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## Addendum to: Implications of the measurements of $B_s - \overline{B_s}$ mixing on SUSY models

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This is an addendum to the previous publication, P. Ko and J.-h. Park, Phys. Rev. **D80**, 035019 (2009). The semileptonic charge asymmetry in  $B_s$  decays is discussed in the context of general MSSM with gluino-mediated flavor and CP violation in light of the recent measurements at the Tevatron.

In this addendum to Ref. [1], we discuss the semileptonic charge asymmetry in the  $B_s$  decays in general SUSY models with gluino-mediated flavor and CP violation, in light of the recent measurements of like-sign dimuon charge asymmetry by DØ Collaboration at the Tevatron. The model is described in Ref. [1], to which we refer for the details of the model and other phenomenological aspects related with  $B_s - \overline{B_s}$  mixing, the branching ratio of and CP asymmetry in  $B \to X_s \gamma$ ,  $B_d \to \phi K_S$  and CP asymmetry in  $B_s \to J/\psi\phi$ .

One can define the semileptonic charge asymmetry in the decay of  $B_q$  mesons as

$$a_{\rm sl}^q \equiv \frac{\Gamma(\overline{B_q^0}(t) \to \mu^+ X) - \Gamma(B_q^0(t) \to \mu^- X)}{\Gamma(\overline{B_q^0}(t) \to \mu^+ X) + \Gamma(B_q^0(t) \to \mu^- X)}, \quad (1)$$

for q = d, s. In terms of the matrix elements of the effective Hamiltonian describing the damped oscillation between  $B_q^0$  and  $\overline{B_q^0}$ , the asymmetry  $a_{\rm sl}^q$  is given by

$$a_{\rm sl}^q = {\rm Im} \frac{\Gamma_{12}^q}{M_{12}^q} = \frac{|\Gamma_{12}^q|}{|M_{12}^q|} \sin \phi_q, \qquad (2)$$

where  $\phi_q \equiv \arg(-M_{12}^q/\Gamma_{12}^q)$ . That is, this is another observable measuring CP violation in  $B_q-\overline{B_q}$  mixing. We take the approximation,  $\Gamma_{12}^q = \Gamma_{12}^{q,\text{SM}}$ , since the leading contribution comes from the absorptive part of the box diagrams for  $B_q-\overline{B_q}$  mixing and there is no new common final state into which both  $B_q$  and  $\overline{B_q}$  can decay in our scenario. The size of  $M_{12}^q$  is fixed by the  $\Delta M_q$  data up to hadronic uncertainties. Then,  $a_{\text{sl}}^q$  can be regarded as a sine function of  $\phi_q$ , multiplied by the factor  $|\Gamma_{12}^q|/|M_{12}^q|$ . This curve is traversed as one allows for arbitrary supersymmetric contributions to  $M_{12}^q$  obeying the  $\Delta M_q$  constraint. Combining the SM predictions [2],

$$\begin{aligned} |\Gamma_{12}^{s,\text{SM}}| / |M_{12}^{s,\text{SM}}| &= (49.7 \pm 9.4) \times 10^{-4}, \\ \phi_s^{\text{SM}} &= (4.2 \pm 1.4) \times 10^{-3}, \end{aligned} \tag{3}$$

one finds the vanishingly small asymmetry  $a_{\rm sl}^{s,{\rm SM}} \sim 2 \times 10^{-5}$ .

Recently, the DØ collaboration reported a measurement of like-sign dimuon charge asymmetry [3]. They interpreted the result as coming from the mixing of neutral B mesons and have found an evidence for an anomaly in the asymmetry,

$$A_{\rm sl}^b \equiv \frac{N_b^{++} - N_b^{--}}{N_b^{++} + N_b^{--}},\tag{4}$$

where  $N_b^{++}$  and  $N_b^{--}$  are the number of events where decays of two *b* hadrons yield two positive and two negative muons, respectively. Their result shows a discrepancy of  $3.2\sigma$  from the SM expectation. This asymmetry consists of  $a_{\rm sl}^d$  coming from  $B_d$  decays as well as  $a_{\rm sl}^s$  from  $B_s$ . One can extract the asymmetry relevant to the  $B_s$  meson using the measured value of  $a_{\rm sl}^d$  and the result by DØ is

$$a_{\rm sl}^s = -0.0146 \pm 0.0075. \tag{5}$$

This is  $1.9\sigma$  away from the SM prediction. We shall use this data in the following discussion.

This DØ result has drawn interest in new physics explanations [4–8]. (For earlier works, see e.g. Refs. [9–11].) Some of the works consider extra contributions to  $\Gamma_{12}^q$  since the dimuon charge asymmetry depends on it as well as on  $M_{12}^q$  [5, 6]. This approach also has a possibility of altering  $|\Delta\Gamma_s|$  even though its current experimental value is in agreement with the SM one,  $2 |\Gamma_{12}^{s,\text{SM}} \cos \phi_s^{\text{SM}}|$  [2, 12, 13]. As we said,  $\Gamma_{12}^q$  is fixed in the present work and we are left only with the option of modifying  $M_{12}^s$ . Therefore,  $|\Delta\Gamma_s|$  shall become smaller than its SM prediction as  $|\phi_s|$  grows up to  $\mathcal{O}(1)$ .

We perform the numerical analysis in the same way as in the main article [1]. The crucial ingredient for evaluating  $a_{sl}^s$  is the range of  $\phi_s$  to be used. Following the latest reports from DØ [3] and CDF [14], there have been a couple of attempts to make a global fit of  $B_s - \overline{B_s}$  mixing parameters including  $\phi_s$  [4, 6]. However, the official combination is not available yet. Partly because of this reason and partly for the sake of coherent presentation, we keep using the range used in Refs. [1, 15],

$$\phi_s \in [-1.10, -0.36] \cup [-2.77, -2.07].$$
 (6)

As a matter of fact, this range is not very different from the  $2\sigma$  interval found in Ref. [6]. As for  $\Gamma_{12}^{s,\text{SM}}/M_{12}^{s,\text{SM}}$ , we take its central value from Eqs. (3). Considering the error in this ratio could add 20% more of uncertainty to the thickness of the  $a_{\text{sl}}^s$  band in the following figures.

We show  $a_{sl}^s$  as a function of  $\phi_s$  for  $\tan\beta = 3$  in Figs. 1. The four plots are for the *LL*, the *RR*, the *LL* = *RR*, and



FIG. 1. Plots of  $a_{\rm sl}^s$  as a function of  $\phi_s$  for the four different cases with  $\tan\beta = 3$ . The hatched gray region leads to the lightest squark mass < 100 GeV. The hatched region is excluded by the  $B \to X_s \gamma$  constraint. The light gray region (cyan online) is allowed by  $\Delta M_s$ . The dark gray region (blue online) is allowed both by  $\Delta M_s$  and  $\phi_s$ . The black square is the SM point. The dashed and solid lines (both red online) mark the  $1\sigma$  and  $2\sigma$  ranges of  $a_{\rm sl}^s$ , respectively.

the LL = -RR cases, respectively. One can immediately notice the aforementioned sinusoidal dependence of  $a_{\rm sl}^s$  on  $\phi_s$ , coming from Eq. (2) and the  $\Delta M_s$  constraint. This feature is not only true of all the cases shown here but also of any new physics model that does not affect  $\Gamma_{12}^s$ . The nonzero thickness of the band arises from the uncertainty in  $\Delta M_s$ . The difference between  $a_{sl}^s$  and its central value is at least about  $1.0\sigma$ . This discrepancy becomes worse but only slightly after  $\phi_s$  is restricted inside its preferred ranges (colored in blue). If one incorporates the  $B \rightarrow$  $X_s \gamma$  constraint, substantial part of the blue regions is excluded, in particular in the upper two cases with one insertion. Even then, however, the lowest possible value of  $a_{\rm sl}^s \simeq -0.006$  within the blue region does not change. In the lower two cases with two insertions,  $B \to X_s \gamma$ does not play an important role since the supersymmetric effect on  $B_s - \overline{B_s}$  mixing is enhanced.

Plots for  $\tan\beta = 10$  are displayed in Figs. 2. The model-independent characteristics dictated by Eq. (2) remain exactly the same as in the previous set of figures. The only difference is the stronger  $B \to X_s \gamma$  constraint due to higher  $\tan\beta$ . Here, it excludes more part of the blue regions. Again, this is particularly true of the upper two cases in which  $a_{\rm sl}^{\rm sl}$  is restricted closer to its SM value. In Fig. 2(a),  $\Delta M_s$ ,  $\phi_s$ , and  $B \to X_s \gamma$ , together allow  $a_{\rm sl}^{\rm s}$  to be as low as -0.003. In Fig. 2(b), there is no solution satisfying all the three constraints. One could get  $a_{\rm sl}^{\rm sl} \simeq -0.0006$  if  $\phi_s$  were not limited. In the lower two cases, the lowest  $a_{\rm sl}^{\rm sl}$ , compatible with  $\Delta M_s$  and  $\phi_s$ , is almost the same as in Figs. 1.

We summarize. We have examined how  $a_{\rm sl}^s$  is influenced by the *LL* and/or *RR* mass insertions. For  $\tan\beta = 3$ , one can reduce the discrepancy between  $a_{\rm sl}^s$ and its SM expectation from  $1.9\sigma$  down to  $1.0\sigma$  in each



FIG. 2. Plots with  $\tan\beta = 10$ . The meaning of each region is the same as in Figs. 1.

of the *LL*, *RR*, *LL* = *RR*, and *LL* = -RR cases, obeying the  $\Delta M_s$ ,  $B \to X_s \gamma$ , and  $\phi_s$  constraints. This amounts to reduction of the  $A^b_{\rm sl}$  tension from  $3.2\sigma$  down to  $2.2\sigma$ if one assumes no new physics in the  $b \to d$  transition. For  $\tan\beta = 10$ , it becomes difficult for the *LL* and *RR* cases whereas the *LL* = *RR* and *LL* = -RR cases are less limited by  $B \to X_s \gamma$ .

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## NOTE ADDED

While we were waiting for the approval for submission, a paper by J. K. Parry appeared on the e-print archive that employs a related model [8]. However, the flavor structure of the squark mass matrix therein is different from any of those here. As far as squarks are concerned, he considers only one case where  $(\delta_{23}^d)_{RR}$  is a variable parameter and  $(\delta_{23}^d)_{LL}$  is fixed to a value that comes from renormalization group running. This way of parameter scan is not covered in this work. He does not display the  $B \to X_s \gamma$  constraint on his plots, but it may not be very restrictive in his case depending on  $\mu$  and  $\tan\beta$ . (See e.g. Fig. 4 in Ref. [16].)

- [1] P. Ko and J.-h. Park, Phys. Rev. D 80, 035019 (2009)
   [arXiv:0809.0705 [hep-ph]].
- [2] A. Lenz and U. Nierste, JHEP 0706, 072 (2007) [arXiv:hep-ph/0612167].

- [3] V. M. Abazov *et al.* [D0 Collaboration], Phys. Rev. D 82, 032001 (2010) [arXiv:1005.2757 [hep-ex]].
- [4] Z. Ligeti, M. Papucci, G. Perez and J. Zupan, Phys. Rev. Lett. 105, 131601 (2010) [arXiv:1006.0432 [hep-ph]].
- [5] A. Dighe, A. Kundu and S. Nandi, Phys. Rev. D 82, 031502 (2010) [arXiv:1005.4051 [hep-ph]].
- [6] C. W. Bauer and N. D. Dunn, arXiv:1006.1629 [hep-ph].
- [7] B. A. Dobrescu, P. J. Fox and A. Martin, Phys. Rev. Lett. **105**, 041801 (2010) [arXiv:1005.4238 [hep-ph]];
  A. J. Buras, M. V. Carlucci, S. Gori and G. Isidori, JHEP **1010**, 009 (2010) [arXiv:1005.5310 [hep-ph]]; C. H. Chen and G. Faisel, arXiv:1005.4582 [hep-ph]; D. Choudhury and D. K. Ghosh, arXiv:1006.2171 [hep-ph]; N. G. Deshpande, X. G. He and G. Valencia, Phys. Rev. D **82**, 056013 (2010) [arXiv:1006.1682 [hep-ph]]; C. H. Chen, C. Q. Geng and W. Wang, JHEP **1011**, 089 (2010) [arXiv:1006.5216 [hep-ph]].
- [8] J. K. Parry, arXiv:1006.5331 [hep-ph].
- [9] G. L. Kane, P. Ko, H. b. Wang, C. Kolda, J.-h. Park and L. T. Wang, Phys. Rev. Lett. 90, 141803 (2003)

[arXiv:hep-ph/0304239]; Phys. Rev. D **70**, 035015 (2004) [arXiv:hep-ph/0212092].

- [10] A. Lenz, Phys. Rev. D 76, 065006 (2007) [arXiv:0707.1535 [hep-ph]].
- [11] L. Randall and S. f. Su, Nucl. Phys. B 540, 37 (1999)
   [arXiv:hep-ph/9807377]; K. Kawashima, J. Kubo and A. Lenz, Phys. Lett. B 681, 60 (2009) [arXiv:0907.2302
   [hep-ph]].
- [12] E. Barberio *et al.* [Heavy Flavor Averaging Group], arXiv:0808.1297 [hep-ex].
- [13] S. Esen *et al.*, Phys. Rev. Lett. **105**, 201802 (2010) [arXiv:1005.5177 [hep-ex]].
- [14] L. Oakes (CDF Collaboration), talk at FPCP 2010, May 25-29, Torino, Italy, http://agenda.infn.it/getFile. py/access?contribId=12&resId=0&materialId= slides&confId=2635
- [15] M. Bona *et al.* [UTfit Collaboration], PMC Phys. A 3, 6 (2009) [arXiv:0803.0659 [hep-ph]].
- [16] P. Ko, J.-h. Park and M. Yamaguchi, JHEP 0811, 051 (2008) [arXiv:0809.2784 [hep-ph]].