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Parametrization of the q^2 Dependence of $\gamma_v p$ Total Cross Sections
in the Resonance Region

F.W. Brasse, W. Flauger, J. Gayler, S.P. Goel*),
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ABSTRACT

All existing data on $\gamma_v p$ total cross sections in the resonance region are fitted in the absolute value of the three momentum transfers $|\vec{q}|$ independently for small bins of W in the range $1.11 \leq W \leq 1.99$ GeV. The data are divided into three ranges of the polarization ϵ : $\epsilon \geq 0.9$, $0.9 > \epsilon > 0.6$, $\epsilon \leq 0.6$.

Taking into account statistical and possible systematic errors, a comparison of the fits for different ranges of ϵ indicate that longitudinal contributions to the cross sections in the resonance region are in general small. In the range $1.3 < W < 1.5$ GeV for $q^2 > 2$ GeV^2 and above $W = 1.6$ GeV for $q^2 > 0.5$ GeV^2 up to 20 % contribution of σ_ℓ is possible.

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I. INTRODUCTION

The total cross section Σ for the absorption of virtual photons on protons in the region of the main nucleon resonances has been determined in many inelastic electron-proton scattering experiments¹⁻²³⁾ where only the scattered electron is detected. Σ is written as

$$\Sigma(q^2, W, \epsilon) = \frac{1}{\Gamma_t} \frac{d^2\sigma}{d\Omega dE} = \sigma_t(W, q^2) + \epsilon \sigma_\ell(W, q^2) \quad (1)$$

where, as usual³⁾,

q^2 is the four-momentum transfer squared, W the mass of the outgoing hadronic system, Γ_t the flux of transverse polarized virtual photons, ϵ the degree of polarization of the virtual photons, and σ_t and σ_ℓ are the absorption cross section for transverse and longitudinal polarized virtual photons, respectively.

A parametrization of the q^2 dependence of the total cross section by a fitting procedure applied to all existing data allows one to study its behaviour with the best available statistics. As the total cross section Σ consists of two parts σ_t and σ_ℓ a parametrization of the measured cross sections Σ has to take into account their variation with ϵ . We have divided therefore all data with respect to different regions of ϵ and have applied fits to these different sets of cross sections. As a result one gets information of the contributions of σ_ℓ to the total cross section.

First fits of the type presented in this report have been shown earlier²⁴⁾ *. Now some more inelastic e-p scattering data in the resonance region have become available^{9-12, 16-18)}. Also, more data^{9, 10, 12, 18, 20)} are

* This reference will henceforth be referred to as (I).

now available between $q^2 = 0$ and $q^2 = 0.5 \text{ GeV}^2$, a region in which, earlier, there was generally a lack of data as only a small portion of data from references 2,6,7) belonged to this range. Furthermore, photoproduction data²⁵⁾ can now be used for $q^2 = 0$ in place of the SLAC photon-proton cross sections (used earlier in I) obtained from inelastic electron-proton scattering cross sections by extrapolation to the limit of zero four-momentum transfer⁸⁾. We have, therefore, done the fitting of the entire electroproduction data available in the resonance region and present the results obtained therefrom in this report. It must be noted however that data for small values of ϵ are still very scarce and that there is a need for more measurements.

2. THE PROCEDURE OF FITTING

The procedure of fitting is the same as used in (I), which is given below:

The cross section Σ from early results shows the following behaviour²⁻⁶⁾:

$$\Sigma = G_D^2(q^2) \cdot A(W) \cdot |\vec{q}|^{b(W)}, \quad (2)$$

where

$G_D^2(q^2)$ = dipole form factor of the nucleon,
 \vec{q} = three momentum transfer to the hadronic system, and

$A(W)$ and $b(W)$ are parameters dependent only on W .

For the validity of eq. (2), $|\vec{q}|$ must be small as for large values of $|\vec{q}|$, Σ has values smaller than those obtained from the above equation. If the outgoing hadronic system gets a definite total angular momentum, eq. (2) represents the threshold behaviour of Σ for $|\vec{q}| \rightarrow 0$. A plot of $\log(\Sigma/G_D^2)$ vs. $\log|\vec{q}|$ gives a straight line; quadratic or higher powers of $\log|\vec{q}|$ may be introduced to account for deviations from this behaviour at high momentum transfers. One additional term is found to be sufficient (I) and, therefore, the following equation is used to fit the data:

$$\log(\Sigma/G_D^2) = a(W) + b(W) \cdot \log(|\vec{q}|/|\vec{q}_0|) + c(W) \cdot \left| \log(|\vec{q}|/|\vec{q}_0|) \right|^{d(W)}, \quad (3)$$

where $|\vec{q}_0|$ is the value of $|\vec{q}|$ at $q^2 = 0$ for the same W .

The fits done in this work enable us to study the dependence of Σ on ϵ . The data with $\epsilon \geq 0.9$ are taken from references 6 to 12 and 17 to 19 while those with $\epsilon \leq 0.6$ from references 3 to 5, 8, 9, 11, 19, and 20. For the fit in the range $0.6 < \epsilon < 0.9$, the relevant data from the above mentioned references as well as from ref. (2) are used. In W the fits were restricted to the range $1.11 \leq W \leq 1.99$ GeV. The Daresbury photoproduction cross sections²⁵⁾, having W in the relevant range, have been included in all ϵ ranges as σ_ℓ is zero at $q^2 = 0$. We have used two bin sizes for W , one, $\Delta W = 0.015$ GeV, across the bumps of the resonances up to $W = 1.755$ GeV, and the other, $\Delta W = 0.020$ GeV, for $1.770 \leq W \leq 1.990$.

The errors as given in the various references have been increased to include possible systematic errors, since most of the papers give only statistical errors. Furthermore, the final bin size in W introduces an error which is not negligible around the resonances where the cross section changes rapidly with W . Finally, there may be differences in the absolute normalization of the measurements from different experimental arrangements. Therefore, to the errors of the cross sections in the range $\epsilon \geq 0.9$ with $W > 1.755$ GeV a 5 % error has been added quadratically, to all others a 10 % error. Using the above errors the cross sections resulting from the fits were practically not different from those, where only the original errors were used.

3. RESULTS

Most of the additional data (9-12, 17, 18, 20) available after the completion of (I) belong to the range $\epsilon \geq 0.9$, and therefore the results of this work for the range $\epsilon \leq 0.6$ are not very different from the corresponding results given in (I).

3.1 The parameter d

For the range $\epsilon \leq 0.6$ the amount of data is not sufficient to determine the four parameters in eq. (3) for each W bin. Further, b and c cannot be separated if d is around 1. This happens for some W bins in the case $\epsilon \leq 0.6$, when cross sections for high values of q^2 are missing or when the errors are too large. Therefore in a first step d was left free for $\epsilon \geq 0.9$ only. From these fits an average value of d = 3.0 for all bins of W and all ranges of ϵ was taken for the final fits. The choice of d turned out not to be critical for the cross sections resulting from the fits.

3.2 Coefficients a, b and c

The values of the coefficients a, b and c as well as their errors and correlations as determined from the least squares fits of eq. (3) are given in Tables I, II, and III for the three ranges of ϵ . Also given are the weighted average values of ϵ for each W bin, the χ^2 per degree of freedom and the number of degrees of freedom.

3.3 Behaviour of the fits as a function of $|\vec{q}|^2$

In figs. 1 to 6 the measured values of Σ/G_D^2 are shown as a function of $|\vec{q}|^2$ separately for the two extreme ranges of ϵ for six values of W, i.e., for W = 1.230, 1.380, 1.530, 1.605, 1.695, and 1.890 GeV. Also shown for each W value is the scale of q^2 . As mentioned in sec. 1, there are now available some experimental data for the range $\epsilon \geq 0.9$, between $q^2 = 0$ and $q^2 = 0.5$ GeV², as can be seen from figs. 1 to 6.

Besides the experimental cross sections we have also shown the fit for the two ranges of ϵ for each value of W in figs. 1-6. The fits at W = 1.230, 1.530 and 1.605 for both ranges of ϵ almost pass through the photoproduction experimental points (Fig. 1, 3 and 4). The fits at W = 1.380 GeV pass below, while those at W = 1.695 and 1.890 GeV pass

above the corresponding experimental points but generally touching or passing through the error bars. It should, however, be pointed out that for $\epsilon \leq 0.6$ there are no experimental data below $q^2 = 0.5$ to guide the fit while for $\epsilon \geq 0.9$ the expression used to obtain the fit is somewhat inadequate to account for a σ_ℓ contribution restricted to small values of q^2 .

The corresponding results for $0.6 < \epsilon < 0.9$ are shown in Figs. 7 to 9. The behaviour of the fits for the three ranges of ϵ is practically the same. Also, the inclusion or non-inclusion of the photoproduction data does not affect the behaviour of the fits in any appreciable manner.

3.4 Dependence of Σ on W and q^2

In Figs. 10-18 we show the cross section Σ for fixed values of q^2 , i.e., $q^2 = 0.1, 0.3, 0.5, 1.0, 2.0, 3.0, 4.0, 5.0$ and 6.0 GeV^2 across the range $1.110 \leq W \leq 1.990 \text{ GeV}$ as computed from our fits. Whereas for $q^2 < 0.5 \text{ GeV}^2$ we only show results for $\epsilon \geq 0.9$, in the figures for $q^2 \geq 0.5 \text{ GeV}^2$ the values of Σ are compared for the two extreme ranges of ϵ . The errors are calculated with the complete error matrix of the coefficients a , b and c according to the following expression:

$$(\Delta\Sigma)^2 = (\Delta a)^2 \cdot \left(\frac{\delta\Sigma}{\delta a}\right)^2 + (\Delta b)^2 \cdot \left(\frac{\delta\Sigma}{\delta b}\right)^2 + (\Delta c)^2 \cdot \left(\frac{\delta\Sigma}{\delta c}\right)^2 - 2 \cdot C_{ab} \cdot \\ \cdot \frac{\delta\Sigma}{\delta a} \cdot \frac{\delta\Sigma}{\delta b} + 2 \cdot C_{ac} \cdot \frac{\delta\Sigma}{\delta a} \cdot \frac{\delta\Sigma}{\delta c} - 2 \cdot C_{bc} \frac{\delta\Sigma}{\delta b} \cdot \frac{\delta\Sigma}{\delta c},$$

where Δa , Δb , Δc , C_{ab} , C_{ac} and C_{bc} are the coefficients of the error matrix as given in tables I, II and III.

We think that the fits are not sufficient to calculate cross sections beyond $q^2 = 4.0 \text{ GeV}^2$ for $\epsilon \leq 0.6$ and beyond $q^2 = 6.0 \text{ GeV}^2$ for $\epsilon \geq 0.9$.

The results show the faster decrease of the first resonance with increasing q^2 compared to the nonresonant background around this resonance and compared to the two other bumps of resonances around 1.5 and 1.67 GeV.

The second bump on the contrary increases slightly in relation to the background at the same energy with increasing q^2 whereas the third one stays practically constant. The position of the third bump moves to higher energies with increasing q^2 as has been noticed earlier¹⁸⁾.

3.5 The Ratio $R = \sigma_\ell / \sigma_t$

The difference between $\Sigma_{\epsilon \leq 0.9}$ and $\Sigma_{\epsilon \geq 0.6}$ is a direct measure of σ_ℓ and is very close to $\frac{1}{2} \cdot \sigma_\ell$ (eq.1), since the average values of ϵ in the two cases are close to 1.0 and 0.5 (Tables I and III). Therefore, the comparison of Σ for the two ranges of ϵ in the Figs. 12-15 allows one to determine σ_ℓ . There are systematic differences outside the error bars between the two sets of cross sections. For $q^2 = 0.5 \text{ GeV}^2$ with $W < 1.5 \text{ GeV}$ the difference $\Sigma_{\geq 0.9} - \Sigma_{\leq 0.6}$ has unphysical values. This is most likely due to the uncertainty of the fit for $\epsilon < 0.6$ for small values of q^2 because of lack of data. It also may show the problem of estimating systematic errors for the different set of measurements, measured with different experimental arrangements. The same conclusion could be made for $W > 1.6 \text{ GeV}$ and small q^2 values, where $\Sigma_{\epsilon \leq 0.6}$ is systematically smaller than $\Sigma_{\epsilon \geq 0.9}$. However there the difference stays on to larger values of q^2 up to 3 GeV^2 , indicating a possible longitudinal contribution of about 10-20 %. This is consistent with a direct determination of R at $q^2 = 1 \text{ GeV}^2$ ¹¹⁾.

For $W < 1.6 \text{ GeV}$ and $0.5 < q^2 < 2 \text{ GeV}^2$ no difference outside the error bars between the two sets of measurement is visible except for a small effect on the sides of the first resonance, which might be due to uncertainties in the energy W . A 10 % contribution of σ_ℓ at the first resonance around $q^2 = 0.5 \text{ GeV}^2$ is indicated by direct determinations^{20,21)}. Above $q^2 = 2 \text{ GeV}^2$ and for $1.3 < W < 1.5 \text{ GeV}$ a longitudinal contribution of about 20 % is not excluded.

LIST OF TABLES

Table I: Results of the fits for $\epsilon \geq 0.9$

Table II: Results of the fits for $0.9 > \epsilon > 0.6$

Table III: Results of the fits for $\epsilon \leq 0.6$

REFERENCES

1. H.L. Lynch, J.W. Allaby and D.M. Ritson, HEPL-494 B (1967)
2. A.A. Cone, K.W. Chen, J.R. Dunning, G. Hartwig, N.F. Ramsey, J.K. Walker, R. Wilson, Phys. Rev. 156 (1967) 149o
3. F.W. Brasse, J. Engler, E. Ganßauge, M. Schweitzer, Nuovo Cim. X, 55A, 679 (1968) and DESY 67/34 (1967)
4. W. Albrecht, F.W. Brasse, H. Dorner, W. Flauger, K.-H. Frank, J. Gayler, H. Hultschig, J. May, E. Ganßauge, Phys. Lett. 28B (1968) 225 and DESY 68/48 (1968)
5. W. Albrecht, F.W. Brasse, H. Dorner, W. Flauger, K.-H. Frank, J. Gayler, H. Hultschig, J. May, E. Ganßauge, Nucl. Phys. B13, (1969) 1 and DESY 69/7
6. W. Bartel, B. Dudelzak, H. Krehbiel, J. McElroy, U. Meyer-Berkhout, W. Schmidt, V. Walther, G. Weber, Phys. Lett. 28B (1968) 148; W. Bartel, thesis, DESY internal report F22-69/3 (1969)
7. E.D. Bloom, G. Buschhorn, R.L. Cottrell, D.H. Coward, H. De Staebler, J. Drees, C.L. Jordon, G. Miller, L. Mo, H. Piel, R.E. Taylor, M. Breidenbach, W.R. Ditzler, J.I. Friedman, G.C. Hartmann, H.W. Kendall, J.S. Pucher, SLAC-PUB-795 (1970);
M. Breidenbach, Thesis; R.E. Taylor, private communication
8. E.D. Bloom, R.L. Cottrell, D.H. Coward, H. DeStaebler, Jr., J. Drees, G. Miller, L.W. Mo, R.E. Taylor, J.I. Friedman, G.C. Hartmann, and H.W. Kendall, SLAC-PUB 653 (1969)
9. J. Moritz, K.H. Schmidt, D. Wegener, J. Bleckwenn and E. Engles, Jr. DESY 71/61; Nucl. Phys. B41 (1972) 336
10. J. Bleckwenn, thesis, DESY internal report F23-71/2 (1971)
11. J.-C. Alder, F.W. Brasse, E. Chazelas, W. Fehrenbach, W. Flauger, K.-H. Frank, E. Ganßauge, J. Gayler, W. Krechlok, V. Korbel, J. May, M. Merkwitz, P.D. Zimmerman, Nucl. Phys. B48 (1972), 487 and DESY 72/38 (1972)

12. M. K  bberling, J. Moritz, K.H. Schmidt, D. Wegener, D. Zeller, J. Bleckwenn, F.H. Heimlich, Karlsruhe KFK 1822; Nucl. Phys. B82 (1974) 201
13. E.D. Bloom, D.H. Coward, H. DeStaebler, J. Drees, G. Miller, L.W. Mo, R.E. Taylor, M. Breidenbach, J.I. Friedman, G.C. Hartmann, H.W. Kendall, Phys. Rev. Lett. 23 (1969) 930
14. M. Breidenbach, J.I. Friedman, H.W. Kendall, E.D. Bloom, D.H. Coward, H. DeStaebler, J. Drees, L.W. Mo, and R.E. Taylor, Phys. Rev. Lett. 23 (1969) 935
15. G. Miller, E.D. Bloom, G. Buschhorn, D.H. Coward, H. DeStaebler, J. Drees, C.L. Jordan, L.W. Mo, R.E. Taylor, J.I. Friedman, G.C. Hartmann, H.W. Kendall and R. Verdier, Phys. Rev. D5 (1972) 528
16. J.S. Poucher, M. Breidenbach, R. Ditzler, J.I. Friedman, H.W. Kendall, E.D. Bloom, R.L.A. Cottrell, D.H. Coward, H. DeStaebler, C.L. Jordan, H. Piel, and R.E. Taylor, SLAC PUB 1309 (1973); Phys. Rev. Lett. 32 (1974) 118
17. Arie Bodek, Ph. D. thesis, MIT (1972);
Edard M. Riordan, Ph. D. thesis, MIT (1973) and also private communication
18. S. Stein, W.B. Atwood, E.D. Bloom, R.L.A. Cottrell, H. DeStaebler, C.L. Jordan, H.G. Piel, C.Y. Prescott, R. Siemann, R.E. Taylor, SLAC-PUB-1528 (1975)
19. W. Bartel, B. Dudelzak, H. Krehbiel, J. McElroy, U. Meyer-Berkhout, W. Schmidt, V. Walter, G. Weber, Phys. Lett. 27B (1968) 660
20. W. Bartel, F.W. B  sser, W.R. Dix, R. Felst, D. Harms, H. Krehbiel, P.E. Kuhlmann, J. McElroy, J. Meyer, G. Weber, Phys. Lett. 35B (1971) 181 and W.R. Dix, Thesis, Universit  t Hamburg (1971)
21. K. B  tzner, U. Beck, K.H. Becks, C. Berger, J. Drees, G. Knop, M. Leenen, K. Moser, C. Nietzel, E. Schlosser, H.E. Stier, Phys. Lett. 39B(1972) 575
22. W. Albrecht, F.W. Brasse, H. Dorner, W. Flauger, K.-H. Frank, J. Gayler, H. Hultschig, V. Korbel, J. May, DESY 69/46 (1969)

23. E.M. Riordan, A. Bodek, M. Breidenbach, D.L. Dubin, J.E. Elias, J.I. Friedman, H.W. Kendall, J.S. Poucher, M.R. Sogard, D.H. Coward, SLAC-PUB 1417 (1974)
24. F.W. Brasse, W. Fehrenbach, W. Flauger, K.-H. Frank, J. Gayler, V. Korbel, J. May, P.D. Zimmerman and E. Ganßauge, DESY 71/2 (1971)
25. T.A. Armstrong, W.R. Hogg, G.M. Lewis, A.W. Robertson, G.R. Brookes, A.S. Clough, J.H. Freeland, W. Galbraith, A.F. King, W.R. Rawlinson, N.R.S. Tait, J.C. Thompson and D.W.L. Tolfree, Daresbury DNPL/P 88 (1971)

FIGURE CAPTIONS

- Fig. 1: Σ/G_D^2 as a function of $|\vec{q}|^2$ in a double logarithmic plot for $W = 1.230$ GeV and with $d = 3.0$. Shown is also the scale of q^2 . The curves are our fit with eq. (2) and with the parameters in tables I and III.
For $\epsilon \geq 0.9$, error bars are given for extreme positions while for $\epsilon \leq 0.6$, they are given for each data point.
- Fig. 2: same as 1, but for $W = 1.380$ GeV
- Fig. 3: same as 1, but for $W = 1.530$ GeV
- Fig. 4: same as 1, but for $W = 1.605$ GeV
- Fig. 5: same as 1, but for $W = 1.695$ GeV
- Fig. 6: same as 1, but for $W = 1.890$ GeV
- Fig. 7: same as 1, but for $0.6 < \epsilon < 0.9$, $W = 1.230$ and 1.380 GeV
- Fig. 8: Same as 7, but for $W = 1.530$ and 1.605 GeV
- Fig. 9: same as 7, but for $W = 1.695$ and 1.890 GeV
- Fig. 10: Σ calculated from our fits for $\epsilon \geq 0.9$ as a function of W for $q^2 = 0.1$ GeV 2
- Fig. 11: Same as 10, but for $q^2 = 0.3$ GeV 2
- Fig. 12: Σ calculated from our fits for $\epsilon \geq 0.9$ and $\epsilon \leq 0.6$, as a function of W for $q^2 = 0.5$ GeV 2
- Fig. 13: same as 12, but for $q^2 = 1.0$ GeV 2
- Fig. 14: same as 12, but for $q^2 = 2.0$ GeV 2
- Fig. 15: same as 12, but for $q^2 = 3.0$ GeV 2
- Fig. 16: same as 12, but for $q^2 = 4.0$ GeV 2
- Fig. 17: Σ calculated from our fits for $\epsilon \geq 0.9$ as a function of W for $q^2 = 5.0$ GeV 2
- Fig. 18: same as 17, but for $q^2 = 6.0$ GeV 2

	ϵ	a	b	$(\Delta b) \cdot 10^4$	c	$(\Delta c) \cdot 10^4$	$c_{ab} \cdot 10^4$	$c_{ac} \cdot 10^4$	$c_{bc} \cdot 10^4$	χ^2	NF
W	1.110	5.045	251.7	3.798	217.3	2.9	223.2	23.8	2.3	2.3	57
	1.125	0.962	1.63.5	1.052	166.4	2.5	163.5	16.9	20.3	2.3	70
	1.140	C.963	5.350	1.92.8	1.213	0.024	110.1	13.5	15.4	2.2	71
	1.155	C.956	5.621	83.1	1.334	0.023	36.0	2.0	1.1	2.1	71
	1.170	C.959	5.943	42.5	1.397	0.013	36.0	1.4	4.4	1.3	69
	1.185	C.957	5.952	45.3	1.727	0.023	7.9	7.8	7.8	1.1	83
	1.200	C.960	6.139	31.8	1.750	0.060	1.5	4.9	6.7	1.1	76
	1.215	C.958	6.178	26.7	1.878	0.080	1.4	31.0	6.7	0.9	83
	1.230	C.961	6.125	27.3	1.887	0.085	1.8	32.6	5.2	0.7	85
	1.245	C.960	2.959	23.1	1.927	0.056	1.5	28.0	4.4	0.8	89
	1.260	C.955	5.769	25.5	2.041	0.065	2.3	32.8	5.5	1.3	83
	1.275	C.962	2.622	21.5	2.384	0.050	2.4	23.6	5.3	1.1	80
	1.290	C.960	5.431	22.0	2.148	0.043	2.4	23.5	5.1	1.5	76
	1.305	C.956	2.238	1.946	2.205	0.034	2.8	26.4	5.3	9.5	85
	1.320	C.957	5.175	17.1	2.344	0.054	3.2	25.0	5.3	10.2	85
	1.335	C.958	2.131	15.7	2.324	0.013	3.2	22.8	4.9	9.6	90
	1.350	C.957	2.053	21.9	2.535	0.046	5.6	33.5	7.9	15.8	76
	1.365	C.954	2.075	16.0	2.464	0.042	4.4	24.8	6.4	12.1	86
	1.380	C.957	5.345	15.7	2.554	0.045	4.2	24.5	6.6	11.5	77
	1.395	C.957	5.078	13.9	2.610	0.011	7.4	3.4	9.7	1.3	71
	1.410	C.960	5.145	16.0	2.639	0.032	5.9	26.1	6.4	16.8	90
	1.425	C.957	5.156	14.8	2.673	0.046	5.4	27.3	6.1	13.4	75
	1.440	C.957	5.224	16.6	2.771	0.084	6.2	27.8	6.3	15.4	67
	1.455	C.959	2.258	15.7	2.899	0.115	7.1	26.2	7.0	16.1	85
	1.470	C.958	2.571	15.0	2.982	0.103	7.6	26.1	7.1	17.2	67
	1.485	C.956	5.457	15.3	3.157	0.152	5.4	25.0	6.2	14.1	69
	1.500	C.956	5.423	15.4	3.163	0.164	8.5	27.3	6.3	13.3	79
	1.515	C.927	5.212	16.7	3.315	0.181	7.5	26.1	7.1	17.1	67
	1.530	C.947	5.224	16.6	3.315	0.181	7.5	26.1	7.0	17.6	75
	1.545	C.957	5.354	15.4	3.375	0.213	7.4	23.7	6.8	24.5	67
	1.560	C.955	2.364	16.7	3.450	0.113	12.9	31.3	6.6	24.5	63
	1.575	C.955	2.341	14.0	3.477	0.245	8.4	26.0	7.3	18.4	65
	1.590	C.954	5.320	15.6	3.471	0.264	16.9	31.1	6.1	29.1	65
	1.605	C.927	5.212	16.7	3.554	0.239	17.0	32.1	6.1	23.5	57
	1.620	C.947	5.212	16.7	3.554	0.239	17.0	32.1	6.1	23.5	55
	1.635	C.956	2.462	16.6	3.633	0.213	14.4	32.2	6.8	27.1	61
	1.650	C.957	5.364	16.7	3.695	0.16	19.9	32.0	6.3	33.1	59
	1.665	C.955	2.341	15.0	3.804	0.212	19.4	34.4	6.2	26.6	52
	1.680	C.955	5.320	15.6	3.900	0.383	22.1	31.6	6.2	35.1	51
	1.695	C.954	2.258	17.0	3.933	0.393	16.9	29.5	6.2	27.1	50
	1.710	C.954	2.276	16.9	3.933	0.393	16.9	29.5	6.2	27.1	50
	1.725	C.955	5.295	15.2	4.704	0.302	14.4	32.2	6.8	34.7	43
	1.740	C.955	5.223	14.3	4.757	0.302	14.4	32.2	6.8	34.7	43
	1.755	C.954	5.478	14.3	4.047	0.393	16.9	29.5	6.2	27.1	51
	1.770	C.955	5.495	14.2	4.210	0.38	21.3	29.3	6.2	34.7	43
	1.785	C.954	5.350	13.7	4.519	0.517	23.0	29.1	6.9	34.7	43
	1.800	C.953	5.330	15.7	4.695	0.519	20.3	32.0	6.3	33.1	59
	1.815	C.953	2.313	13.3	4.704	0.519	19.4	33.1	6.7	33.1	59
	1.830	C.952	5.295	14.2	4.757	0.302	14.4	32.2	6.8	34.7	43
	1.845	C.954	5.223	13.6	4.840	0.393	16.9	29.5	6.2	27.1	50
	1.860	C.955	5.443	14.2	4.210	0.38	21.3	29.3	6.2	34.7	43
	1.875	C.955	5.350	13.7	4.519	0.517	23.0	29.1	6.9	34.7	43
	1.890	C.953	5.330	15.7	4.695	0.519	20.3	32.0	6.3	33.1	59
	1.905	C.953	2.313	13.3	4.704	0.519	19.4	33.1	6.7	33.1	59
	1.920	C.952	5.295	14.2	4.757	0.302	14.4	32.2	6.8	34.7	43
	1.935	C.954	5.223	13.6	4.840	0.393	16.9	29.5	6.2	27.1	50
	1.950	C.955	5.443	14.2	4.210	0.38	21.3	29.3	6.2	34.7	43
	1.965	C.955	5.350	13.7	4.519	0.517	23.0	29.1	6.9	34.7	43
	1.980	C.953	5.330	15.7	4.695	0.519	20.3	32.0	6.3	33.1	59
	1.995	C.953	2.313	13.3	4.704	0.519	19.4	33.1	6.7	33.1	59

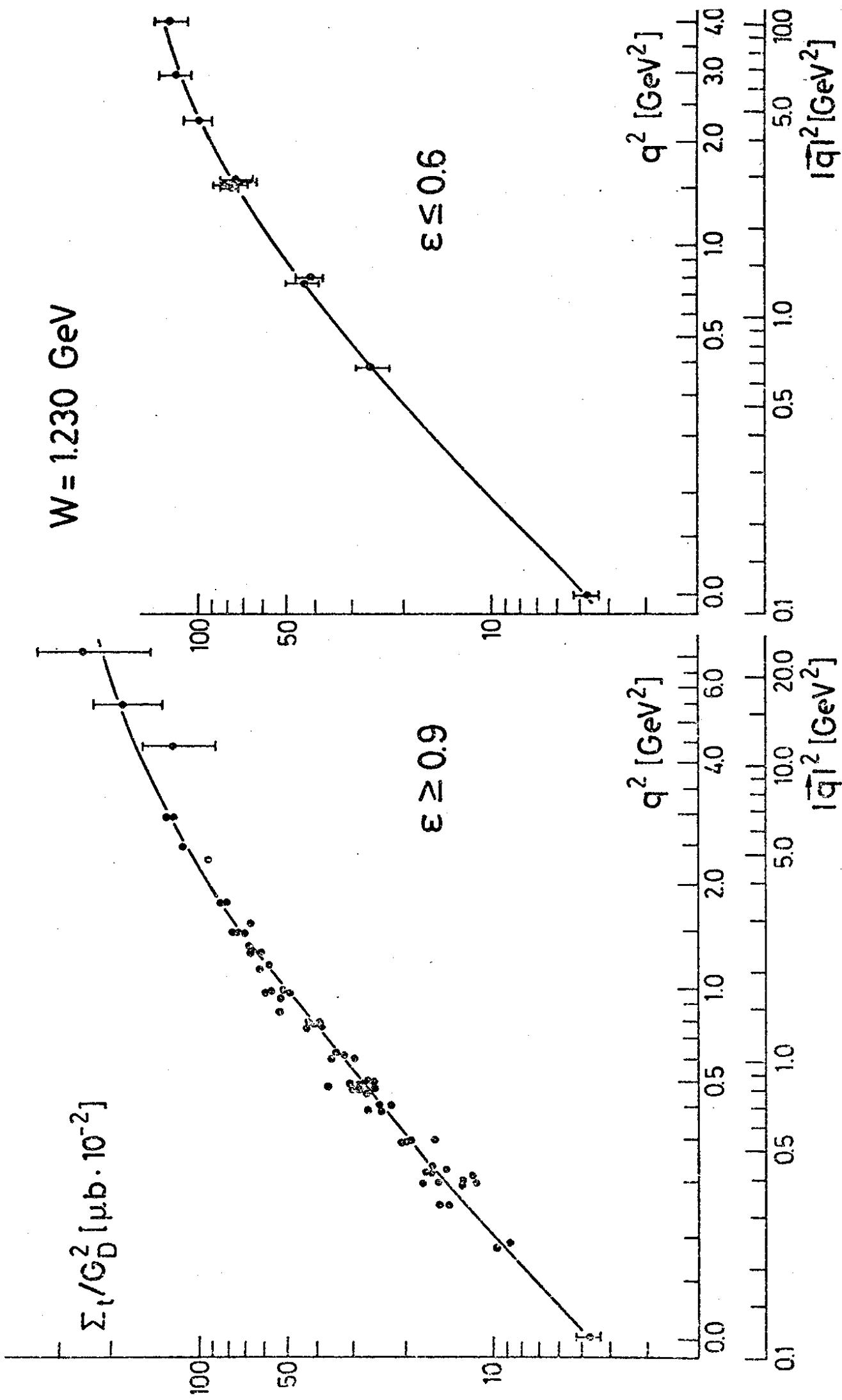
Table I: Results of the fits for $\epsilon = 0.9$

	χ^2	NF	$\epsilon_c \cdot 10^4$	$C_{bc} \cdot 10^4$	$C_{ac} \cdot 10^4$	$(\Delta c)^2 \cdot 10^4$	$C_{ab} \cdot 10^4$	c	$(\Delta b)^2 \cdot 10^4$	b	$(\Delta a)^2 \cdot 10^4$	a	ϵ
W	1.110	0.835	16677.8	4.849	9928.3	-0.285	73.8	12807.8	7415.9	42.8	7415.9	42.8	2.3
	1.125	0.820	1.082	9586.8	4.120	5790.4	-0.212	610.8	2895.1	24.6	272.3	486.4	1.7
	1.140	0.837	4.119	1435.7	2.244	2468.2	-0.100	238.1	2135.3	20.9	209.8	238.1	2.5
	1.155	0.859	3.898	2482.6	2.824	1864.3	-0.155	189.3	79.5	3.5	13.7	13.7	2.1
	1.170	0.854	5.990	90.1	1.257	90.9	-0.006	13.3	79.5	7.6	101.6	101.6	1.9
	1.185	0.844	6.033	1077.1	1.548	910.2	-0.025	11.5	982.8	101.6	11.5	11.5	2.5
	1.200	0.854	6.160	85.7	1.730	93.4	-0.066	9.9	13.7	9.9	13.7	9.9	1.6
	1.215	0.857	6.219	84.6	1.805	95.1	-0.086	4.0	82.5	9.6	15.1	15.1	0.5
	1.230	0.864	6.117	81.6	1.866	97.5	-0.071	2.5	84.3	10.4	13.7	13.7	1.1
	1.245	0.859	5.959	77.9	1.914	82.8	-0.054	1.5	75.8	7.7	9.5	9.5	1.1
	1.260	0.851	5.451	584.1	2.272	723.0	-0.084	15.8	645.1	89.9	103.2	103.2	3.1
	1.275	0.848	5.675	79.8	1.840	121.0	-0.012	4.9	92.4	13.6	20.9	20.9	1.6
	1.290	0.858	5.417	85.8	2.102	127.6	-0.037	5.4	96.8	14.2	22.4	22.4	0.5
	1.305	0.845	5.238	82.2	2.221	134.1	-0.046	7.0	98.4	15.9	25.7	25.7	1.1
	1.320	0.781	5.084	71.9	2.368	105.8	-0.058	2.7	81.7	12.4	15.1	15.1	2.4
	1.335	0.783	4.913	57.5	2.622	76.1	-0.086	1.0	62.3	5.6	7.5	7.5	3.1
	1.350	0.798	4.458	407.6	3.146	69.5	-0.131	25.1	526.1	93.7	127.3	127.3	3.2
	1.365	0.856	5.012	69.0	2.537	148.3	-0.031	10.3	71.7	17.5	34.0	34.0	1.1
	1.380	0.859	5.128	86.1	2.473	172.3	-0.031	9.8	111.5	20.0	35.3	35.3	1.4
	1.395	0.851	5.140	75.9	2.533	158.3	-0.031	13.6	101.1	20.3	38.5	38.5	1.3
	1.410	0.828	5.261	38.2	2.610	67.4	-0.093	2.2	46.2	6.0	10.2	10.2	2.2
	1.425	0.853	5.370	44.7	2.461	121.3	-0.002	13.5	66.1	15.6	34.9	34.9	1.1
	1.440	0.851	5.416	42.3	2.570	114.9	-0.052	11.5	62.3	14.1	31.5	31.5	1.2
	1.455	0.788	5.466	24.5	2.700	77.7	-0.083	8.1	39.5	9.0	21.8	21.8	0.7
	1.470	0.824	5.508	30.2	2.799	97.0	-0.093	10.4	47.7	11.5	27.9	27.9	0.9
	1.485	0.847	5.578	29.6	2.929	102.4	-0.104	13.0	49.2	12.9	31.8	31.8	2.7
	1.500	0.839	5.629	29.4	3.016	111.7	-0.061	19.7	50.7	14.7	39.8	39.8	1.1
	1.515	0.847	5.623	20.4	3.128	71.3	-0.163	7.1	33.1	7.6	19.6	19.6	2.4
	1.530	0.841	5.555	23.2	3.215	103.4	-0.173	22.2	43.3	14.3	41.2	41.2	3.3
	1.545	0.851	5.503	24.4	3.271	118.3	-0.192	26.0	46.5	15.5	48.1	48.1	2.6
	1.560	0.852	5.471	19.1	3.288	90.7	-0.191	13.3	35.5	11.3	35.5	35.5	0.7
	1.575	0.819	5.419	21.8	3.328	111.7	-0.202	29.0	43.7	15.9	51.2	51.2	3.2
	1.590	0.846	5.390	21.5	3.393	109.6	-0.163	7.1	33.1	7.6	19.6	19.6	3.3
	1.605	0.845	5.423	17.9	3.376	110.9	-0.158	34.2	37.7	14.7	53.2	53.2	3.4
	1.620	0.840	5.396	18.6	3.567	106.6	-0.236	27.1	39.4	14.2	47.2	47.2	3.3
	1.635	0.765	5.451	12.7	3.600	50.5	-0.239	2.6	21.3	3.5	9.7	9.7	3.6
	1.650	0.821	5.403	18.4	3.495	136.6	-0.408	36.3	43.7	15.9	51.2	51.2	2.5
	1.665	0.927	5.446	16.3	4.346	127.4	-0.402	2.9	42.3	14.1	43.5	43.5	0.5
	1.680	0.690	5.448	13.1	4.124	63.5	-0.355	5.9	24.2	5.2	16.2	16.2	3.5
	1.695	0.813	5.421	12.7	4.270	78.7	-0.400	10.9	26.5	7.2	24.8	24.8	3.4
	1.710	0.819	5.076	9.8	5.112	82.0	-0.768	41.7	24.9	13.3	51.4	51.4	0.8
	1.725	0.810	5.308	20.5	4.667	161.0	-0.633	52.2	50.5	22.5	81.1	81.1	3.2
	1.740	0.814	5.193	16.0	4.856	122.9	-0.642	46.5	47.9	21.0	75.7	75.7	1.1
	1.755	0.818	5.099	9.0	5.069	74.0	-0.773	27.9	22.6	10.8	40.7	40.7	2.9
	1.770	0.819	5.076	9.8	5.112	82.0	-0.768	41.7	24.9	13.3	51.4	51.4	0.8
	1.785	0.823	5.076	14.3	5.819	146.4	-1.223	85.6	40.9	25.2	121.5	121.5	2.9
	1.800	0.775	5.054	9.3	5.236	72.7	-0.879	33.8	23.2	12.2	44.2	44.2	3.9
	1.815	0.821	5.064	8.4	5.331	74.5	-0.941	38.3	21.8	11.3	47.3	47.3	3.1
	1.830	0.818	5.028	13.4	5.609	109.5	-0.637	31.1	39.0	14.4	53.9	53.9	3.3
	1.845	0.823	5.380	9.7	5.628	86.7	-0.773	27.9	22.6	10.8	40.7	40.7	2.9
	1.860	0.823	5.076	9.8	5.112	82.0	-0.768	41.7	24.9	13.3	51.4	51.4	0.8
	1.875	0.810	5.173	9.4	5.461	60.4	-0.812	12.5	20.0	6.8	44.2	44.2	2.3
	1.890	0.775	5.011	13.5	6.391	160.0	-1.581	103.7	40.2	25.0	114.4	114.4	1.5
	1.905	0.833	5.136	19.2	5.995	156.6	-1.219	60.2	49.0	25.0	89.8	89.8	1.7
	1.920	0.811	5.055	14.1	6.455	184.0	-1.605	132.1	43.7	29.3	141.6	141.6	2.1
	1.935	0.781	5.035	14.1	6.565	154.2	-1.055	19.7	40.9	12.4	50.3	50.3	1.5
	1.950	0.648	5.035	16.8	7.063	116.7	-4.060	3950.9	2.2	421.2	6320.8	6320.8	1.5

Table III: Results of the fits for $0.9 > \epsilon > 0.6$

ϵ	a	$(\Delta a)^2 \cdot 10^4$	b	$(\Delta b)^2 \cdot 10^4$	c	$(\Delta c)^2 \cdot 10^4$	$C_{ab} \cdot 10^4$	$C_{ac} \cdot 10^4$	$C_{bc} \cdot 10^4$	χ^2	NF		
W	0.428	21932.4	10961.8	-3.350	49.5	15445.2	1014.9	728.0	2.0	8			
	3.795	30057.5	1.935	-3.005	153.2	24640.8	2113.8	1707.9	1.5	10			
	6.003	15772.4	5.774	1014.7	83.0	12622.9	1112.6	907.7	2.2	10			
	6.339	2102.8	0.653	1420.6	3.054	9.7	1714.5	134.7	114.1	2.6	11		
	6.371	92.2	1.017	113.3	3.053	3.9	31.7	8.6	19.3	1.0	14		
	6.334	4390.1	0.739	4387.4	3.023	80.5	4372.0	577.7	586.9	1.4	15		
	6.429	1.166	37.5	1.695	3.054	2.7	31.6	9.0	14.0	1.2	16		
	6.498	6.239	50.6	1.763	3.059	2.2	31.6	9.0	14.0	1.2	16		
	6.215	5.450	77.9	1.329	12.7	4.7	34.4	10.1	20.9	0.2	11		
	6.430	6.149	83.0	1.244	12.6	4.6	34.4	10.1	20.9	0.2	11		
	6.444	5.988	40.5	1.042	15.9	3.060	11.0	18.1	36.3	0.7	15		
	6.453	5.155	45.3	2.793	537.9	3.058	486.6	65.8	76.1	0.6	19		
	6.475	2.759	27.4	1.242	24.6	2.2	32.2	136.7	33.3	77.5	0.8	15	
	6.260	2.759	27.4	1.242	24.6	2.2	32.2	136.7	33.3	77.5	0.8	15	
	6.185	6.429	4390.1	0.739	4387.4	3.023	80.5	4372.0	577.7	586.9	1.4	15	
	6.230	9.498	6.239	50.6	1.763	3.059	2.2	31.6	9.0	14.0	1.2	16	
	6.213	10961.8	19357.3	-3.350	49.5	15445.2	1014.9	728.0	2.0	8			
	6.213	19357.3	19357.3	-3.005	153.2	24640.8	2113.8	1707.9	1.5	10			
	6.213	10961.8	19357.3	-3.036	83.0	12622.9	1112.6	907.7	2.2	10			
	6.213	19357.3	19357.3	-3.060	11.0	1714.5	134.7	114.1	2.6	11			
	6.213	10961.8	19357.3	-3.017	12.6	31.7	8.6	19.3	1.0	14			
	6.213	19357.3	19357.3	-3.054	3.9	31.7	8.6	19.3	1.0	14			
	6.213	10961.8	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	19357.3	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	10961.8	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	19357.3	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	10961.8	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	19357.3	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	10961.8	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	19357.3	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	10961.8	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	19357.3	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	10961.8	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	19357.3	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	10961.8	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	19357.3	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	10961.8	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	19357.3	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	10961.8	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	19357.3	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	10961.8	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	19357.3	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	10961.8	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	19357.3	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	10961.8	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	19357.3	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	10961.8	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	19357.3	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	10961.8	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	19357.3	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	10961.8	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	19357.3	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	10961.8	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	19357.3	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	10961.8	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	19357.3	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	10961.8	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	19357.3	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	10961.8	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	19357.3	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	10961.8	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	19357.3	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	10961.8	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	19357.3	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	10961.8	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	19357.3	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	10961.8	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	19357.3	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	10961.8	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	19357.3	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	10961.8	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	19357.3	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	10961.8	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	19357.3	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	10961.8	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	19357.3	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	10961.8	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	19357.3	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	10961.8	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	19357.3	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	10961.8	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	19357.3	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	10961.8	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	19357.3	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	10961.8	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	19357.3	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	10961.8	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	19357.3	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	10961.8	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	19357.3	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	10961.8	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	19357.3	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	10961.8	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	19357.3	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	10961.8	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	19357.3	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	10961.8	19357.3	-3.053	3.9	31.7	8.6	19.3	1.0	14			
	6.213	19											

Fig. 1



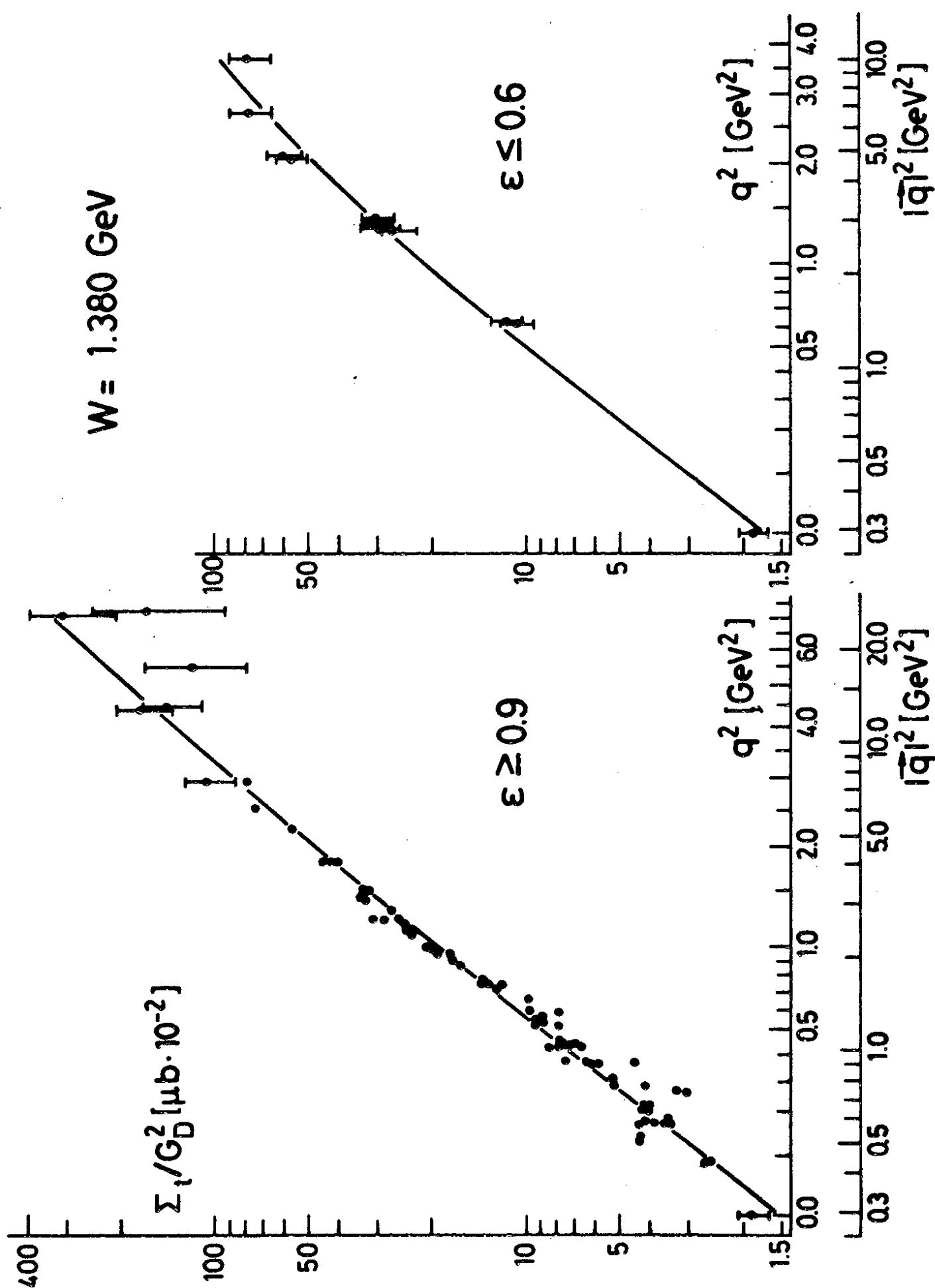


Fig. 2

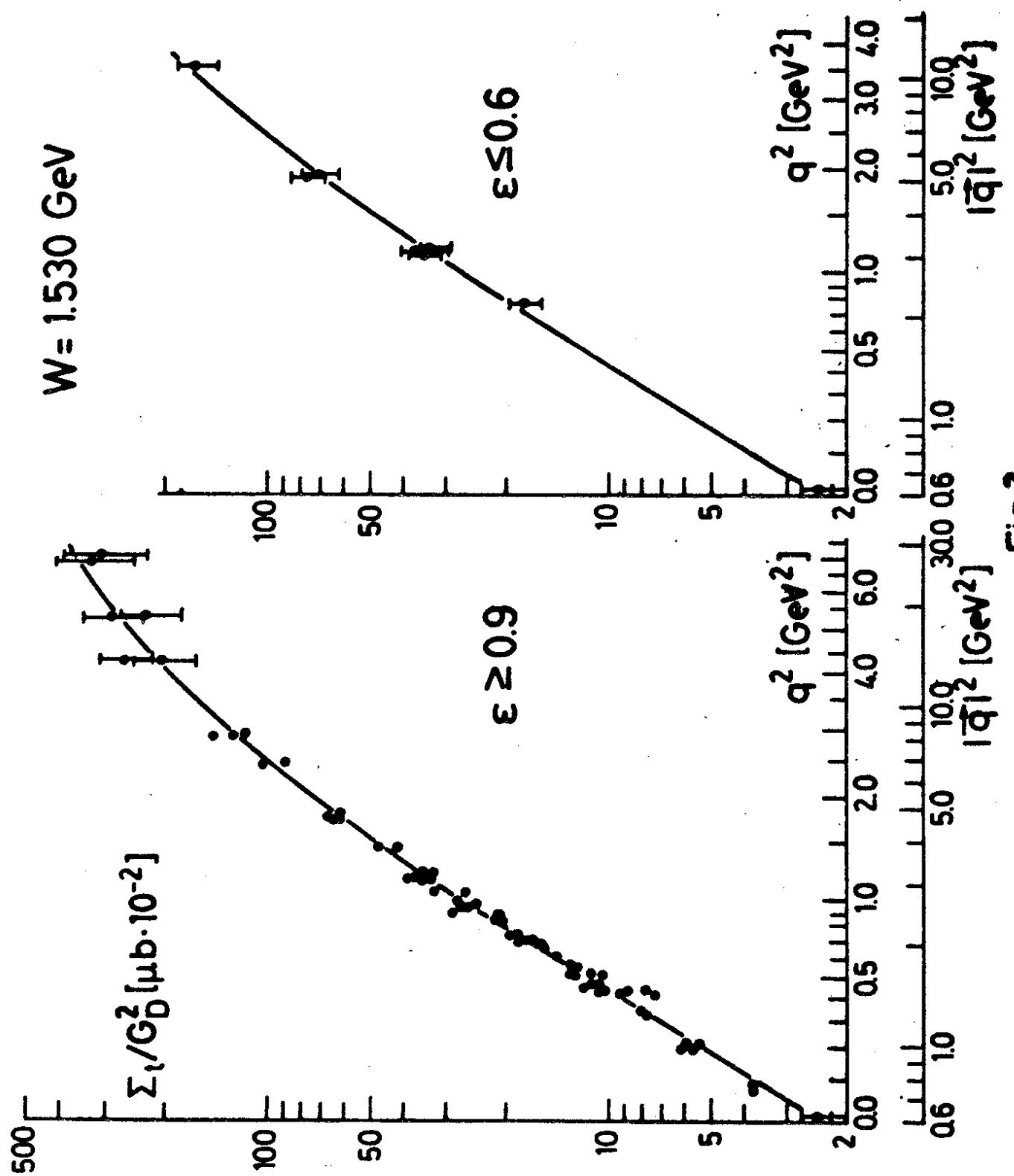


Fig. 3

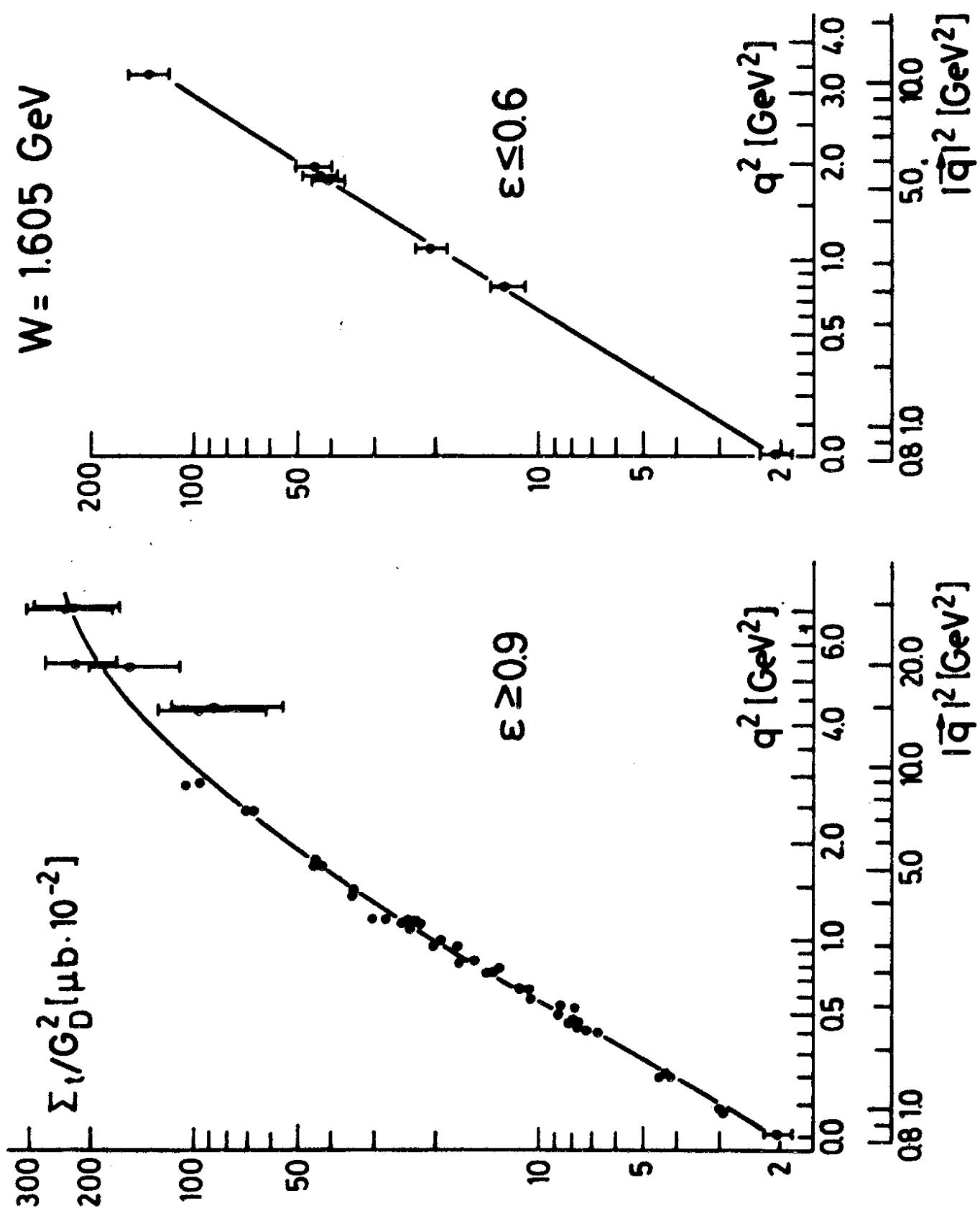


Fig. 4

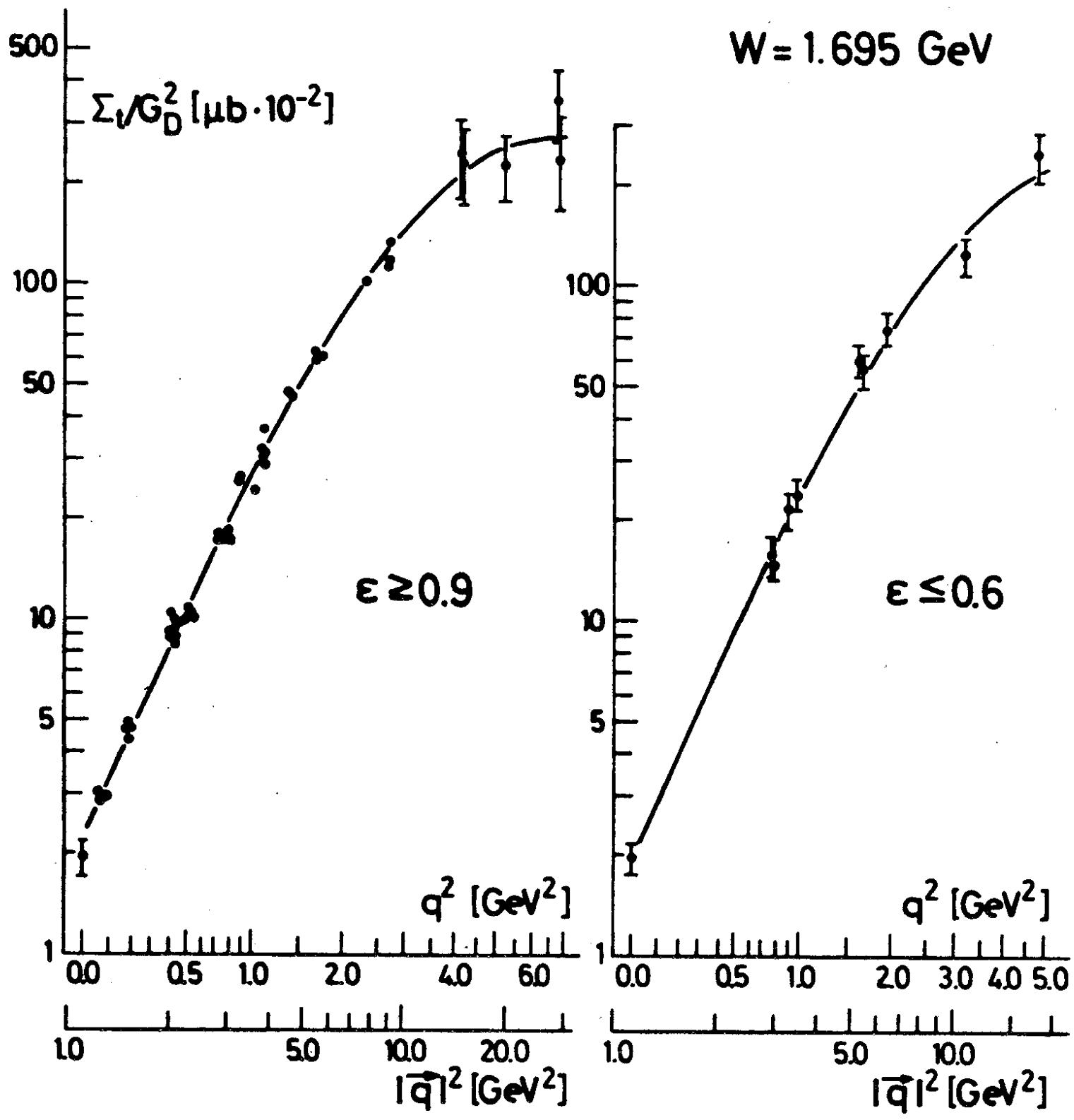


Fig. 5

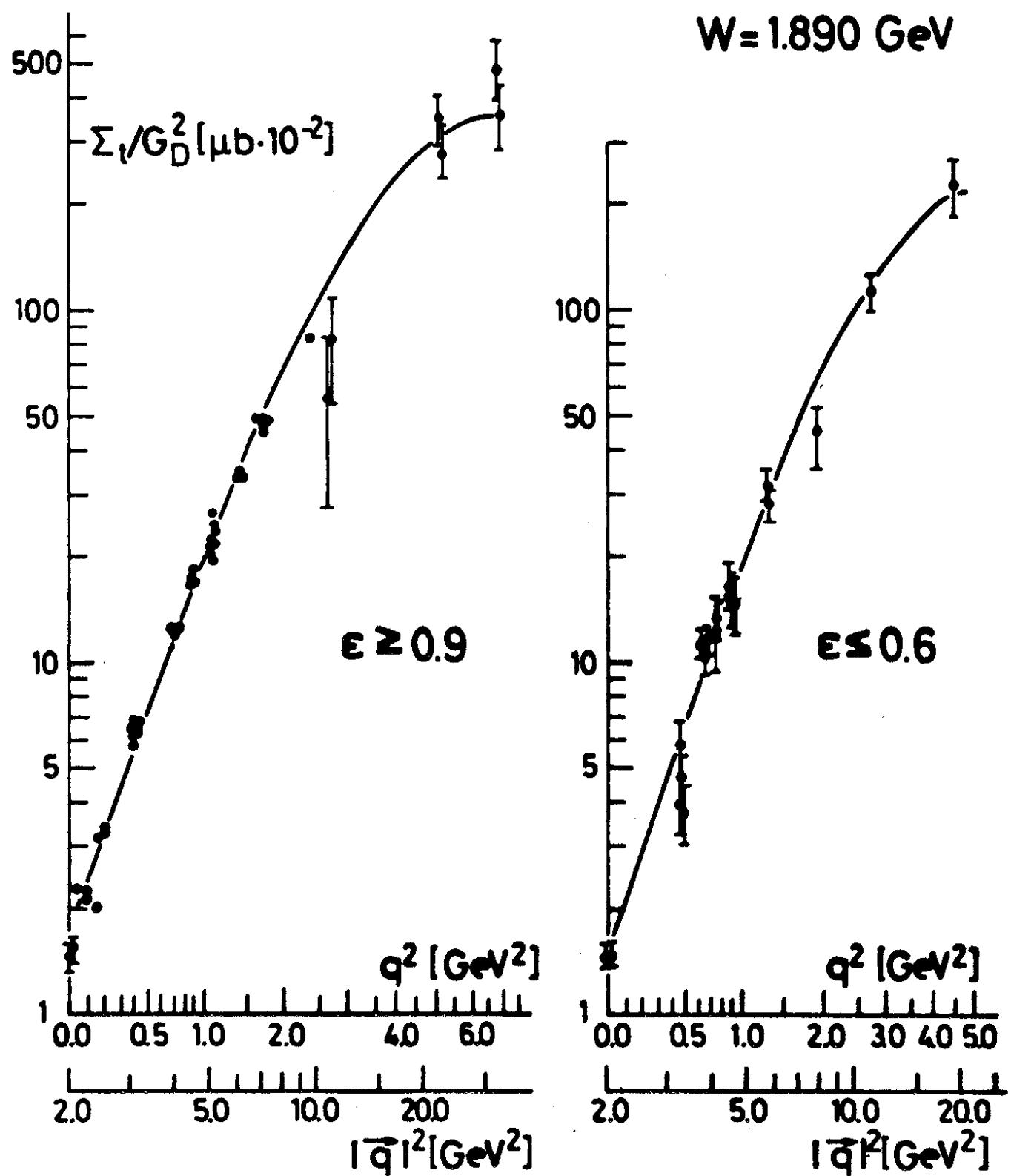


Fig. 6

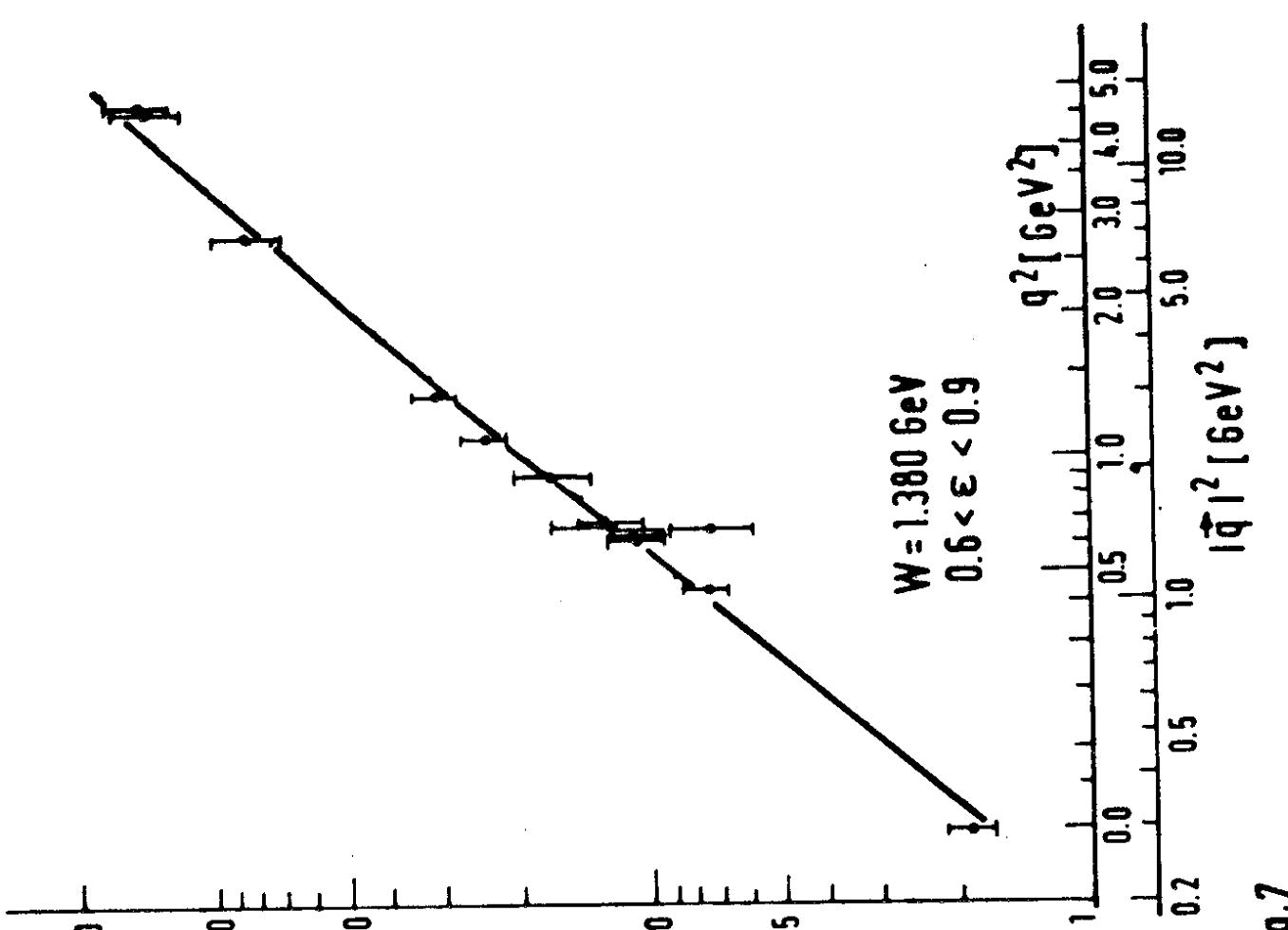
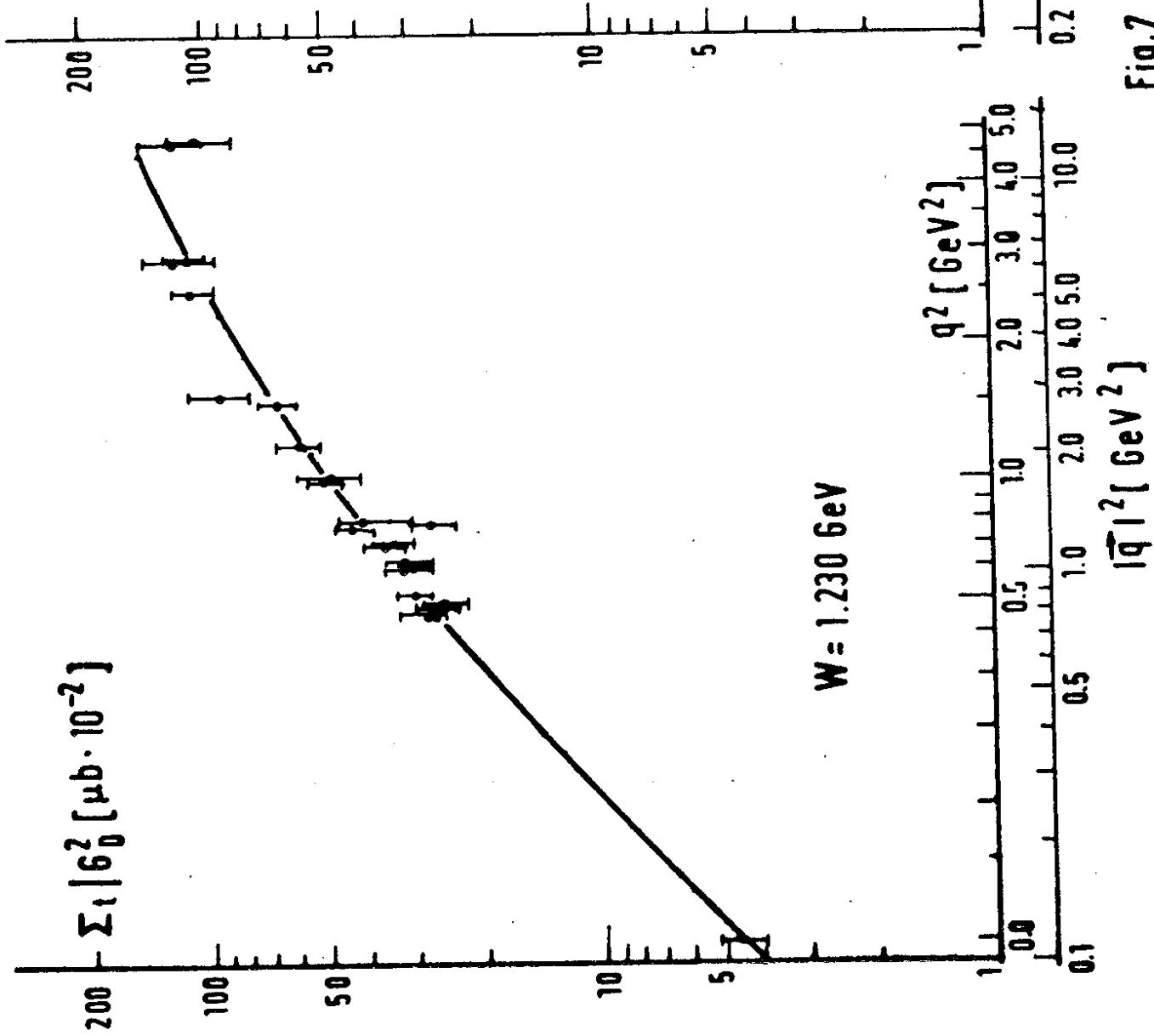
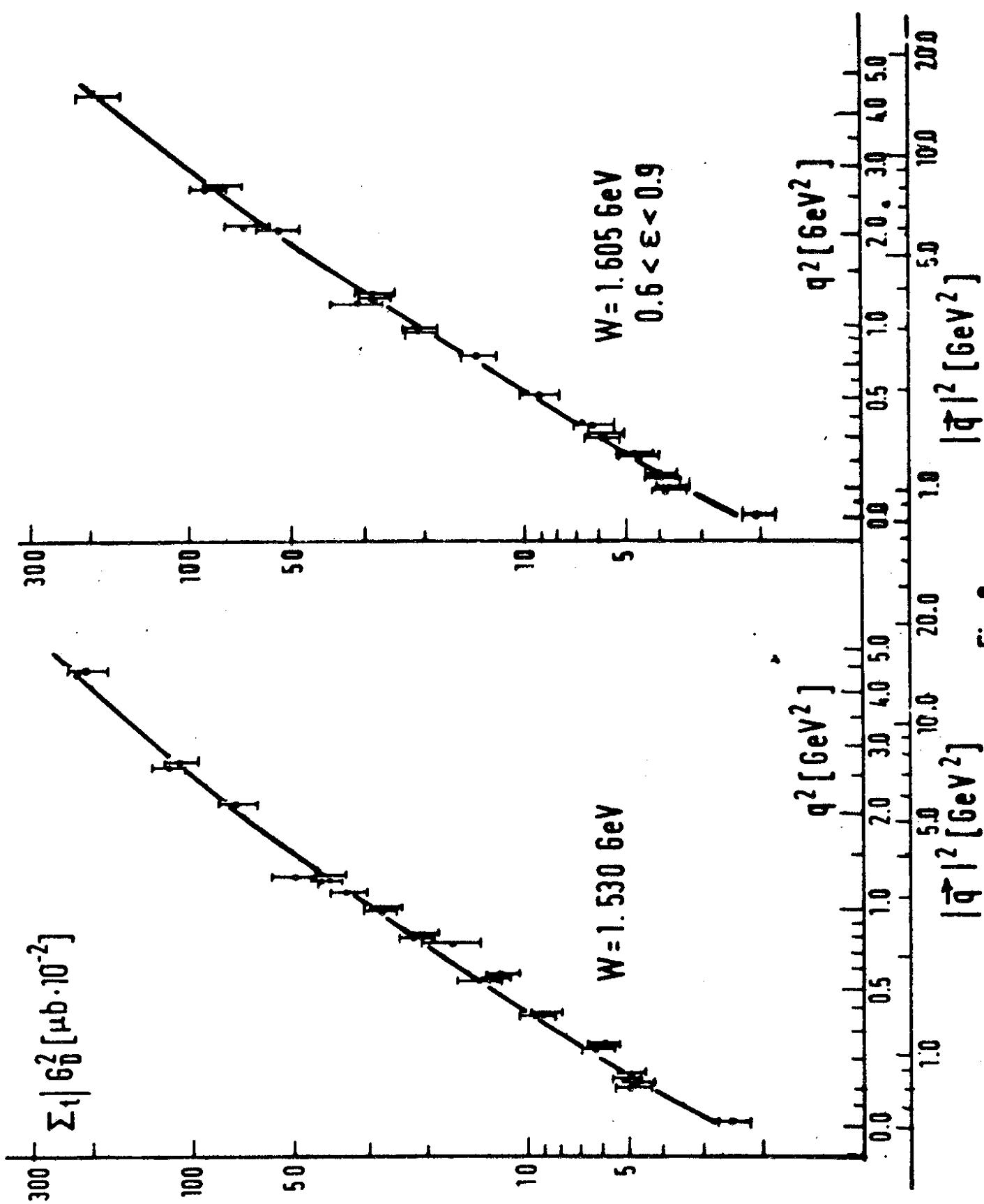


Fig. 7

Fig. 8



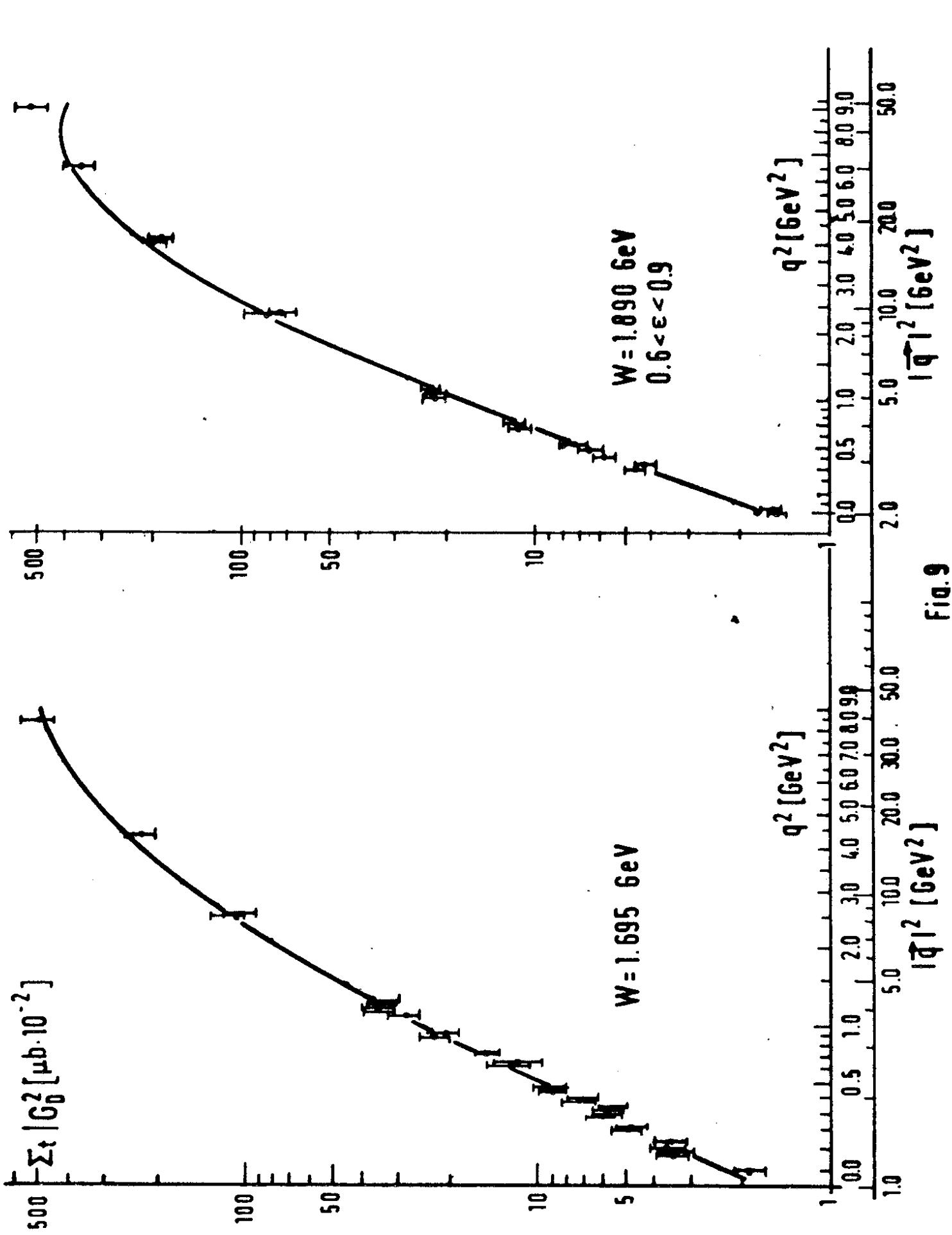
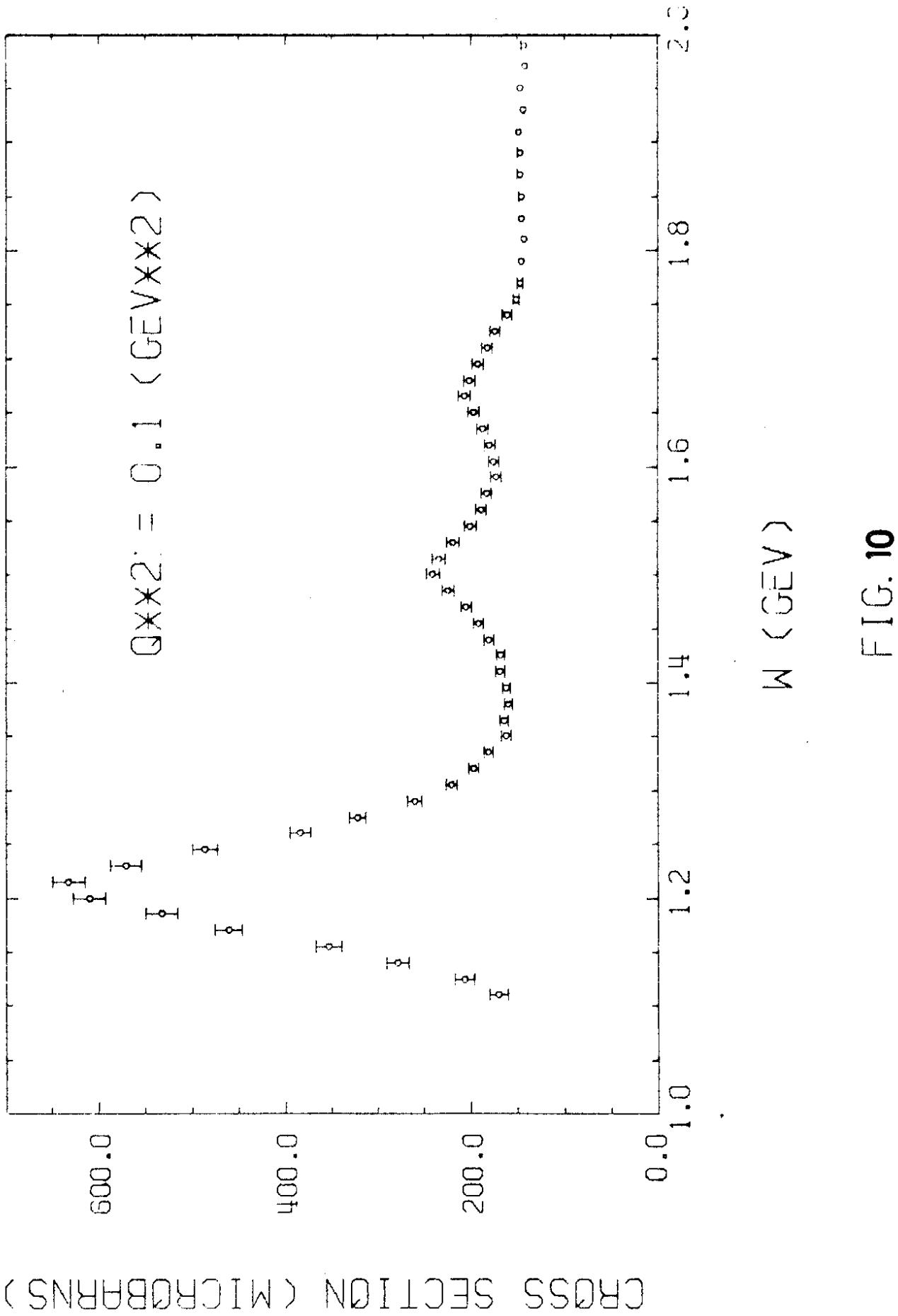


Fig. 9



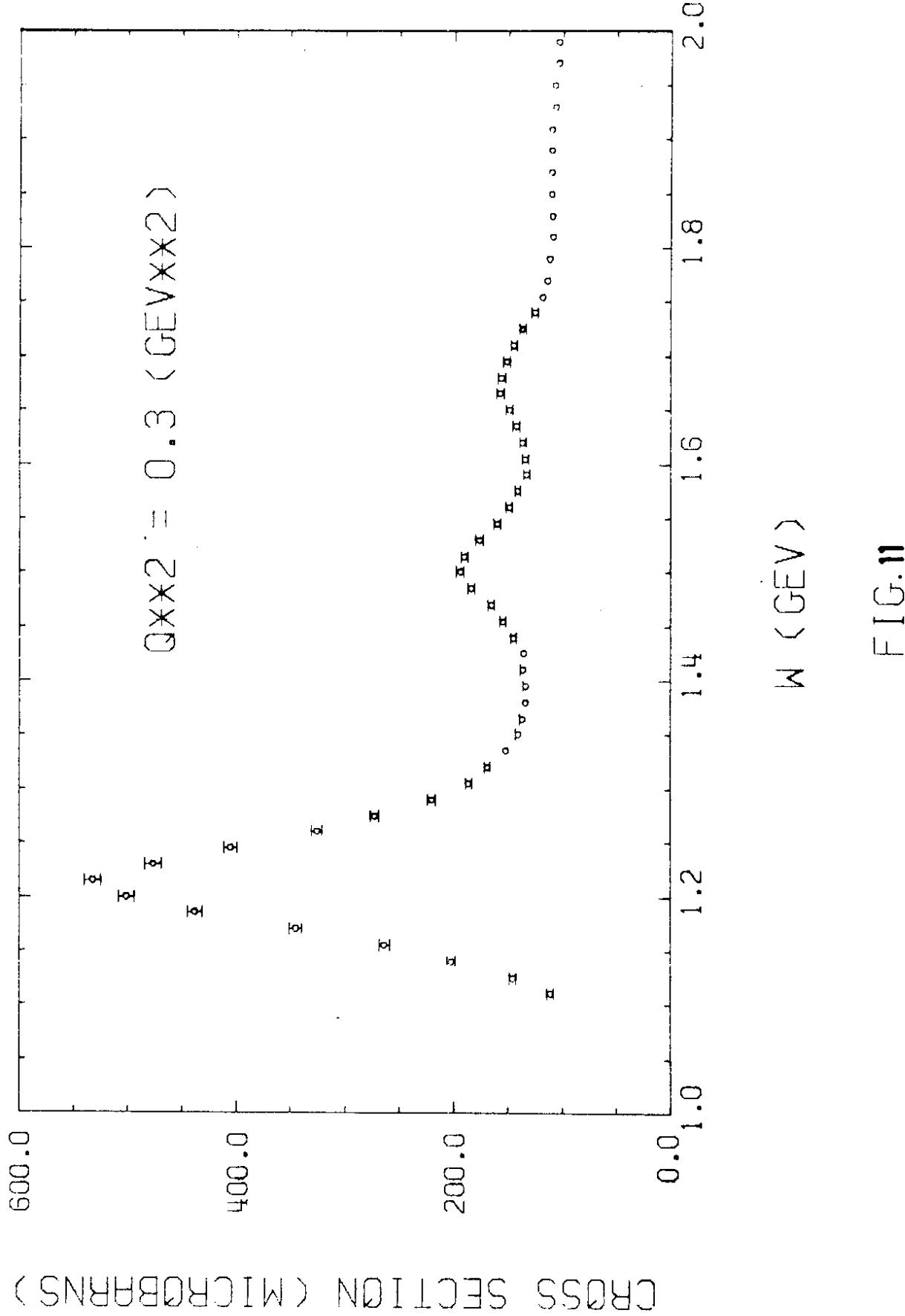


FIG. 11

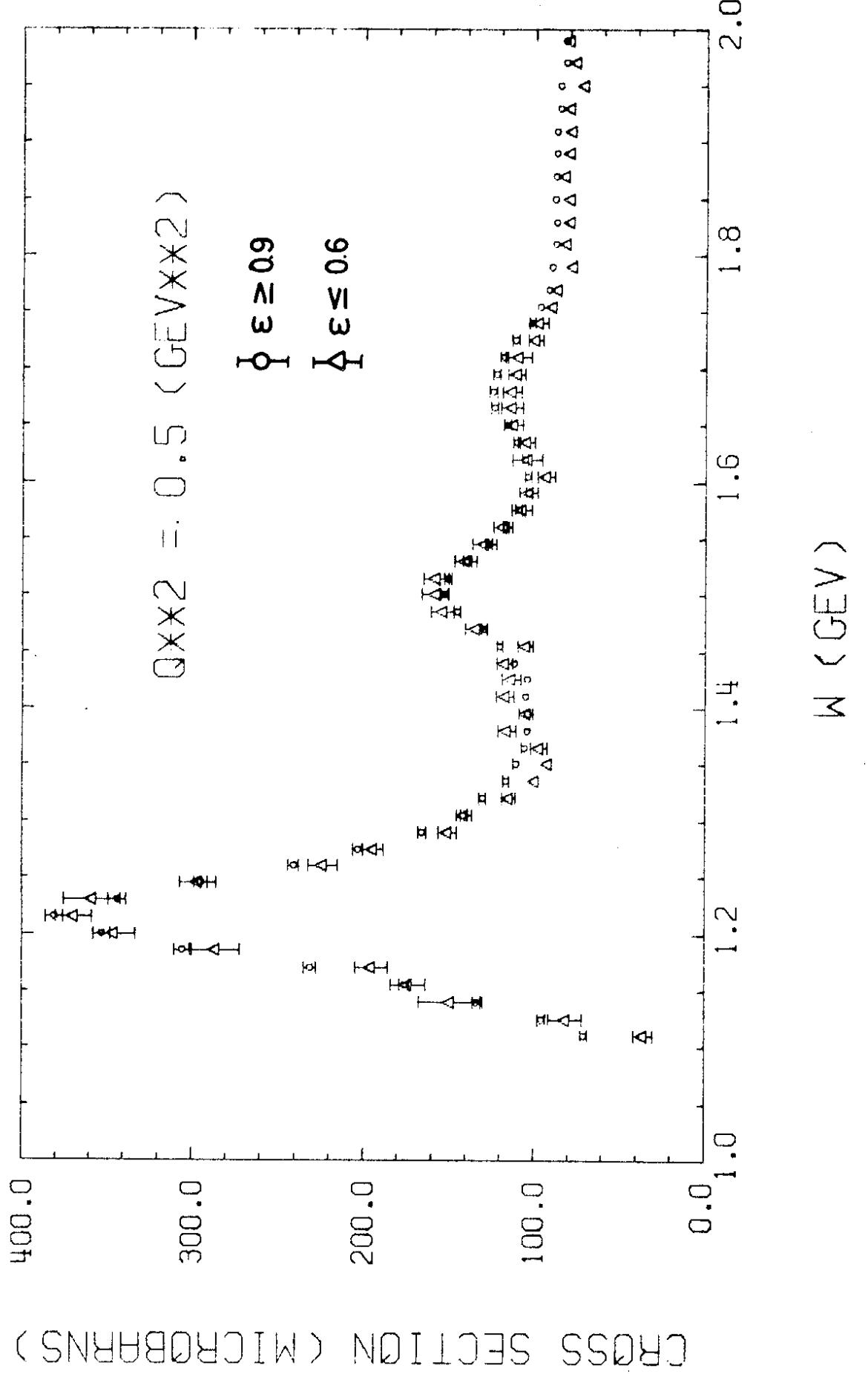


FIG. 12

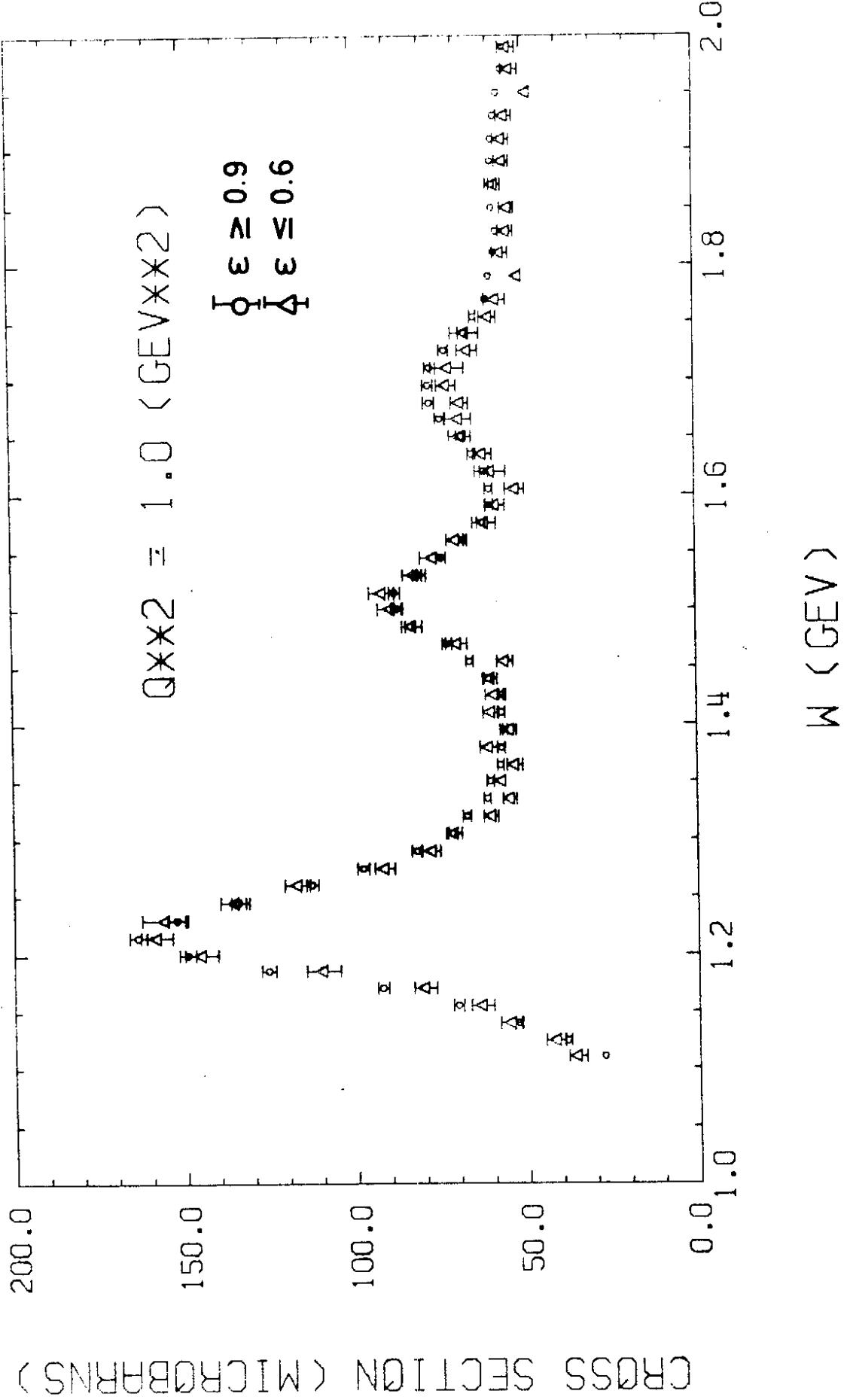


FIG 13

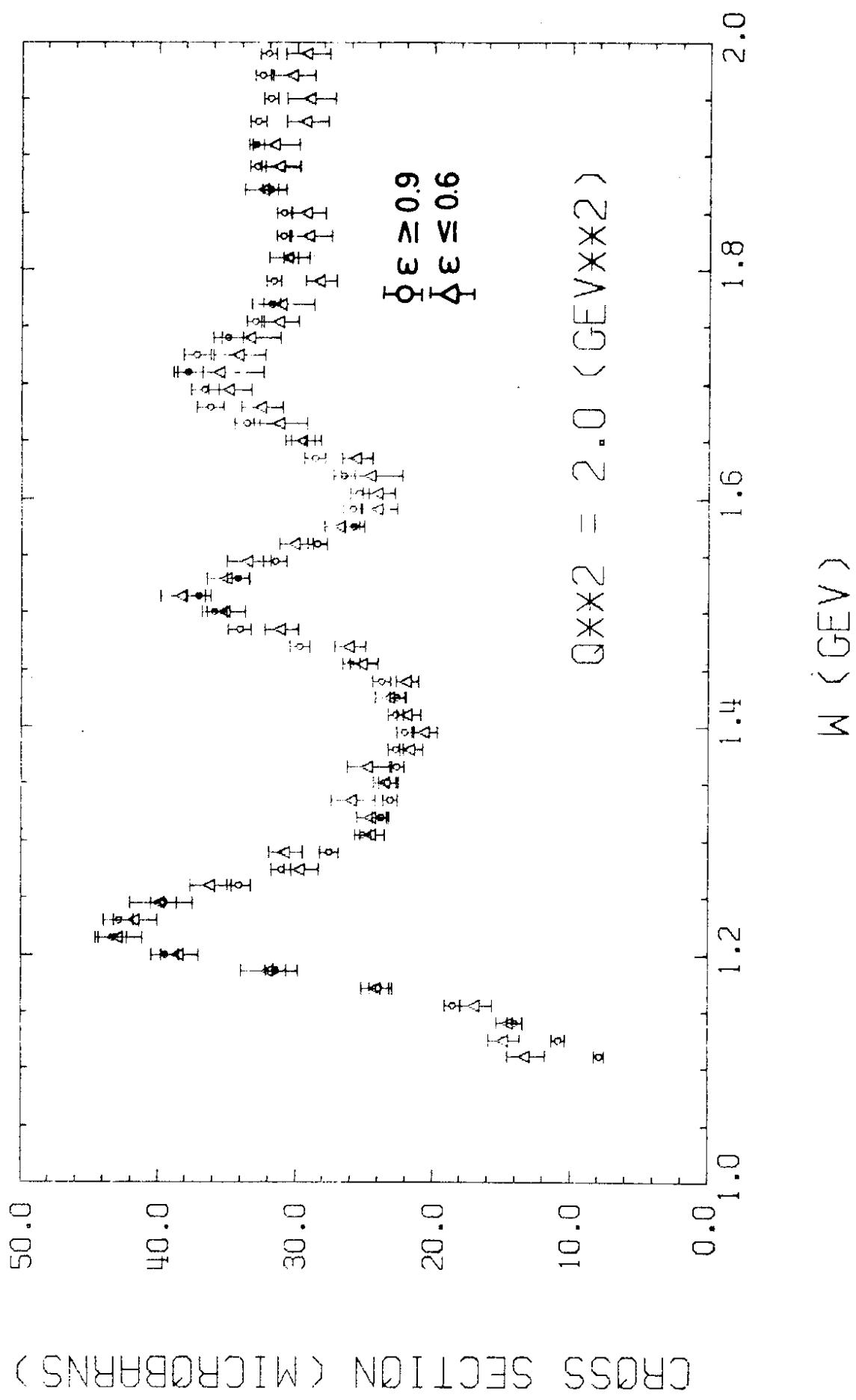
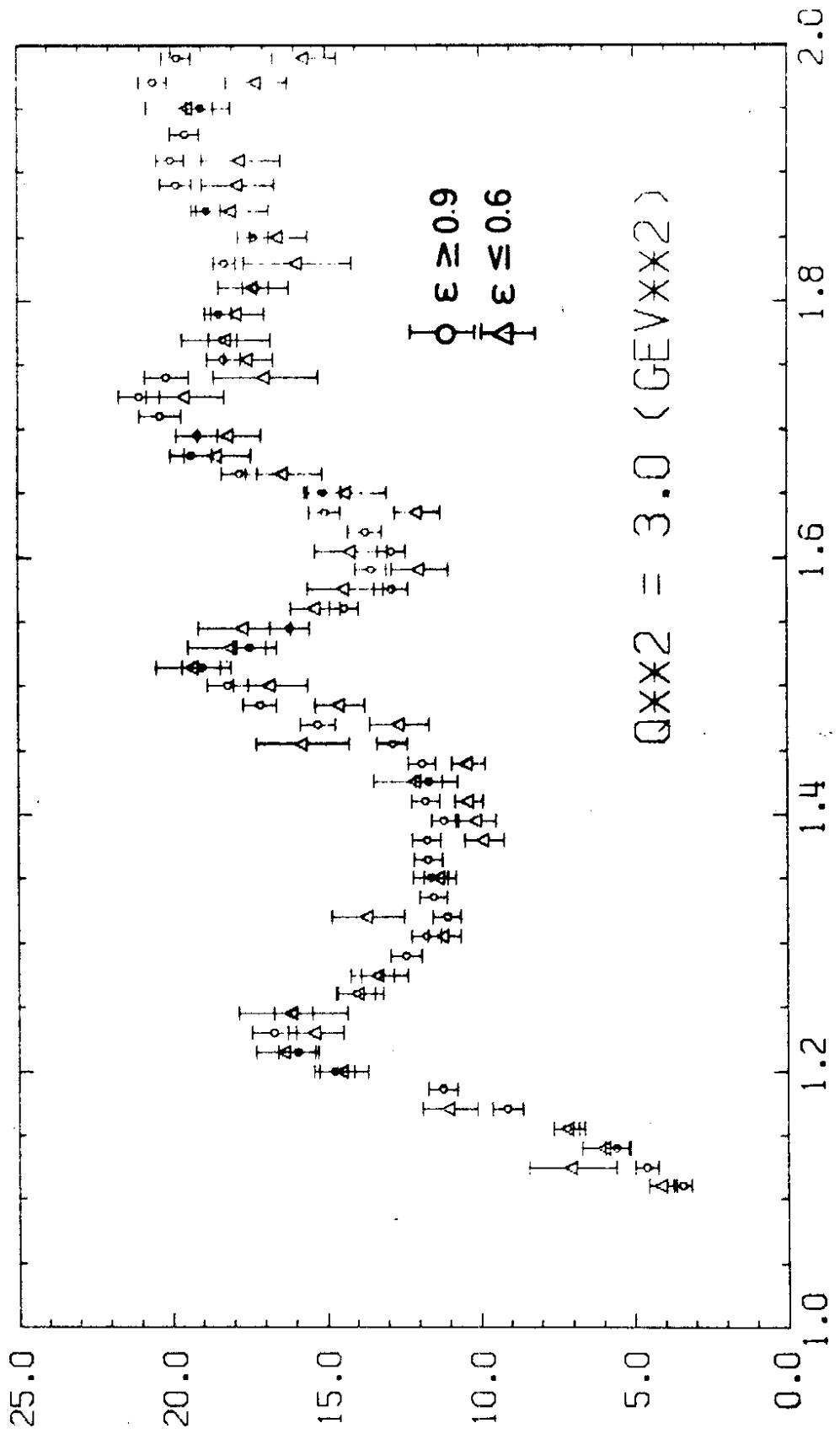


FIG 14



CROSS SECTION (MICROBARNs)

W (GeV)

FIG 15

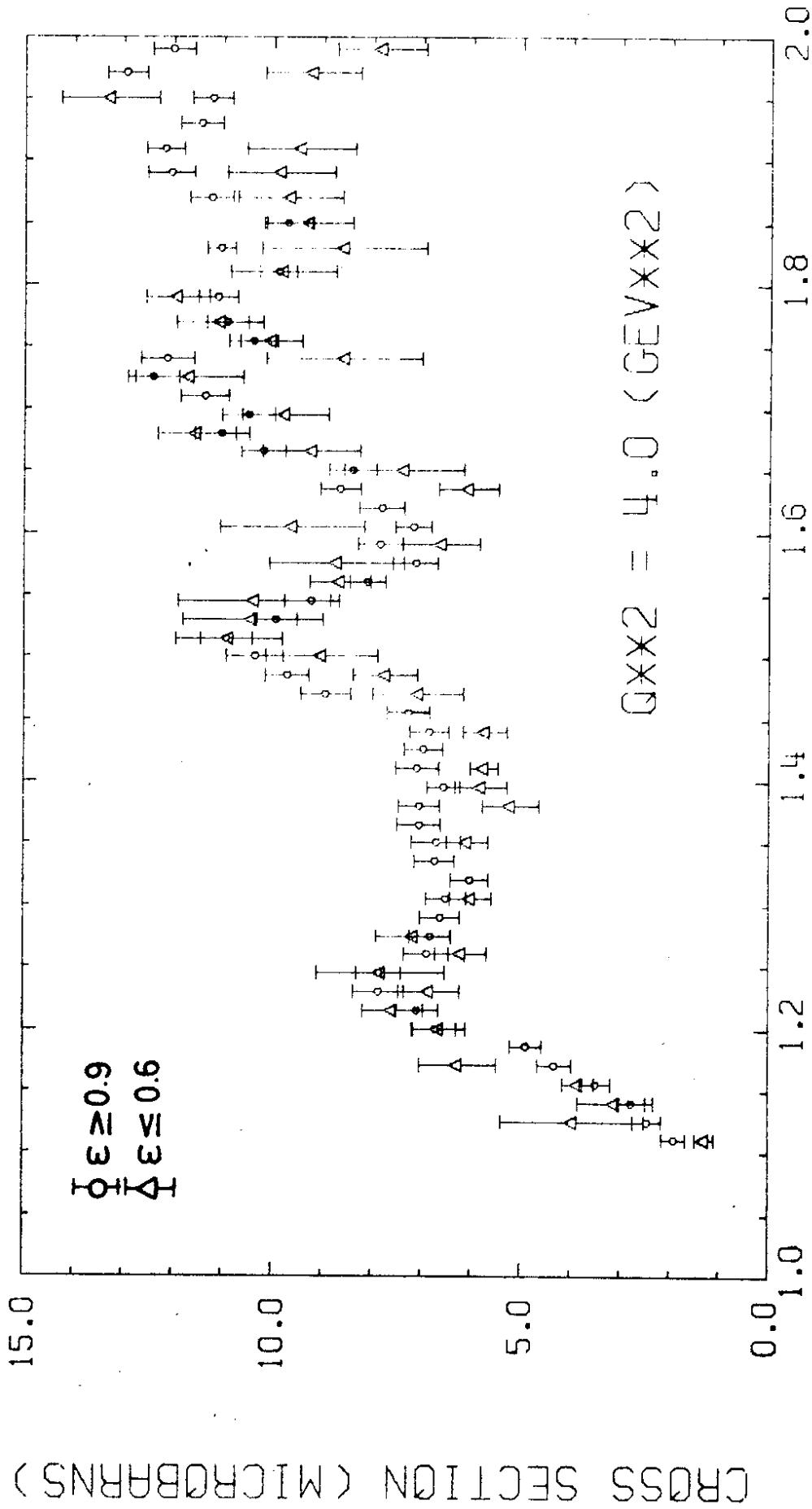


FIG 16

CROSS SECTION (MICROBARNs)

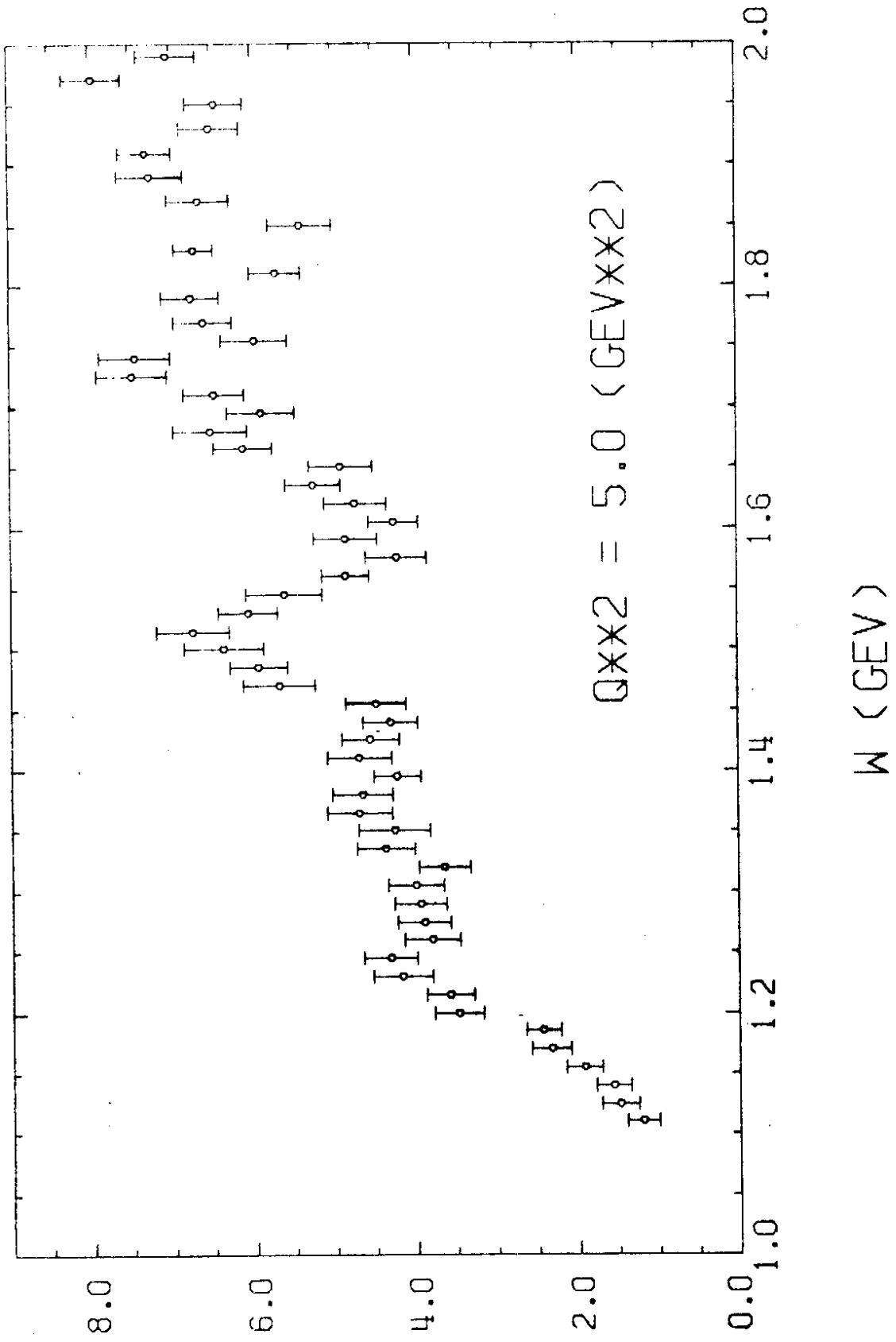


FIG 17

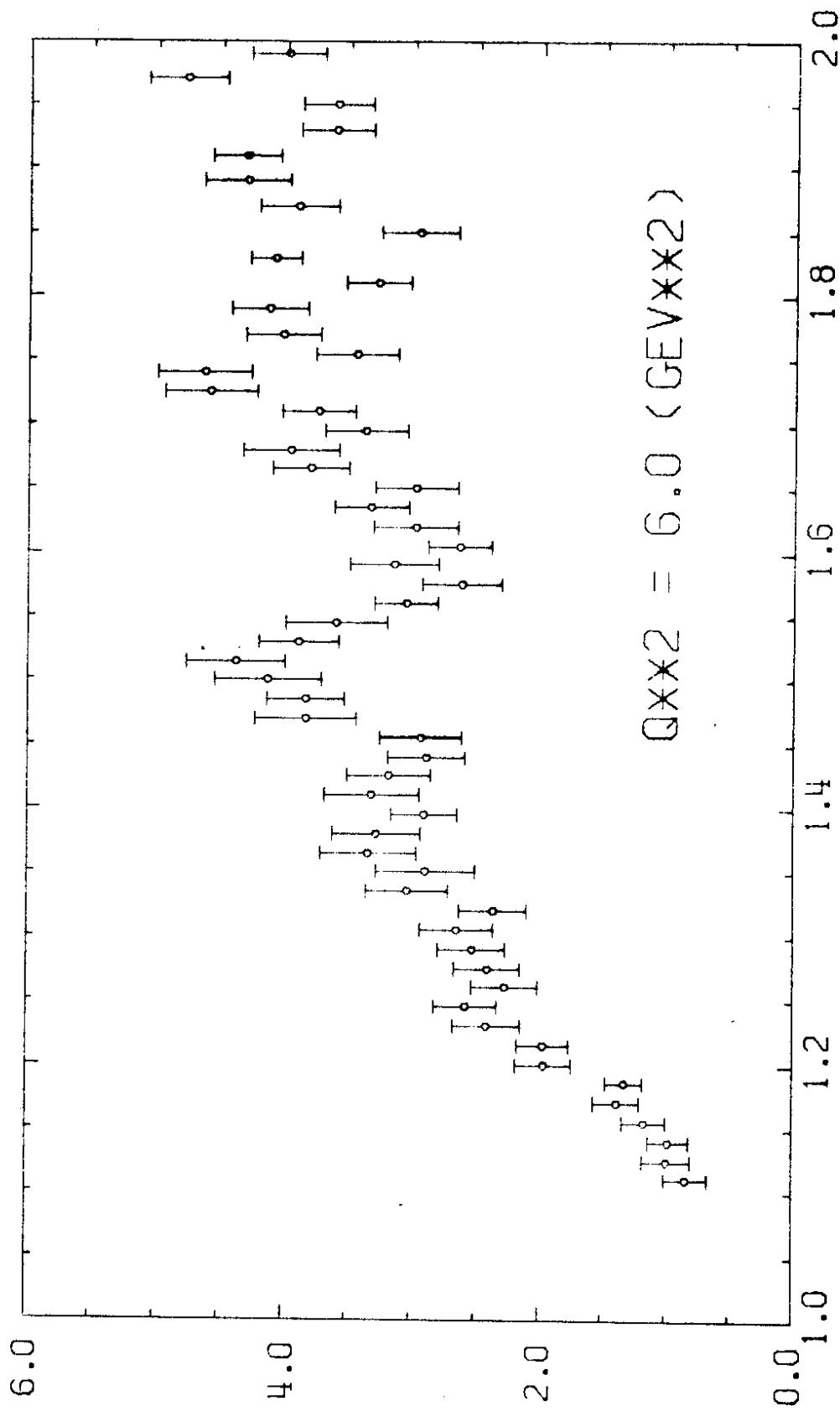


FIG 18