant (4.16) in the orbifold frame, i.e. in the region $\mu \rightarrow-\infty$, where we make contact with the expansion (2.24). This will allow us to compute the exact values of $Z(N, 4)$ for finite $N$, which provides a check of the formula (3.10). In order to proceed we introduce the orbifold periods [5]:

$$
\begin{align*}
\lambda & =\mathrm{i} \frac{\kappa}{8 \pi}{ }_{3} F_{2}\left(\frac{1}{2}, \frac{1}{2}, \frac{1}{2} ; 1, \frac{3}{2} ; \frac{\kappa^{2}}{16}\right) \\
\partial_{\lambda} \mathcal{F}_{0}(\lambda) & =\mathrm{i} \frac{\kappa}{4} G_{3,3}^{2,3}\left(\begin{array}{ll}
\frac{1}{2}, & \frac{1}{2}, \\
0, & \frac{1}{2} \\
0 & -\frac{1}{2}
\end{array} \frac{\kappa^{2}}{16}\right)-\frac{\pi^{2} \kappa}{2}{ }_{3} F_{2}\left(\frac{1}{2}, \frac{1}{2}, \frac{1}{2} ; 1, \frac{3}{2} ; \frac{\kappa^{2}}{16}\right) . \tag{4.20}
\end{align*}
$$

Here, $\mathcal{F}_{0}(\lambda)$ is the genus zero free energy in the orbifold frame, normalized in such a way that is expansion around $\lambda=0$ is given by

$$
\begin{equation*}
\mathcal{F}_{0}(\lambda)=-4 \pi^{2} \lambda^{2}\left(\log (2 \pi \lambda)-\frac{3}{2}-\log (4)\right)+\cdots \tag{4.21}
\end{equation*}
$$

Using modular properties of the periods and analytic continuation we find

$$
\begin{align*}
\Xi(\kappa, 4) & =\exp \left[\frac{1}{4} \tanh ^{-1}\left(\frac{\mathrm{e}^{\mu}}{4}\right)+\frac{\mu}{4}-\frac{1}{2 \pi^{2}}\left(\mathcal{F}_{0}(\lambda)-\lambda \partial_{\lambda} \mathcal{F}_{0}(\lambda)+\frac{1}{2} \lambda^{2} \partial_{\lambda}^{2} \mathcal{F}_{0}(\lambda)\right)\right] \\
& \times \exp \left[\mathcal{F}_{1}+F_{1}^{\mathrm{NS}}-\frac{\mathrm{i} \pi}{8}+\right] \vartheta_{3}\left(\bar{\xi}-\frac{1}{4}, 2 \bar{\tau}\right) \tag{4.22}
\end{align*}
$$

where

$$
\begin{equation*}
\bar{\tau}=-\frac{1}{8 \pi^{3} \mathrm{i}} \partial_{\lambda}^{2} \mathcal{F}_{0}(\lambda), \quad \bar{\xi}=\frac{\mathrm{i}}{4 \pi^{3}}\left(\lambda \partial_{\lambda}^{2} \mathcal{F}_{0}(\lambda)-\partial_{\lambda} \mathcal{F}_{0}(\lambda)\right) \tag{4.23}
\end{equation*}
$$

The genus one free energy in the orbifold frame is given by

$$
\begin{equation*}
\mathcal{F}_{1}=-\log \eta(2 \bar{\tau})-\frac{1}{2} \log 2 \tag{4.24}
\end{equation*}
$$

The small $\kappa$ expansion of (4.16) is now straightforward and one finds

$$
\begin{align*}
\mathcal{F}_{1}+F_{1}^{\mathrm{NS}}(t)-\frac{\mathrm{i} \pi}{8}+\frac{1}{4} \tanh ^{-1}\left(\frac{\mathrm{e}^{\mu}}{4}\right)+\frac{\mu}{4} & =\frac{\kappa}{16}+\frac{\kappa^{3}}{768}+\frac{\kappa^{4}}{32768}+\mathcal{O}\left(\kappa^{5}\right) \\
-\frac{1}{2 \pi^{2}}\left(\mathcal{F}_{0}(\lambda)-\lambda \partial_{\lambda} \mathcal{F}_{0}(\lambda)+\frac{1}{2} \lambda^{2} \partial_{\lambda}^{2} \mathcal{F}_{0}(\lambda)\right) & =-\frac{\kappa^{2}}{64 \pi^{2}}-\frac{\kappa^{4}}{3072 \pi^{2}}+\mathcal{O}\left(\kappa^{5}\right)  \tag{4.25}\\
\vartheta_{3}\left(\bar{\xi}-\frac{1}{4}, 2 \bar{\tau}\right) & =1-\frac{\kappa^{3}}{256 \pi}+\mathcal{O}\left(\kappa^{5}\right)
\end{align*}
$$

Hence we have

$$
\begin{equation*}
\Xi(\kappa, 4)=1+\frac{\kappa}{16}+\frac{\pi^{2}-8}{512 \pi^{2}} \kappa^{2}+\mathcal{O}\left(\kappa^{3}\right), \tag{4.26}
\end{equation*}
$$

which reproduces the computation of the very first $Z(N, 4)$ in [10]. Of course we can push the computation at higher order in $\kappa$ and reproduce all known $Z(N, 4)$ for higher $N$.

## 5. Factorization of the spectral determinant

Since the potential $V_{M}(x)$ appearing in the $\mathrm{ABJ}(\mathrm{M})$ spectral problem is even, the eigenfunctions $\phi_{n}(x)(2.17)$ can be classified according to their parity, as in ordinary one-dimensional quantummechanical problems. The even energy levels correspond to even eigenfunctions, while the odd energy levels correspond to odd eigenfunctions. Therefore, we can refine the spectral determinant (2.19) according to the parity of the eigenfunctions, and we define

$$
\begin{equation*}
\Xi_{+}(\kappa, k, M)=\prod_{n=0}^{\infty}\left(1+\kappa \mathrm{e}^{-E_{2 n}}\right), \quad \Xi_{-}(\kappa, k, M)=\prod_{n=0}^{\infty}\left(1+\kappa \mathrm{e}^{-E_{2 n+1}}\right) . \tag{5.1}
\end{equation*}
$$

If we introduce the operators with even/odd parity

$$
\begin{equation*}
\rho_{ \pm}\left(x_{1}, x_{2}\right)=\frac{\rho\left(x_{1}, x_{2}\right) \pm \rho\left(x_{1},-x_{2}\right)}{2}, \tag{5.2}
\end{equation*}
$$

the spectral determinants can be also written as

$$
\Xi_{ \pm}(\kappa, k, M)=\operatorname{det}\left(1+\kappa \hat{\rho}_{ \pm}\right)=\exp \left[-\sum_{n=1}^{\infty} \frac{(-\kappa)^{n}}{n} \operatorname{Tr} \hat{\rho}_{ \pm}^{n}\right] .
$$

Notice that by construction, one immediately gets

$$
\begin{equation*}
\Xi(\kappa, k, M)=\Xi_{+}(\kappa, k, M) \Xi_{-}(\kappa, k, M) . \tag{5.4}
\end{equation*}
$$

In this section we present an exact expression for the spectral determinants (5.1) in the case of $k=1, M=0$. We do not have a first principles derivation of such expressions, so we postulate a form that can be subsequently verified in detail. The expressions we propose are the following,

$$
\begin{align*}
& \Xi_{+}(\kappa, 1)=e^{J_{+}(\mu, 1)} \vartheta_{3}\left(\frac{1}{2}\left(\frac{\xi}{2}+\frac{5}{24}\right), \frac{\tau}{2}\right),  \tag{5.5}\\
& \Xi_{-}(\kappa, 1)=e^{J_{-}(\mu, 1)} \vartheta_{4}\left(\frac{1}{2}\left(\frac{\xi}{2}+\frac{5}{24}\right), \frac{\tau}{2}\right) .
\end{align*}
$$

In these expressions,

$$
\begin{align*}
& J_{+}(\mu, 1)=\frac{f_{0}(\mu)}{2}+f_{1}(\mu)+\frac{7}{16} \mu+A_{+}(1)+\frac{1}{8} \log \left(\frac{1+2 \sqrt{2} e^{-\mu}+4 e^{-2 \mu}}{1-2 \sqrt{2} e^{-\mu}+4 e^{-2 \mu}}\right), \\
& J_{-}(\mu, 1)=\frac{f_{0}(\mu)}{2}+f_{1}(\mu)-\frac{1}{16} \mu+A_{-}(1)-\frac{1}{8} \log \left(\frac{1+2 \sqrt{2} e^{-\mu}+4 e^{-2 \mu}}{1-2 \sqrt{2} e^{-\mu}+4 e^{-2 \mu}}\right), \tag{5.6}
\end{align*}
$$

where

$$
\begin{align*}
f_{0}(\mu) & =\frac{1}{16 \pi^{2}}\left(F_{0}(t)-t \partial_{t} F_{0}(t)+\frac{t^{2}}{2} \partial_{t}^{2} F_{0}(t)\right)  \tag{5.7}\\
f_{1}(\mu) & =F_{1}(t)+F_{1}^{\mathrm{NS}}(t)
\end{align*}
$$

and

$$
\begin{equation*}
A_{+}(1)=-\frac{\zeta(3)}{16 \pi^{2}}-\frac{3}{8} \log 2, \quad A_{-}(1)=-\frac{\zeta(3)}{16 \pi^{2}}+\frac{5}{8} \log 2, \tag{5.8}
\end{equation*}
$$

In the $k=1$ context we should use $t=4 \mu_{\text {eff }}$ and

$$
\begin{equation*}
F_{0}(t)=\frac{t^{3}}{6}+\sum_{d, w} n_{0}^{d} \frac{(-1)^{d w}}{w^{3}} \mathrm{e}^{-d w t} \tag{5.9}
\end{equation*}
$$

The standard and refined genus one free energy are given in terms of $z=\mathrm{e}^{-4 \mu}$ as

$$
\begin{align*}
F_{1}(t) & =-\frac{1}{12} \log [64 z(1+16 z)]-\frac{1}{2} \log \left(\frac{K(-16 z)}{\pi}\right),  \tag{5.10}\\
F_{1}^{\mathrm{NS}}(t) & =\frac{1}{12} \log z-\frac{1}{24} \log (1+16 z) .
\end{align*}
$$

As a first check of the proposal (5.5), let us show that the above expressions lead to the right quantization conditions, i.e. $\Xi_{ \pm}(\kappa, 1)$ vanish when $\mu=E_{2 n}+\mathrm{i} \pi, \mu=E_{2 n+1}+\mathrm{i} \pi$, respectively. Let us first recall the quantization condition for $k=1$ [17]

$$
\begin{equation*}
\frac{\xi}{2}-\frac{1}{24}=m+\frac{3}{4}, \quad m \geq 0 \tag{5.11}
\end{equation*}
$$

It is easy to see that the zeros of $\Xi_{ \pm}(E+\mathrm{i} \pi, 1)$ are given by

$$
\begin{align*}
& \frac{\xi}{2}-\frac{1}{24}=2 m+\frac{3}{4}, \quad m \geq 0  \tag{5.12}\\
& \frac{\xi}{2}-\frac{1}{24}=2(m+1)+\frac{3}{4}, \quad m \geq 0
\end{align*}
$$

which leads precisely to the odd and even energy levels for $k=1$ determined by (5.11).
As a second check, one can verify that

$$
\begin{equation*}
\Xi(\kappa, k)=\Xi_{+}(\kappa, k) \Xi_{-}(\kappa, k) . \tag{5.13}
\end{equation*}
$$

It is important to notice that the total grand potential $J(\mu, 1)$ differs from the sum

$$
\begin{equation*}
J_{+}(\mu, 1)+J_{-}(\mu, 1) \tag{5.14}
\end{equation*}
$$

by a term involving the genus one free energy. More precisely one has

$$
\begin{equation*}
J(\mu, 1)=J_{+}(\mu, 1)+J_{-}(\mu, 1)-F_{1}(t)-F_{1}^{\mathrm{NS}}(t) \tag{5.15}
\end{equation*}
$$

This additional contribution comes from the product of the two theta functions in $\Xi_{ \pm}(\kappa, 1) .{ }^{6}$

[^4]The factor $\vartheta_{4}(0, \tau)$ contributes to the modified grand potential $J(\mu, 1)$.

The third test concerns the large $\mu$ expansion for $J_{ \pm}(\mu, 1)$. If we write

$$
\begin{align*}
& J_{+}(\mu, 1)=\frac{\mu^{3}}{3 \pi^{2}}+\frac{7}{16} \mu+A_{+}(1)+J_{+}^{\mathrm{np}}(\mu, 1)  \tag{5.16}\\
& J_{-}(\mu, 1)=\frac{\mu^{3}}{3 \pi^{2}}-\frac{1}{16} \mu+A_{-}(1)+J_{-}^{\mathrm{np}}(\mu, 1)
\end{align*}
$$

we find from the above expressions,

$$
\begin{align*}
J_{+}^{\mathrm{np}}(\mu, 1)= & \frac{1}{\sqrt{2}} \mathrm{e}^{-\mu}-\frac{4}{3 \sqrt{2}} \mathrm{e}^{-3 \mu}+\frac{16 \mu^{2}+4 \mu+1}{8 \pi^{2}} \mathrm{e}^{-4 \mu}-\frac{16}{5 \sqrt{2}} \mathrm{e}^{-5 \mu}+\frac{64}{7 \sqrt{2}} \mathrm{e}^{-7 \mu} \\
& +\left[-\frac{52 \mu^{2}+\mu / 2+9 / 16}{4 \pi^{2}}+2\right] \mathrm{e}^{-8 \mu}+\frac{256}{9 \sqrt{2}} \mathrm{e}^{-9 \mu}+\mathcal{O}\left(\mathrm{e}^{-11 \mu}\right) \\
J_{-}^{\mathrm{np}}(\mu, 1)= & -\frac{1}{\sqrt{2}} \mathrm{e}^{-\mu}+\frac{4}{3 \sqrt{2}} \mathrm{e}^{-3 \mu}+\frac{16 \mu^{2}+4 \mu+1}{8 \pi^{2}} \mathrm{e}^{-4 \mu}+\frac{16}{5 \sqrt{2}} \mathrm{e}^{-5 \mu}-\frac{64}{7 \sqrt{2}} \mathrm{e}^{-7 \mu}  \tag{5.17}\\
& +\left[-\frac{52 \mu^{2}+\mu / 2+9 / 16}{4 \pi^{2}}+2\right] \mathrm{e}^{-8 \mu}-\frac{256}{9 \sqrt{2}} \mathrm{e}^{-9 \mu}+\mathcal{O}\left(\mathrm{e}^{-11 \mu}\right)
\end{align*}
$$

These expansions can be reproduced from the expressions for the spectral traces of $\rho_{ \pm}$, which were found in [8] up to $n=8$. The very first few values are given by

$$
\begin{array}{ll}
\operatorname{Tr} \rho_{+}=\frac{1}{4 \sqrt{2}}, & \operatorname{Tr} \rho_{-}=\frac{1}{4}-\frac{1}{4 \sqrt{2}} \\
\operatorname{Tr} \rho_{+}^{2}=\frac{1}{16 \pi}, & \operatorname{Tr} \rho_{-}^{2}=\frac{-3+\pi}{16 \pi}, \\
\operatorname{Tr} \rho_{+}^{3}=\frac{3-2 \sqrt{2}}{64}, & \operatorname{Tr} \rho_{-}^{3}=\frac{-12+\pi+2 \sqrt{2} \pi}{64 \pi}  \tag{5.18}\\
\operatorname{Tr} \rho_{+}^{4}=\frac{1}{512}\left(1-\frac{8}{\pi^{2}}\right), & \operatorname{Tr} \rho_{-}^{4}=-\frac{8+32 \pi-11 \pi^{2}}{512 \pi^{2}}
\end{array}
$$

In fact, we can relate our factorized spectral determinants $\Xi_{ \pm}$to these spectral traces directly, by expressing them in terms of orbifold quantities, like in [17]. As shown in appendix B in detail, one finds the following small $\kappa$ expansion:

$$
\begin{equation*}
\log \Xi_{ \pm}(\kappa, 1)=-\sum_{n=1}^{\infty} \frac{(-\kappa)^{n}}{n} \operatorname{Tr} \rho_{ \pm}^{n} \tag{5.19}
\end{equation*}
$$

which reproduces the exact values of $\operatorname{Tr} \rho_{ \pm}^{n}$. We have indeed computed $\operatorname{Tr} \rho_{ \pm}^{n}$ up to $n=44$ and compared them with the ones obtained from the orbifold expansion. The results show a perfect agreement.

## 6. Exact functional equations

In the previous section, we have considered the factorization of the spectral determinant. The reason why we focus on such a factorization is because the factorized spectral determinants $\Xi_{ \pm}(\kappa, k, M)$ have a very rich structure. In particular, these functions satisfy a set of exact functional equations as we will see in this section. A similar property has already been found in
the context of the spectral problem of certain ordinary differential equations (see [24, 25, 39] for example), where it indicates an unexpected connection to integrable models. Our result extends this type of properties to the spectral problem of the Fredholm integral equation (2.17). We hope that our findings may give a clue of a connection to some integrable systems.

### 6.1 Wronskian-like relations

We consider the spectral determinant (2.19) and the factorized ones (5.1). Remarkably, these functions satisfy the following beautiful equations for given $k$ and $M$ :

$$
\begin{align*}
& \mathrm{e}^{\frac{M}{2 k} \pi \mathrm{i}} \Xi_{+}(\mathrm{i} \kappa, k, M+1) \Xi_{-}(-\mathrm{i} \kappa, k, M-1) \\
& \quad-\mathrm{e}^{-\frac{M}{2 k} \pi \mathrm{i}} \Xi_{+}(-\mathrm{i} \kappa, k, M+1) \Xi_{-}(\mathrm{i} \kappa, k, M-1)=2 \mathrm{i} \sin \left(\frac{M \pi}{2 k}\right) \Xi(\kappa, k, M), \tag{6.1}
\end{align*}
$$

and

$$
\begin{align*}
& \mathrm{e}^{-\frac{M}{2 k} \pi \mathrm{i}} \Xi_{+}(\mathrm{i} \kappa, k, M-1) \Xi_{-}(-\mathrm{i} \kappa, k, M+1) \\
& \quad+\mathrm{e}^{\frac{M}{2 k} \pi \mathrm{i}} \Xi_{+}(-\mathrm{i} \kappa, k, M-1) \Xi_{-}(\mathrm{i} \kappa, k, M+1)=2 \cos \left(\frac{M \pi}{2 k}\right) \Xi(\kappa, k, M) \tag{6.2}
\end{align*}
$$

The equation (6.1) is quite similar to the so-called quantum Wronskian relation [40]. We do not have a general proof for these relations but we can test them by computing the spectrum and its spectral traces from the quantization condition (3.19), and by doing small $\kappa$ expansions of the spectral determinants. This can be done for various values of the coupling $k$. Such tests strongly suggest that the functional equations (6.1) and (6.2) are widely valid not only for integral values of $k$, but also for non-integral values.

In particular, if $k$ is an integer, the equations (6.1) and (6.2) are essentially equivalent due to the Seiberg-like duality of ABJ theories [2]. This duality relates the partition function for $(k, M)$ to the one for $(k, k-M)$. In terms of the spectral determinants, it simply says that

$$
\begin{equation*}
\Xi(\kappa, k, M)=\Xi(\kappa, k, k-M), \quad \Xi_{ \pm}(\kappa, k, M)=\Xi_{ \pm}(\kappa, k, k-M) . \tag{6.3}
\end{equation*}
$$

If one considers the case $M=k-m$ in (6.2), one gets

$$
\begin{align*}
& -\mathrm{i} \frac{m}{2 k} \pi \mathrm{i} \Xi_{+}(\mathrm{i} \kappa, k, k-m-1) \Xi_{-}(-\mathrm{i} \kappa, k, k-m+1) \\
& +\mathrm{i} \mathrm{e}^{-\frac{m}{2 k} \mathrm{i} \mathrm{i}} \Xi_{+}(-\mathrm{i} \kappa, k, k-m-1) \Xi_{-}(\mathrm{i} \kappa, k, k-m+1)=2 \sin \left(\frac{m \pi}{2 k}\right) \Xi(\kappa, k, k-m) . \tag{6.4}
\end{align*}
$$

Using the Seiberg-like duality $\Xi_{ \pm}(\kappa, k, k-m \pm 1)=\Xi_{ \pm}(\kappa, k, m \mp 1)$, it is easy to see that this equation is equivalent to (6.1) for $M=m$. We note that for non-integral $k$, the equations (6.1) and (6.2) are independent.

Moreover we conjecture another functional equation, which associates $\Xi_{ \pm}(\kappa, k, 1)$ to $\Xi(\kappa, k)$.

$$
\begin{equation*}
\Xi_{+}(\mathrm{i} \kappa, k, 1) \Xi_{+}(-\mathrm{i} \kappa, k, 1)+\frac{\kappa}{4 k} \Xi_{-}(\mathrm{i} \kappa, k, 1) \Xi_{-}(-\mathrm{i} \kappa, k, 1)=\Xi(\kappa, k), \quad \forall k . \tag{6.5}
\end{equation*}
$$

As before the identity (6.5) can be checked by computing the spectrum and by doing a small $\kappa$ expansion. We tested this relation for various values of the coupling $k$, and conjecture that it is valid for any $k$.

Let us comment on a consequence of these functional equations. For odd $k$, there are $k+1$ independent functions $\Xi_{ \pm}(\kappa, k, M),\left(M=0, \ldots, \frac{k-1}{2}\right)$ due to the Seiberg-like duality. A simple
counting shows that the functional equation (6.1) (or equivalently (6.2)) gives $k-1$ independent constraints. This means that if we know the two functions $\Xi_{ \pm}(\kappa, k)$ in the ABJM theory, all the other functions $\Xi_{ \pm}(\kappa, k, M)$ in ABJ theory are determined by the functional equations. Similarly, for even $k$, the functional equation (6.1) gives $k-1$ independent constraints among the $k+2$ independent functions $\Xi_{ \pm}(\kappa, k, M)\left(M=0, \ldots, \frac{k}{2}\right)$. In this case, the equation (6.1) only does not determine the ABJ spectral determinants from the ABJM ones. Since we have the additional relation (6.5), one might expect that combining these equations, the ABJ spectral determinants are fixed, as for odd $k$. However, this is not the case. One can check that the equations (6.1) and (6.5) are not sufficient to determine all the ABJ spectral determinants only from the ABJM ones. We need more information for even $k .{ }^{7}$

In the rest of this subsection we exploit our exact solution (5.5) to further test (6.5) in the case of $k=1$. More precisely, we are interested in studying the following identity:

$$
\begin{equation*}
\Xi_{+}(\mathrm{i} \kappa, 1) \Xi_{+}(-\mathrm{i} \kappa, 1)+\frac{\kappa}{4} \Xi_{-}(\mathrm{i} \kappa, 1) \Xi_{-}(-\mathrm{i} \kappa, 1)=\Xi(\kappa, 1) . \tag{6.6}
\end{equation*}
$$

Notice that under $\kappa \rightarrow \pm \mathrm{i} \kappa$, the chemical potential change according to

$$
\begin{equation*}
\mu \rightarrow \mu \pm \frac{\pi \mathrm{i}}{2} \tag{6.7}
\end{equation*}
$$

Starting from (5.7) it is easy to see that

$$
\begin{align*}
& f_{0}\left(\mu+\frac{\pi \mathrm{i}}{2}\right)+f_{0}\left(\mu-\frac{\pi \mathrm{i}}{2}\right)=2 f_{0}(\mu)-\frac{1}{4} \partial_{t}^{2} F_{0}(t),  \tag{6.8}\\
& f_{1}\left(\mu+\frac{\pi \mathrm{i}}{2}\right)+f_{1}\left(\mu-\frac{\pi \mathrm{i}}{2}\right)=2 f_{1}(\mu)
\end{align*}
$$

It follows that

$$
\begin{align*}
& J_{+}\left(\mu+\frac{\pi \mathrm{i}}{2}\right)+J_{+}\left(\mu-\frac{\pi \mathrm{i}}{2}\right)=J(\mu, 1)+F_{1}(t)+F_{1}^{\mathrm{NS}}(t)-\log 2+\frac{\mu}{2}-\frac{1}{8} \partial_{t}^{2} F_{0}(t), \\
& J_{-}\left(\mu+\frac{\pi \mathrm{i}}{2}\right)+J_{-}\left(\mu-\frac{\pi \mathrm{i}}{2}\right)=J(\mu, 1)+F_{1}(t)+F_{1}^{\mathrm{NS}}(t)+\log 2-\frac{\mu}{2}-\frac{1}{8} \partial_{t}^{2} F_{0}(t) . \tag{6.9}
\end{align*}
$$

Similarly one has

$$
\begin{equation*}
\xi\left(E \pm \frac{\pi \mathrm{i}}{2}\right)=\xi(E) \pm \frac{\tau}{2}-1, \quad \tau\left(E \pm \frac{\pi \mathrm{i}}{2}\right)=\tau(E) \mp 4 . \tag{6.10}
\end{equation*}
$$

Therefore,

$$
\begin{align*}
& \vartheta_{3}\left(\frac{1}{2}\left(\frac{\xi\left(E \pm \frac{\pi \mathrm{i}}{2}\right)}{2}+\frac{5}{24}\right), \frac{\tau\left(E \pm \frac{\pi \mathrm{i}}{2}\right)}{2}\right)=\vartheta_{3}\left(\frac{\xi}{4}-\frac{7}{48} \pm \frac{\tau}{8}, \frac{\tau}{2}\right), \\
& \vartheta_{3}\left(\frac{1}{2}\left(\frac{\xi\left(E \pm \frac{\pi \mathrm{i}}{2}\right)}{2}-\frac{19}{24}\right), \frac{\tau\left(E \pm \frac{\pi \mathrm{i}}{2}\right)}{2}\right)=\vartheta_{4}\left(\frac{\xi}{4}-\frac{7}{48} \pm \frac{\tau}{8}, \frac{\tau}{2}\right) . \tag{6.11}
\end{align*}
$$

[^5]Using the identities

$$
\begin{align*}
& \vartheta_{3}(x+y, \tau) \vartheta_{3}(x-y, \tau)=\vartheta_{3}(2 x, 2 \tau) \vartheta_{3}(2 y, 2 \tau)+\vartheta_{2}(2 x, 2 \tau) \vartheta_{2}(2 y, 2 \tau) \\
& \vartheta_{4}(x+y, \tau) \vartheta_{4}(x-y, \tau)=\vartheta_{3}(2 x, 2 \tau) \vartheta_{3}(2 y, 2 \tau)-\vartheta_{2}(2 x, 2 \tau) \vartheta_{2}(2 y, 2 \tau) \tag{6.12}
\end{align*}
$$

we get

$$
\begin{align*}
& \vartheta_{3}\left(\frac{\xi}{4}-\frac{7}{48}+\frac{\tau}{8}, \frac{\tau}{2}\right) \vartheta_{3}\left(\frac{\xi}{4}-\frac{7}{48}-\frac{\tau}{8}, \frac{\tau}{2}\right)= \\
& \vartheta_{3}\left(\frac{\xi}{2}-\frac{7}{24}, \tau\right) \vartheta_{3}\left(\frac{\tau}{4}, \tau\right)+\vartheta_{2}\left(\frac{\xi}{2}-\frac{7}{24}, \tau\right) \vartheta_{2}\left(\frac{\tau}{4}, \tau\right) \tag{6.13}
\end{align*}
$$

Similarly

$$
\begin{align*}
& \vartheta_{4}\left(\frac{\xi}{4}-\frac{7}{48}+\frac{\tau}{8}, \frac{\tau}{2}\right) \vartheta_{4}\left(\frac{\xi}{4}-\frac{7}{48}-\frac{\tau}{8}, \frac{\tau}{2}\right)=  \tag{6.14}\\
& \vartheta_{3}\left(\frac{\xi}{2}-\frac{7}{24}, \tau\right) \vartheta_{3}\left(\frac{\tau}{4}, \tau\right)-\vartheta_{2}\left(\frac{\xi}{2}-\frac{7}{24}, \tau\right) \vartheta_{2}\left(\frac{\tau}{4}, \tau\right)
\end{align*}
$$

It follows that

$$
\begin{align*}
\Xi_{+}(\mathrm{i} \kappa, 1) \Xi_{+}(-\mathrm{i} \kappa, 1) & =\frac{1}{2} \Xi(\kappa, 1)\left(\vartheta_{3}\left(\frac{\tau}{4}, \tau\right)+\frac{\vartheta_{2}(v, \tau)}{\vartheta_{3}(v, \tau)} \vartheta_{2}\left(\frac{\tau}{4}, \tau\right)\right) \\
& \times \exp \left[F_{1}+F_{1}^{\mathrm{NS}}+\frac{\mu}{2}-\frac{1}{8} \partial_{t}^{2} F_{0}\right] \\
\Xi_{-}(\mathrm{i} \kappa, 1) \Xi_{-}(-\mathrm{i} \kappa, 1) & =2 \Xi(\kappa, 1)\left(\vartheta_{3}\left(\frac{\tau}{4}, \tau\right)-\frac{\vartheta_{2}(v, \tau)}{\vartheta_{3}(v, \tau)} \vartheta_{2}\left(\frac{\tau}{4}, \tau\right)\right)  \tag{6.15}\\
& \times \exp \left[F_{1}+F_{1}^{\mathrm{NS}}-\frac{\mu}{2}-\frac{1}{8} \partial_{t}^{2} F_{0}\right] .
\end{align*}
$$

By using the above expression one can write (6.6) as

$$
\begin{equation*}
F_{1}(t)+F_{1}^{\mathrm{NS}}(t)+\frac{\mu}{2}-\frac{1}{8} \partial_{t}^{2} F_{0}(t)=-\log \vartheta_{3}\left(\frac{\tau}{4}, \tau\right) \tag{6.16}
\end{equation*}
$$

We can checked this identity order by order in the large $\kappa$ expansion and numerically for finite value of $\kappa$. It would be interesting to confirm the functional equations (6.1) and (6.5) at $k=2$ in the similar way by using the exact solutions.

### 6.2 Relations among different levels

In addition to the general relations found in the previous subsection, there are some accidental relations among the spectral determinants for different values of $k$. We find that the following relations hold:

$$
\begin{align*}
\Xi(\kappa, 4) & =\Xi_{+}(\kappa, 2,1), \\
\Xi(\kappa, 4,1) & =\Xi_{-}(\kappa, 2)=\Xi(\mathrm{i} \sqrt{\kappa}, 1) \Xi(-\mathrm{i} \sqrt{\kappa}, 1), \\
\Xi(\kappa, 4,2) & =\Xi_{-}(\kappa, 2,1),  \tag{6.17}\\
\Xi(\kappa, 8,2) & =\Xi_{-}(\kappa, 4,2) .
\end{align*}
$$

These relations can be checked as follows. Let us recall the relation (4.19) for the quantum volumes. Considering the quantization condition, the relation (4.19) implies the equality

$$
\begin{equation*}
E_{n}(k=4)=E_{2 n}(k=2, M=1), \quad n=0,1,2, \ldots \tag{6.18}
\end{equation*}
$$

Thus we immediately find the first line in (6.17) by definition. Similarly, we find

$$
\begin{align*}
& \operatorname{Vol}(E, 4,1)=\operatorname{Vol}(E, 2)-4 \pi^{2}=4 \operatorname{Vol}\left(\frac{E}{2}, 1\right), \\
& \operatorname{Vol}(E, 4,2)=\operatorname{Vol}(E, 2,1)-4 \pi^{2},  \tag{6.19}\\
& \operatorname{Vol}(E, 8,2)=\operatorname{Vol}(E, 4,2)-8 \pi^{2},
\end{align*}
$$

From these relations, we find the relations on the energy levels, and then get (6.17). As a further test, one can check the equalities around $\kappa=0$. For example, the first and third lines in (6.17) show that the spectral determinant for $k=2, M=1$ splits into two part

$$
\begin{equation*}
\Xi(\kappa, 2,1)=\Xi(\kappa, 4) \Xi(\kappa, 4,2) . \tag{6.20}
\end{equation*}
$$

One can check this equation by substituting the exact values of the partition function computed in $[10,23]$. Notice that we already know the exact spectral determinants for $k=1,2,4, M=0$ as well as for $k=2, M=1$ [17]. These data fix $\Xi(\kappa, 4,1)$ and $\Xi(\kappa, 4,2)$ through (6.17). For example, using (6.20), we get

$$
\begin{equation*}
\Xi(\kappa, 4,2)=\frac{\Xi(\kappa, 2,1)}{\Xi(\kappa, 4)}=\mathrm{e}^{J(\mu, 2,1)-J(\mu, 4)} \frac{\vartheta_{3}\left(\xi-\frac{\tau}{4}-\frac{7}{12}, \tau\right)}{\vartheta_{3}\left(\frac{1}{2}\left(\xi-\frac{\tau}{4}-\frac{1}{12}\right), \frac{\tau}{2}\right)} . \tag{6.21}
\end{equation*}
$$

A simple calculation shows that this is written as

$$
\begin{equation*}
\Xi(\kappa, 4,2)=\mathrm{e}^{J(\mu, 4,2)} \vartheta_{4}\left(\frac{1}{2}\left(\xi-\frac{\tau}{4}-\frac{1}{12}\right), \frac{\tau}{2}\right), \tag{6.22}
\end{equation*}
$$

where

$$
\begin{equation*}
J(\mu, 4,2)=J(\mu, 2,1)-J(\mu, 4)+F_{1}(t)+F_{1}^{\mathrm{NS}}(t), \quad t=2 \mu_{\mathrm{eff}} . \tag{6.23}
\end{equation*}
$$

Similarly, $\Xi(\kappa, 4,1)$ is fixed by the second line in (6.17) by using the exact expreesion for $\Xi(\kappa, 1)$ in [17].

## 7. Conclusion

In this paper we studied the spectral problem appearing in the Fermi gas formulation of $\operatorname{ABJ}(\mathrm{M})$ theory. By generalizing the recent study of maximally supersymmetric $\operatorname{ABJ}(\mathrm{M})$ theories in [17], we found an exact expression for the spectral determinant in terms of a generalized theta function, and then we computed the quantum volume by looking at the zeros of this spectral determinant. This method has the advantage of overcoming many technical difficulties encountered in [15, 16] and leads to an exact quantization condition for the spectrum. Our result also shows that the quantization conditions conjectured in $[15,16]$ are only approximate, although they become exact in the maximally supersymmetric cases. Our quantization condition agrees with a recent numerical analysis in [18], and we tested it against a high precision, numerical computation of the spectrum. As an application of our results, we also conjectured some functional equations for the spectral determinants. Note that the formalism we used in this paper is very powerful and completely general. As explained in [20], it also has important applications in topological string theory.

This work can be extended in many ways. First of all, it would be interesting to understand the structure of the spectral determinant in other cases with $\mathcal{N}=6$ supersymmetry. In the ABJM
theory with $k=4$, we could resum the all-genus expansion of the modified grand potential in order to understand the expansion of the spectral determinant at small fugacity. It would be very interesting to understand if this resummation can be done in general. This will probably require a better understanding of the modular properties of the modified grand potential and of the generalized theta function at finite $k$.

Another avenue to explore is the generalization of these results to other Chern-Simonsmatter theories. This is not completely straightforward: although our results for the spectral determinant and the quantization conditions are quite general and can be easily extended to other models, our detailed computations rely on a detailed knowledge of the modified grand potential, which so far has been only achieved for $\mathrm{ABJ}(\mathrm{M})$ theory. Nevertheless, we hope that the results obtained in this paper will be useful to further understand the non-perturbative structure of Chern-Simons-matter theories and their large $N$ duals.

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## A. The finite part of $\tilde{J}_{b}\left(\mu_{\text {eff }}, k, M\right)$

We want to show that, when $k$ is integer, the coefficient $\widetilde{b}_{\ell}(k)$ defined in (2.40) has no finite part. More precisely, let us consider the expansion of $\widetilde{b}_{\ell}(k)$ around an integer $k_{0}$. Since $\widetilde{b}_{\ell}(k)$ has a simple pole there, one has

$$
\begin{equation*}
\widetilde{b}_{\ell}(k)=\frac{\tilde{b}_{\ell}^{-1}\left(k_{0}\right)}{k-k_{0}}+\tilde{b}_{\ell}^{0}\left(k_{0}\right)+\mathcal{O}\left(k-k_{0}\right) . \tag{A.1}
\end{equation*}
$$

The finite part of $\widetilde{b}_{\ell}(k)$ as $k \rightarrow k_{0}$ is $\tilde{b}_{\ell}^{0}\left(k_{0}\right)$. We want to show that $\tilde{b}_{\ell}^{0}\left(k_{0}\right)=0$. Let us start by looking at the case in which $k_{0}$ is even. From (2.40) one can see that the finite part is

$$
\begin{equation*}
\tilde{b}_{\ell}^{0}\left(k_{0}\right)=-\frac{\ell}{2 \pi} \sum_{j_{L}, j_{R}} \sum_{\ell=d w} \sum_{d_{1}+d_{2}=d} N_{j_{L}, j_{R}}^{d_{1}, d_{2}} \frac{\mathrm{i}_{L} m_{L}\left(d_{1}-d_{2}\right) \cos \left(\frac{1}{2} \pi k_{0} w\left(d_{1}+d_{2}+m_{L}+m_{R}+1\right)\right)}{2 \pi w^{2}} \tag{A.2}
\end{equation*}
$$

where we have denoted

$$
\begin{equation*}
m_{L, R}=2 j_{L, R}+1 \tag{A.3}
\end{equation*}
$$

Since the BPS invariants of local $\mathbb{P}^{1} \times \mathbb{P}^{1}, N_{j_{L}, j_{R}}^{d_{1}, d_{2}}$, are symmetric under the exchange $d_{1} \leftrightarrow d_{2}$, the above quantity vanishes. When $k_{0}$ is odd, it is convenient to split the sum over $w$ in (2.40) into even $w$ and odd $w$. The sum over even $w$ can be reduced to (A.2) and therefore vanishes. The sum over odd $w$ gives instead a contribution of the form

$$
\begin{equation*}
-\frac{\ell}{2 \pi} \sum_{j_{L}, j_{R}} \sum_{\ell=d w} \sum_{d_{1}+d_{2}=d} N_{j_{L}, j_{R}}^{d_{1}, d_{2}} \frac{e^{\frac{1}{2} i \pi\left(k\left(d_{1} w-d_{2} w-1\right)-w\right)} \sin \left(\frac{1}{2} \pi k_{0} m_{L} w\right) \sin \left(\frac{1}{2} \pi k_{0} w m_{R}\right)}{2 \pi w^{2}} . \tag{A.4}
\end{equation*}
$$

For local $\mathbb{P}^{1} \times \mathbb{P}^{1}$, the non-vanishing BPS invariants are such that

$$
\begin{equation*}
m_{L}+m_{R}=2 n+1, \quad n \in \mathbb{Z} \tag{A.5}
\end{equation*}
$$

It follows that

$$
\begin{equation*}
\sin \left(\frac{1}{2} \pi k_{0} m_{L} w\right) \sin \left(\frac{1}{2} \pi k_{0} w m_{R}\right)=0 . \tag{A.6}
\end{equation*}
$$

Hence, (A.4) also vanishes. This proves our statement.

## B. The spectral determinant at the orbifold point

Let us compute the expansions of the spectral determinants $\Xi_{ \pm}(\kappa, 1)$, given in (5.5), around the orbifold point $\kappa=0$. One finds

$$
\begin{align*}
\Xi_{+}(\kappa, 1) & =\exp \left[\mathcal{F}_{1}+F_{1}^{\mathrm{NS}}-\frac{1}{8 \pi^{2}}\left(\mathcal{F}_{0}(\lambda)-\lambda \partial_{\lambda} \mathcal{F}_{0}(\lambda)+\frac{\lambda^{2}}{2} \partial_{\lambda}^{2} \mathcal{F}_{0}(\lambda)\right)\right] \\
& \times \exp \left[\frac{1}{8} \log \left(\frac{4+2 \sqrt{2} \kappa+\kappa^{2}}{4-2 \sqrt{2} \kappa+\kappa^{2}}\right)+\frac{7 \log (\kappa)}{16}-\frac{15}{8} \log (2)\right] \bar{\theta}^{(1)}(\xi, \tau), \\
\Xi_{-}(\kappa, 1) & =\exp \left[\mathcal{F}_{1}+F_{1}^{\mathrm{NS}}-\frac{1}{8 \pi^{2}}\left(\mathcal{F}_{0}(\lambda)-\lambda \partial_{\lambda} \mathcal{F}_{0}(\lambda)+\frac{\lambda^{2}}{2} \partial_{\lambda}^{2} \mathcal{F}_{0}(\lambda)\right)\right]  \tag{B.1}\\
& \times \exp \left[-\frac{1}{8} \log \left(\frac{4+2 \sqrt{2} \kappa+\kappa^{2}}{4-2 \sqrt{2} \kappa+\kappa^{2}}\right)-\frac{\log (\kappa)}{16}-\frac{7}{8} \log (2)\right] \bar{\theta}^{(2)}(\xi, \tau)
\end{align*}
$$

where $F_{1}^{\mathrm{NS}}$ is given in (5.10) and $\mathcal{F}_{1}, \mathcal{F}_{0}$ can be obtained from (4.20), (4.24) by replacing

$$
\begin{equation*}
\kappa \rightarrow-\mathrm{i} \kappa^{2} . \tag{B.2}
\end{equation*}
$$

In these equations, we defined

$$
\begin{align*}
\bar{\theta}^{(1)}(\xi, \tau)= & {\left[\mathrm{e}^{\frac{\mathrm{i} \pi}{4}} \vartheta_{2}\left(\frac{\xi}{8 \tau}-\frac{1}{96 \tau}+\frac{1}{4}, 8 \tau\right)+\vartheta_{2}\left(\frac{\xi}{8 \tau}-\frac{1}{96 \tau}, 8 \tau\right)\right.} \\
& \left.-\mathrm{ie}^{\frac{\mathrm{i} \pi}{4}} \vartheta_{1}\left(\frac{\xi}{8 \tau}-\frac{1}{96 \tau}+\frac{1}{4}, 8 \tau\right)-\mathrm{i} \vartheta_{1}\left(\frac{\xi}{8 \tau}-\frac{1}{96 \tau}, 8 \tau\right)\right],  \tag{B.3}\\
\bar{\theta}^{(2)}(\xi, \tau)= & {\left[\mathrm{e}^{\frac{-3 i \pi}{4}} \vartheta_{2}\left(\frac{\xi}{8 \tau}-\frac{1}{96 \tau}+\frac{1}{4}, 8 \tau\right)+\vartheta_{2}\left(\frac{\xi}{8 \tau}-\frac{1}{96 \tau}, 8 \tau\right)\right.} \\
& \left.-\mathrm{ie} \frac{-\frac{3 i \pi}{4}}{} \vartheta_{1}\left(\frac{\xi}{8 \tau}-\frac{1}{96 \tau}+\frac{1}{4}, 8 \tau\right)-\mathrm{i} \vartheta_{1}\left(\frac{\xi}{8 \tau}-\frac{1}{96 \tau}, 8 \tau\right)\right] .
\end{align*}
$$

Let us look at the series expansion of the terms appearing in $\Xi_{+}(\kappa, 1)$. We have:

$$
\begin{align*}
\mathcal{F}_{1}+F_{1}^{\mathrm{NS}}+\frac{1}{8} \log \left(\frac{4+2 \sqrt{2} \kappa+\kappa^{2}}{4-2 \sqrt{2} \kappa+\kappa^{2}}\right)+\frac{\log (\kappa)}{2} & =\frac{\kappa}{4 \sqrt{2}}-\frac{\kappa^{3}}{48 \sqrt{2}}+\mathcal{O}\left(\kappa^{5}\right) \\
\frac{1}{8 \pi^{2}}\left(-\mathcal{F}_{0}(\lambda)+\lambda \partial_{\lambda} \mathcal{F}_{0}(\lambda)-\frac{\lambda^{2}}{2} \partial_{\lambda}^{2} \mathcal{F}_{0}(\lambda)\right) & =\frac{\kappa^{4}}{256 \pi^{2}}-\frac{\kappa^{8}}{12288 \pi^{2}}+\mathcal{O}\left(\kappa^{12}\right)  \tag{B.4}\\
\exp \left[-\frac{15}{8} \log (2)-\frac{\log (\kappa)}{16}\right] \bar{\theta}^{(1)}(\xi, \tau) & =1-\frac{\kappa^{2}}{32 \pi}+\frac{\kappa^{3}}{64}-\frac{\pi^{2}-1}{2048 \pi^{2}} \kappa^{4}+\mathcal{O}\left(\kappa^{5}\right) .
\end{align*}
$$

This leads to

$$
\begin{equation*}
\log \Xi_{+}(\kappa, 1)=\frac{\kappa}{4 \sqrt{2}}-\frac{\kappa^{2}}{32 \pi}+\frac{3-2 \sqrt{2}}{192} \kappa^{3}+\frac{1}{2048}\left(\frac{8}{\pi^{2}}-1\right) \kappa^{4}+\mathcal{O}\left(\kappa^{5}\right) \tag{B.5}
\end{equation*}
$$

A similar computation holds for $\Xi_{-}(\kappa, 1)$.

## C. Numerical calculation of the spectrum

In this Appendix we explain how to compute numerically the first two energy levels of the spectrum (2.17), with high precision. We introduce the following two functions:

$$
\begin{align*}
& \phi_{\ell}^{+}(x)=\frac{1}{\cosh \left(\frac{x}{2 k}\right)} \int_{-\infty}^{\infty} \frac{d x^{\prime}}{2 \pi k} \frac{\cosh \left(\frac{x^{\prime}}{2 k}\right)}{2 \cosh \left(\frac{x-x^{\prime}}{2 k}\right)} V_{M}\left(x^{\prime}\right) \phi_{\ell-1}^{+}\left(x^{\prime}\right),  \tag{C.1}\\
& \phi_{\ell}^{-}(x)=\frac{1}{\cosh \left(\frac{x}{2 k}\right)} \int_{-\infty}^{\infty} \frac{d x^{\prime}}{2 \pi k} \frac{\sinh \left(\frac{x^{\prime}}{2 k}\right) \tanh \left(\frac{x^{\prime}}{2 k}\right)}{2 \cosh \left(\frac{x-x^{\prime}}{2 k}\right)} V_{M}\left(x^{\prime}\right) \phi_{\ell-1}^{-}\left(x^{\prime}\right),
\end{align*}
$$

where $V_{M}(x)$ is defined by (2.11). As shown in $[8,10,23]$, these functions are building blocks to construct the matrix elements $\rho_{ \pm}^{\ell}(x, y)$, and one can compute the spectral traces $\operatorname{Tr} \rho_{ \pm}^{\ell}$ from $\phi_{ \pm}^{\ell}(x)$. Recalling that the traces $\operatorname{Tr} \rho_{ \pm}^{\ell}$ are also given by

$$
\begin{equation*}
\operatorname{Tr} \rho_{+}^{\ell}=\sum_{n=0}^{\infty} \mathrm{e}^{-\ell E_{2 n}}, \quad \operatorname{Tr} \rho_{-}^{\ell}=\sum_{n=0}^{\infty} \mathrm{e}^{-\ell E_{2 n+1}} \tag{C.2}
\end{equation*}
$$

one can compute the first two energies levels from

$$
\begin{equation*}
\mathrm{e}^{-E_{0}}=\lim _{\ell \rightarrow \infty} \frac{\operatorname{Tr} \rho_{+}^{\ell}}{\operatorname{Tr} \rho_{+}^{\ell-1}}, \quad \mathrm{e}^{-E_{1}}=\lim _{\ell \rightarrow \infty} \frac{\operatorname{Tr} \rho_{-}^{\ell}}{\operatorname{Tr} \rho_{-}^{\ell-1}}, \tag{C.3}
\end{equation*}
$$

where we have used that

$$
\begin{equation*}
\mathrm{e}^{-\ell E_{0}}>\mathrm{e}^{-\ell E_{1}}>\mathrm{e}^{-\ell E_{2}}>\ldots \tag{C.4}
\end{equation*}
$$

From a practical point of view, there is another simpler way to compute $E_{0}$ and $E_{1}$. This way is based on the observation that the two functions $\phi_{ \pm}^{\ell}(0)$ already contain all the information of the spectrum. Indeed we have

$$
\begin{align*}
& \phi_{\ell}^{+}(0)=C_{0} \mathrm{e}^{-\ell E_{0}}+C_{2} \mathrm{e}^{-\ell E_{2}}+C_{4} \mathrm{e}^{-\ell E_{4}}+\ldots, \\
& \phi_{\ell}^{-}(0)=C_{1} \mathrm{e}^{-\ell E_{1}}+C_{3} \mathrm{e}^{-\ell E_{3}}+C_{5} \mathrm{e}^{-\ell E_{5}}+\ldots, \tag{C.5}
\end{align*}
$$

where $C_{n}$ are constant coefficients. Using this observation, one immediately finds

$$
\begin{equation*}
\mathrm{e}^{-E_{0}}=\lim _{\ell \rightarrow \infty} \frac{\phi_{\ell}^{+}(0)}{\phi_{\ell-1}^{+}(0)}, \quad \mathrm{e}^{-E_{1}}=\lim _{\ell \rightarrow \infty} \frac{\phi_{\ell}^{-}(0)}{\phi_{\ell-1}^{-}(0)} . \tag{C.6}
\end{equation*}
$$

These expressions are technically useful because we do not need to perform any integral over $x$. The integral equations (C.1) can be solved numerically for given $k$ and $M$ with high precision. Once we get $\phi_{ \pm}^{\ell}(x)$ up to some values of $\ell$, we can estimate the energies by (C.6). This method is very powerful to compute $E_{0}$ and $E_{1}$ numerically. In fact, we have checked that this method reproduces the energies computed from the exact quantization condition in [17] for $(k, M)=$ $(1,0),(2,0),(2,1)$ with very high (at least 100 -digit) precision.

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[^0]:    ${ }^{1}$ Note that our spectral determinant involves the spectrum of $\mathrm{e}^{-\hat{H}}$, rather than the spectrum of $\hat{H}$ itself, as in other definitions of spectral determinants $[24,25,30]$. Our definition is more convenient from the point of convergence properties, since it does not require regularization.

[^1]:    ${ }^{2}$ The coefficients $d_{m}(k, M)$ differ from those in [16] by a factor $(-1)^{m}$.

[^2]:    ${ }^{3}$ As mentioned in [27], the total ABJ partition function (2.4) vanishes for $k<M$. This is because the ChernSimons factor $Z_{\mathrm{CS}}(M, k)$ becomes zero in this regime. However the normalized partition function $\hat{Z}(N, k ; M)$ defined by (2.9) is still non-zero even for $k<M$. Therefore one can consider the classical limit once going to the Fermi gas system.
    ${ }^{4}$ Strictly speaking $\rho_{\mathrm{cl}}(x, p)$ is not fully classical because $U(x, M)$ still depends on the Planck constant when $M>0$.

[^3]:    ${ }^{5}$ As noted in [30], p. 672, "the smart way to determine the eigenvalues of linear operators is by determining zeros of their spectral determinants."

[^4]:    ${ }^{6}$ We have used the identity

    $$
    \vartheta_{3}\left(\frac{v}{2}, \frac{\tau}{2}\right) \vartheta_{4}\left(\frac{v}{2}, \frac{\tau}{2}\right)=\vartheta_{4}(0, \tau) \vartheta_{4}(v, \tau)
    $$

[^5]:    ${ }^{7}$ If we additionally give the traces of the odd powers of $\rho_{+}$for $M=1$, for example, then the other traces in ABJ theories are fixed.

